

1. What is SUPERSYMMETRY?

1.1 Introduction

Reminder: How do we describe symmetries in particle physics?

$U(1)$ - Gauge Symmetry: $\psi \rightarrow \psi'(x) = e^{iQ\alpha(x)} \psi(x)$

Q : charge operator
generator of symmetry

For electron: $Q\psi_e(x) = (-e)\psi_e(x)$

$SU(2)$ - Gauge Symmetry:

$$\psi_{(x)} = \begin{pmatrix} \psi_1(x) \\ \psi_2(x) \end{pmatrix} \rightarrow \begin{pmatrix} \psi'_1(x) \\ \psi'_2(x) \end{pmatrix} = e^{-i\frac{\theta_a(x)}{2}\tau^a} \begin{pmatrix} \psi_1(x) \\ \psi_2(x) \end{pmatrix}$$

τ^a : Pauli matrices, $\tau^1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$, $\tau^2 = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}$, $\tau^3 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$

$$\tau^\pm = \tau^1 \pm i\tau^2, \quad \tau^+ = \begin{pmatrix} 0 & 1 \\ 0 & 0 \end{pmatrix}, \quad \tau^- = \begin{pmatrix} 0 & 0 \\ 1 & 0 \end{pmatrix}$$

$$\tau^+ \begin{pmatrix} 1 \\ 0 \end{pmatrix} = \begin{pmatrix} 1 \\ 0 \end{pmatrix}, \quad \tau^- \begin{pmatrix} 1 \\ 0 \end{pmatrix} = \begin{pmatrix} 0 \\ 1 \end{pmatrix}$$

$$\mathcal{L}_4 = \bar{\psi}(x) i\gamma^\mu (\partial_\mu - ig \frac{\vec{\tau} \cdot \vec{W}_\mu}{2}) \psi(x) - m \bar{\psi}(x) \psi(x)$$

τ^\pm flips the isospin (and "has" isospin)
 $\rightsquigarrow W^\pm$

In Supersymmetry: generator: Q, \bar{Q}

$$Q | \text{BOSON} \rangle_{S=0} = | \text{FERMION} \rangle_{S=\frac{1}{2}}$$

$$\bar{Q} | \text{FERMION} \rangle_{S=\frac{1}{2}} = | \text{BOSON} \rangle_{S=0}$$

change spin by $\pm \frac{1}{2}$, Q : raises spin
 \bar{Q} : lowers spin ($\bar{Q} = Q^\dagger$)

Q : is a (spin- $\frac{1}{2}$) spinor (2-comp. Weyl spinor)

Want a Lagrangian: $\mathcal{L}_{\text{SUSY}}$

$$\text{for which: } \int d^4x \delta \mathcal{L}_{\text{SUSY}} = \delta S_{\text{SUSY}} = 0$$

invariant under SUSY transformations.

\rightarrow not possible within SM — need to double spectrum

EXAMPLE: ELECTRON

(6)

$$(\psi_e)_L (s=\frac{1}{2}) \longleftrightarrow \varphi_{\tilde{e}_L} (s=0)$$

Scalar Electron
"selectron"

$$(\psi_e)_R (s=\frac{1}{2}) \longleftrightarrow \varphi_{\tilde{e}_R} (s=0)$$

\tilde{e}_L : "left-handed" selectron

\tilde{e}_R : "right-handed" selectron

These are scalars — no spin — and thus have no helicity. "L" & "R" are reminders of the SU(2)

transformation properties

$$SU(2) \updownarrow \begin{pmatrix} \tilde{\nu}_L \\ \tilde{e}_L \end{pmatrix} - \text{doublet}; \quad \tilde{e}_R - \text{singlet} \quad (\tilde{\nu}_R)$$

The superpartners (e_L, \tilde{e}_L) have the same gauge properties:

$$[Q, \text{"gauge generators"}] = 0$$

$$[Q, T^a] = 0$$

EXTERNAL SYMMETRIES

"Space-Time"

Ex: Lorentz group

SUPERSYMMETRY

INTERNAL SYMMETRIES ⁽⁷⁾

"Gauge Symmetries"

SU(2)

Coleman-Mandula / Haag-Kopuzanski-Sohnius Theorem:

"Supersymmetry is the only possible external symm. of the S-matrix, beyond Lorentz symmetry for which the S-matrix is not trivial."

$$S_{fi} = \langle \text{final} | \text{initial} \rangle = \mathbb{1} - i T_{fi}$$

End of the line for external symmetries

1.2. Supersymmetry Algebra

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Reminder: $U(1)$: $[Q, Q] = 0$ only 1 generat.
algebra is trivial

$$SU(2): \left[\frac{T_i}{2}, \frac{T_j}{2} \right] = i \epsilon_{ijk} \frac{T_k}{2}$$

Supersymmetry is an external symmetry

→ transforms non-trivially under
Lorentz transf^{ns}

Lorentz-Generators: P^μ — translations (Poincaré)

$$J^{\mu\nu} \begin{cases} \text{rotations: } J^i = \frac{1}{2} \epsilon^{ijk} J^{jk} \\ \text{boosts: } K^i = J^{0i} \end{cases}$$

$$[P^\mu, P^\nu] = 0 \quad \text{translations commute}$$

$$[J^{\mu\nu}, P^\lambda] = i(g^{\nu\lambda} P^\mu - g^{\mu\lambda} P^\nu)$$

$$[J^{\mu\nu}, J^{\lambda\rho}] = i(g^{\nu\lambda} J^{\mu\rho} - g^{\mu\lambda} J^{\nu\rho} + g^{\mu\rho} J^{\nu\lambda} - g^{\nu\rho} J^{\mu\lambda})$$

Examples : $[J^{\mu\nu}, P^\lambda] = i(g^{\nu\lambda} P^\mu - g^{\mu\lambda} P^\nu)$ ①

$(\mu\nu) = (ij), \lambda = 0 \quad [J^{ij}, P^0] = i(g^{j0} P^i - g^{i0} P^j) = 0$

rot^{ns} commute w/ time translation

$(\mu\nu) = (12), \lambda = 3 \quad [J^{12}, P^3] = i(g^{23} P^1 - g^{13} P^2) = 0$

rotⁿ around 3-axis commutes w/ translⁿ along 3-axis.

$(\mu\nu) = (03), \lambda = 0 \quad [J^{03}, P^0] = i(g^{30} P^0 - g^{00} P^3) = -i P^3$

boost along 3-axis does not commute with time translation

Now what happens when we include spinor Q ?

First make 4-comp. Majorana spinor

$$Q_M = \begin{pmatrix} Q_\alpha \\ \bar{Q}^{\dot{\alpha}} \end{pmatrix}$$

a) $[Q_M, P^M] = 0$

spin is unaffected by translations.

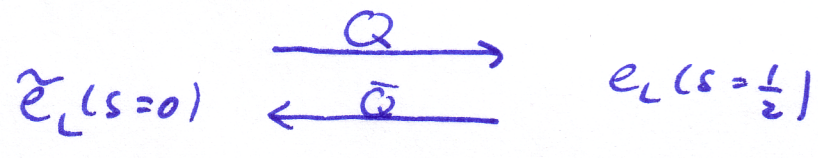
b) $[Q_M, J^{MN}] = \frac{1}{2} \sigma^{MN} Q_M$, $\sigma^{MN} = \frac{i}{4} [\gamma^M, \gamma^N]$

c) $\{Q_M, \bar{Q}_N\} = 2 \gamma^M P_M$, d) $\{Q_M, Q_N\} = 0$

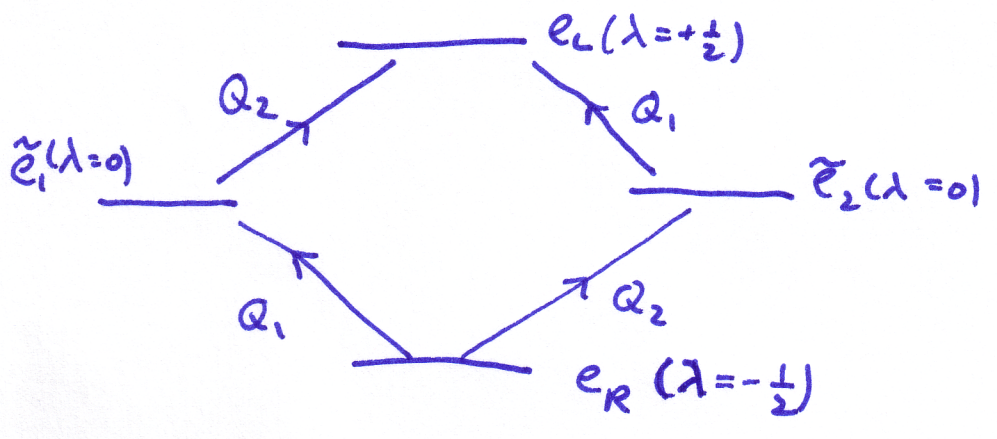
b) spin is affected by rot^{ns} — Thomas precession

c) $\{Q_M, \bar{Q}_N\} \rightarrow$ ^{space-time} translation

Q: raises by spin $\frac{1}{2}$, \bar{Q} : lowers by spin $\frac{1}{2}$



In principle could have more than one
susy Generator: Q_1, Q_2 (\bar{Q}_1, \bar{Q}_2)



$$\{Q_1, Q_2\} = 0$$

Since $[Q_{1,2}, T^a] = 0 \Rightarrow e_L$ & e_R would
have same gauge quantum numbers. \Downarrow

This is called $N=2$ supersymmetry and it is
phenomenologically excluded.

$$[Q, P^\mu] = 0$$

$$\Rightarrow [Q, P^2] = P_\mu [Q, P^\mu] + [Q, P_\mu] P^\mu = 0$$

$$\Rightarrow [Q, (\text{Mass})^2] = 0$$

$$\Rightarrow \text{Mass}(e^-) = \text{Mass}(\tilde{e}^-)$$

↑
selectron

⇒ Supersymmetry predicts a
(spin=0, Q=-1, m=511 keV) - particle

- EXPERIMENTALLY EXCLUDED.

⇒ supersymmetry must be broken

$$\text{Mass}(\tilde{e}^-) \gg \text{Mass}(e^-)$$

↑
how much bigger?

1.3. HIERARCHY PROBLEM

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Consider first electron as a charged sphere

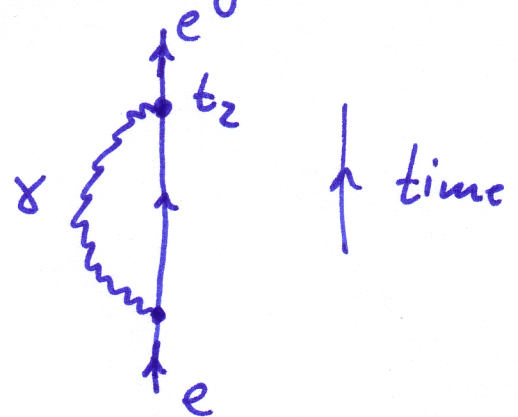


Using Maxwell's Eq^{ns} compute Coulomb self-energy

$$E_{\text{self}} = \frac{3}{5} \frac{e^2}{r_e}$$

- This is linearly divergent for $r_e \rightarrow 0$, $E_{\text{self}} \rightarrow \infty$
- This is the problem of a "point particle" in classical E & M
- In Quantum Mechanics compute the diagram

- single electron theory
- use time ordered perturbation theory



$$\Rightarrow m_e c^2 = m_e^0 c^2 + E_{\text{self}}$$

↑ as above

$$r_e \rightarrow 0 \Rightarrow \text{lower } m_e^0 \text{ to keep } m_e c^2 = 511 \text{ keV}$$

But at $r_e = 4 \text{ fm} \Rightarrow E_{\text{self}} \geq 511 \text{ keV}$

$$\Rightarrow m_e^0 < 0 ?$$

At LEP (CERN) [Large Electron Positron Collider
 $E_{\text{cm}} = 90 - 200 \text{ GeV}$]

point-like nature of electron is confirmed

for $\tau_e \leq 10^{-3} \text{ fm} \ll 4 \text{ fm}$

What to do?

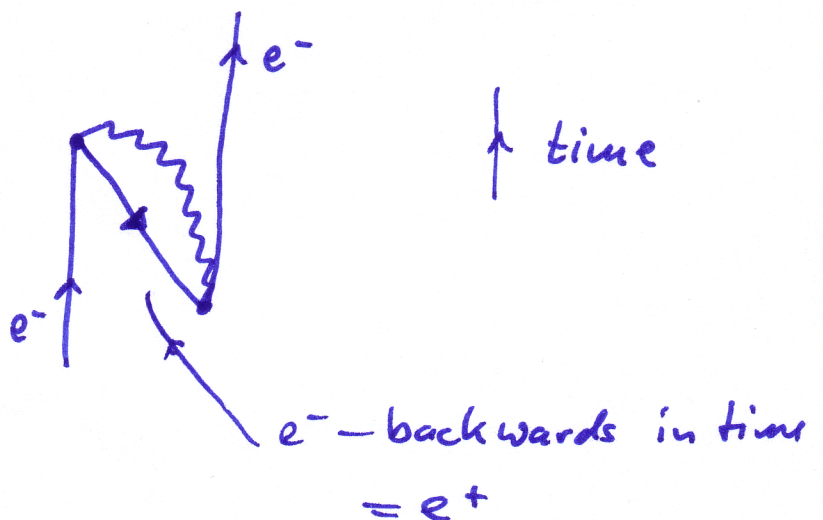
Introduce a new particle: e^+ - Positron

$$m(e^+) = m(e^-)$$

$$Q(e^+) = -Q(e^-)$$

Same quantum numbers - except charge, which is opposite

New diagram:



$$\Rightarrow E_{self} = \frac{3}{4\pi} \frac{e^2}{r_{compt}} \ln \left(\frac{m_e c \cdot r_e}{\hbar} \right)$$

only log divergence

Historically this was first done by Viktor Weisskopf, 1934 — he actually got it wrong first, but was helped out by Furry.



If I set $r_e = l_{Planck}$ (the smallest length scale in physics)

$$\Rightarrow E_{self} = \mathcal{O}(0.1) m_e c^2 \quad \text{— a wild correction}$$

In QED just one diagram:



$$\Delta m_e = \frac{\alpha}{\pi} m_e \ln \left(\frac{\Lambda^2 + m_e^2}{m_e^2} \right), \quad k < \Lambda$$

$$\Lambda = M_{pl} \Rightarrow \Delta m_e = \frac{1}{4} m_e$$

What have we done?

We have doubled particle content
→ reduced divergence

We have also extended the symmetry.

We now have an extra CPT-invariance.

In the SM we have particles

$S=1$ — Gauge bosons

$S=\frac{1}{2}$ — Quarks & Leptons

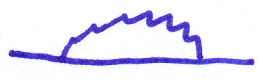
$S=0$ — Higgs boson

$S=1$



correction to photon mass
 $= 0$, protected by gauge symm.

$S=\frac{1}{2}$



$\sim m_e \cdot \ln(\frac{\Lambda}{m_e})$

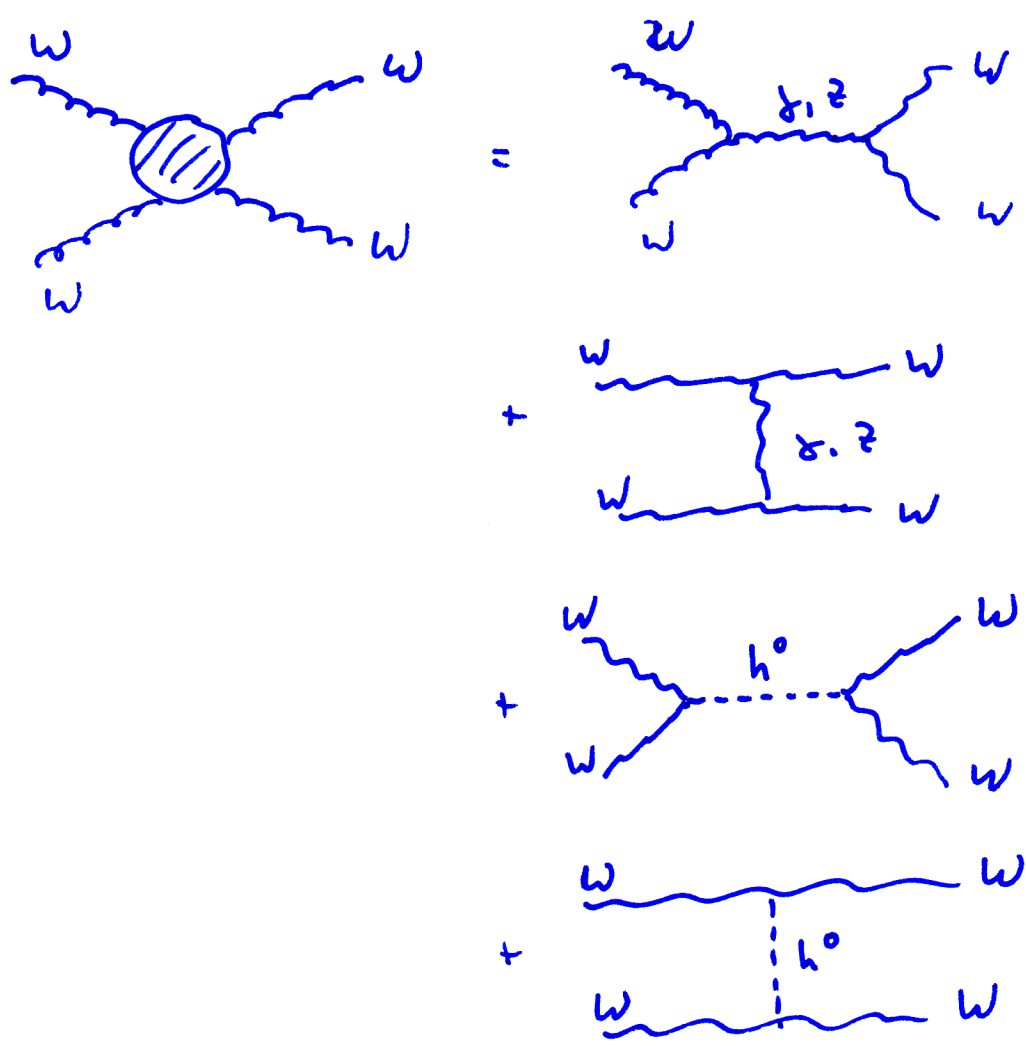
mild log. corr.
vanishes when $m_e \rightarrow 0$
chiral symm.

φ is not protected by a symmetry.

This is a disaster in SM, need

$$M_{\text{Higgs}} \leq \mathcal{O}(1 \text{ TeV})$$

by unitarity



violates unitarity if $M_h \geq 1 \text{ TeV}$

Introducing Λ only makes sense if there is

a new scale: $\Lambda \sim M_{\text{GUT}} (10^{16} \text{ GeV})$

$M_{\text{see-saw}} (10^{10} \text{ GeV})$

$M_{\text{Planck}} (10^{19} \text{ GeV}) \leftarrow$

If there is no new scale, then can absorb the Λ^2 -divergence into m_h^2 - renormalise - just like the log divergence.

But we are confident, that new physics must arise at the latest at M_{Planck} .

And even if we can absorb the ∞ when including the new physics - we have the finite correction:

$$\Delta m_h^2 = \frac{N(f) \lambda_f^2}{4\pi^2} m_f^2 \left(\frac{\Lambda^2}{\Lambda^2 + m_f^2} \right)$$

If there are any new particles at scale Λ coupling to h we are doomed again.

One could get rid of correction via a counter term in \mathcal{L}

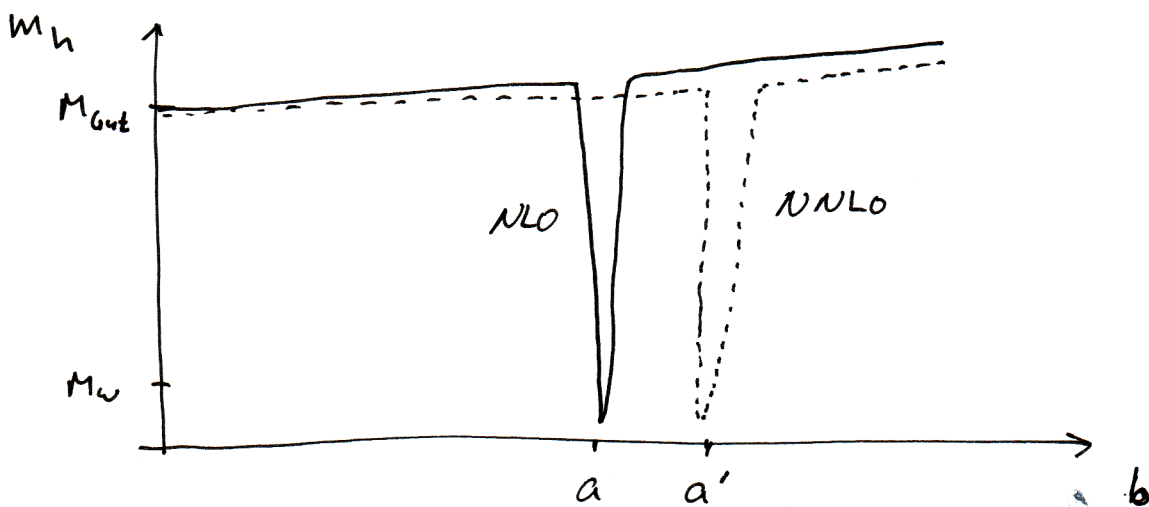
This counter term: $-b \cdot \Lambda^2$ would have to be fine-tuned to a fantastic precision

$$(m_h)^2 = (m_h^0)^2 + \underbrace{a \Lambda^2}_{\text{rad. corr.}} - b \Lambda^2$$

$$\Rightarrow (a-b) \cdot \Lambda^2 = \mathcal{O}((1 \text{ TeV})^2)$$

$$\Rightarrow a-b = \mathcal{O}(10^{-29}) \quad !$$

This is a huge fine-tuning



Different fine-tuning in each order of pert. theory

Prediction?

Standard Modell has a deep problem

→ need new physics

First, let us ask if there is a higher scale.

* Surprisingly, this question is experimentally accessible.

Plot →

For $130 \text{ GeV} \lesssim m_h \lesssim 200 \text{ GeV}$ avoid these problems.

Hamberger & Riesselmann
hep-ph/9708416

But if we find $m_h = 400 \text{ GeV}$, for example

⇒ $\Lambda \lesssim 100 \text{ TeV}$

Further hints of a higher scale

- Gravity — G_N : Newton's constant

$$G_N =$$

$$\Rightarrow r = \ell_{Pl} = \frac{1}{M_{Pl}} \quad \text{Gravity} \sim \text{order } 1$$

- coupling constant unification

$$\text{Measure } \left. \begin{array}{l} \alpha_{QED}^{-1}(M_z) = 127.922 \pm 0.027 \\ \sin^2 \theta_W(M_z) = 0.2228 \pm 0.0004 \end{array} \right\} \begin{array}{l} \rightarrow \alpha_1 \\ \rightarrow \alpha_2 \end{array}$$

$$\alpha_s(M_z) = 0.1172 \pm 0.0020$$

Plot \rightarrow

- Neutrino Masses : $m_{\nu_i} \ll m_{\ell_i^\pm}$ (more later)

$$\text{see-saw: } m_{\nu_i} = \mathcal{O}\left(\frac{m_{u_i}^2}{M}\right), \quad M = \mathcal{O}(10^{10} \text{ GeV})$$

- Baryon asymmetry

$$\frac{N_B - N_{\bar{B}}}{N_\gamma} \approx 10^{-10}$$

Not possible in SM

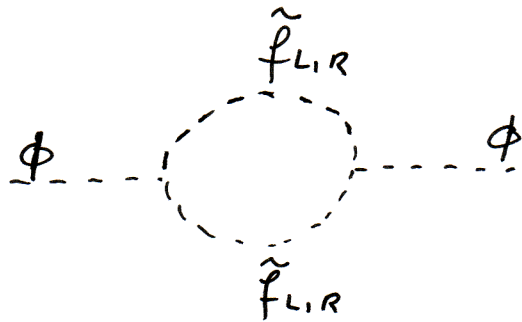
Supersymmetric Solution to the hierarchy problem

RECALL: WE DOUBLED THE SPECTRUM

$$(e_L^-, e_R^-) \oplus (\tilde{e}_L^-, \tilde{e}_R^-)$$

⇒ extra diagrams

$$\mathcal{L}_{\phi\tilde{f}} = -\frac{\tilde{\lambda}_f^2}{2} \Phi^2 (|\tilde{f}_L|^2 + |\tilde{f}_R|^2) - v \tilde{\lambda}_f \Phi (|\tilde{f}_L|^2 + |\tilde{f}_R|^2)$$



Scalar (boson) loops ⇒ opposite sign

$$\begin{aligned} \Pi_{\phi\phi}^{\tilde{f}}(0) &= \tilde{\lambda}_f^2 N(\tilde{f}) \int \frac{d^4k}{(2\pi)^4} \left[\frac{1}{k^2 - m_{\tilde{f}_L}^2} + \frac{1}{k^2 - m_{\tilde{f}_R}^2} \right] \quad (*) \\ &+ (\tilde{\lambda}_f v)^2 N(\tilde{f}) \int \frac{d^4k}{(2\pi)^4} \left[\frac{1}{(k^2 - m_{\tilde{f}_L}^2)^2} + \frac{1}{(k^2 - m_{\tilde{f}_R}^2)^2} \right] \end{aligned}$$

(*) Exactly cancels Λ^2 -divergence, if

$$\tilde{\lambda}_f = \lambda_f, \quad N(\tilde{f}_L) = N(\tilde{f}_R) = N(f)$$

Symmetry

These are just the conditions for supersymmetry!

Same quantum numbers \Rightarrow same couplings

Same degrees of freedom $\Rightarrow N(\tilde{f}) = N(f)$

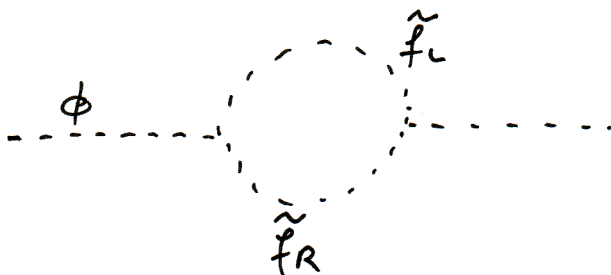
Note: Cancellation of Λ^2 -diverg. independent of $m_{\tilde{f}_{L,R}}^2$

SUSY - BREAKING: (we will come back to this.)

$$\Delta \mathcal{L} = \Delta m_{\tilde{f}_L}^2 |\tilde{f}_L|^2 + \Delta m_{\tilde{f}_R}^2 |\tilde{f}_R|^2 + \left(\frac{\lambda_f}{\sqrt{2}} A_f \phi \tilde{f}_L \tilde{f}_R + \text{h.c.} \right)$$

A_f : tri-linear susy breaking parameter

$\dim[A_f] = \text{GeV}$ (energy)



$$\Delta \Pi_{\phi\phi}^{\tilde{f}}(0) = (\lambda_f A_f)^2 N(\tilde{f}) \int \frac{d^4 k}{(2\pi)^4} \frac{1}{(k^2 - m_{\tilde{f}_L}^2)(k^2 - m_{\tilde{f}_R}^2)}$$

also no Λ^2 -divergence.

Final answer (for $m_{\tilde{f}_L} = m_{\tilde{f}_R} = m_{\tilde{f}}$)

$$\begin{aligned} \Pi_{\phi\phi}^{f+\tilde{f}} &= i \frac{\lambda_f^2 N(f)}{16\pi^2} \left[-2m_f^2 \left(1 - \log \frac{m_f^2}{\mu^2}\right) + 4m_f^2 \log \frac{m_f^2}{\mu^2} \right. \\ &\quad \left. + 2m_{\tilde{f}}^2 \left(1 - \log \frac{m_{\tilde{f}}^2}{\mu^2}\right) - 4m_f^2 \log \frac{m_{\tilde{f}}^2}{\mu^2} - |A_f|^2 \log \frac{m_{\tilde{f}}^2}{\mu^2} \right] \\ &= 0 \quad , \quad \left. \begin{array}{l} \text{for } A_f = 0 \\ m_{\tilde{f}} = m_f \end{array} \right\} \text{SUSY limit} \end{aligned}$$

Note: $m_{\tilde{f}}^2$ - term above

⇒ if SUSY is broken only have

$$m_{\tilde{f}}^2 \leq (1 \text{ TeV})^2 \quad \text{if} \quad m_{\tilde{f}}^2 \leq (1 \text{ TeV})^2 \cdot \left(\frac{16\pi^2}{\lambda^2 N}\right)$$

⇒ SUSY must be broke at $\mathcal{O}(1 \text{ TeV})$
 to solve hierarchy problem

This is the motivation for low-energy SUSY