

The Auger Engineering Radio Array

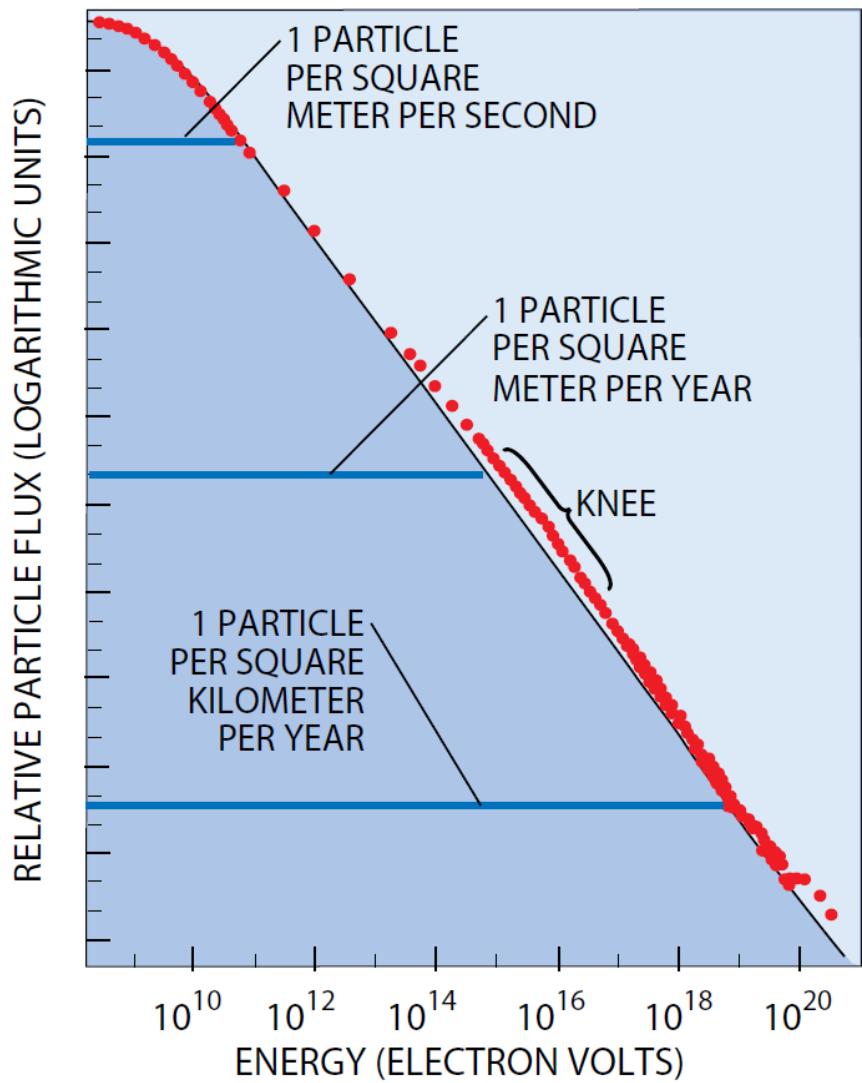


Anna Nelles

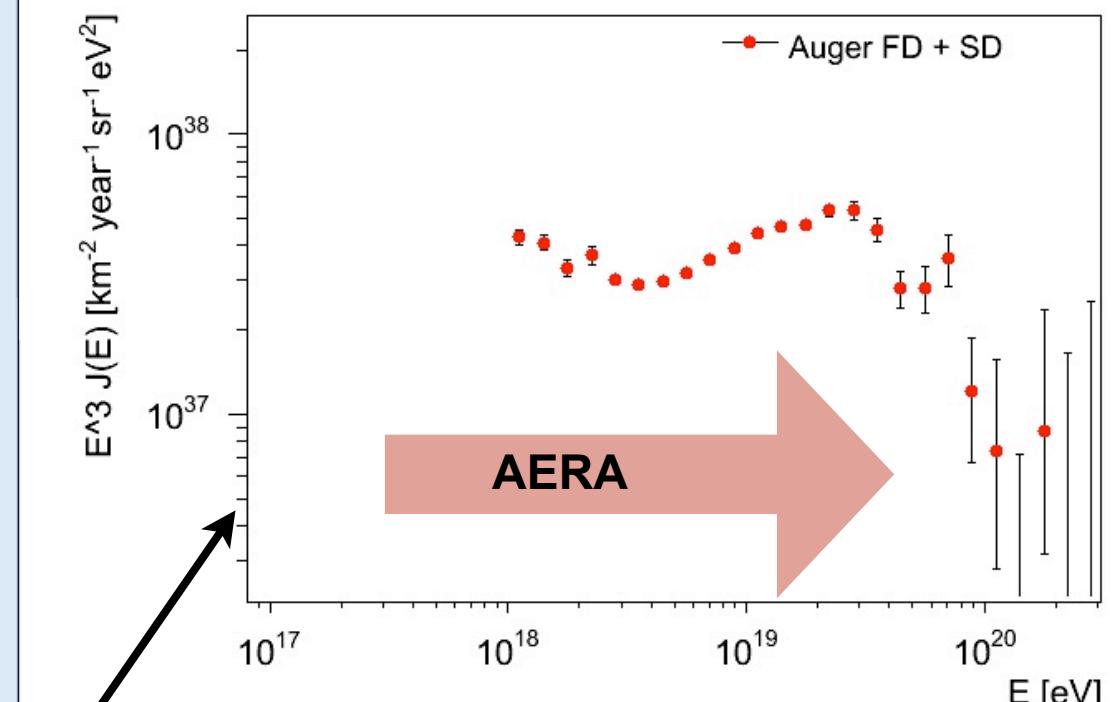
Radboud University Nijmegen



Cosmic Rays

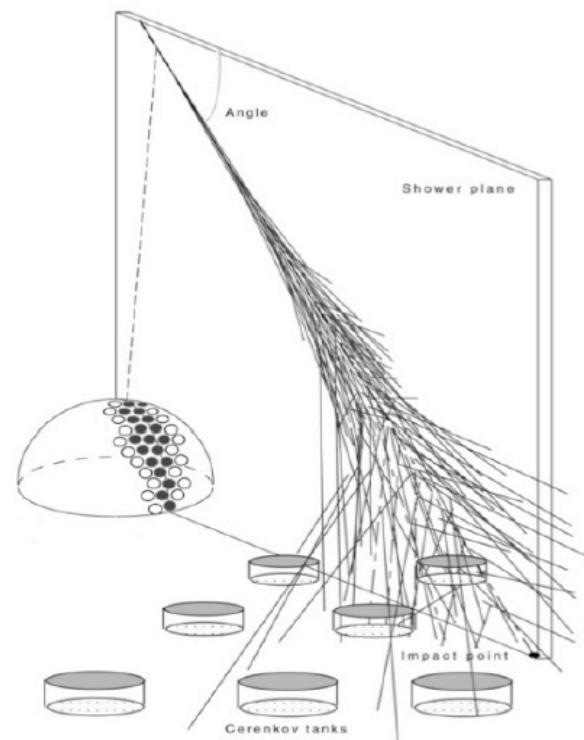


Cronin, Gaisser, Swordy 1997



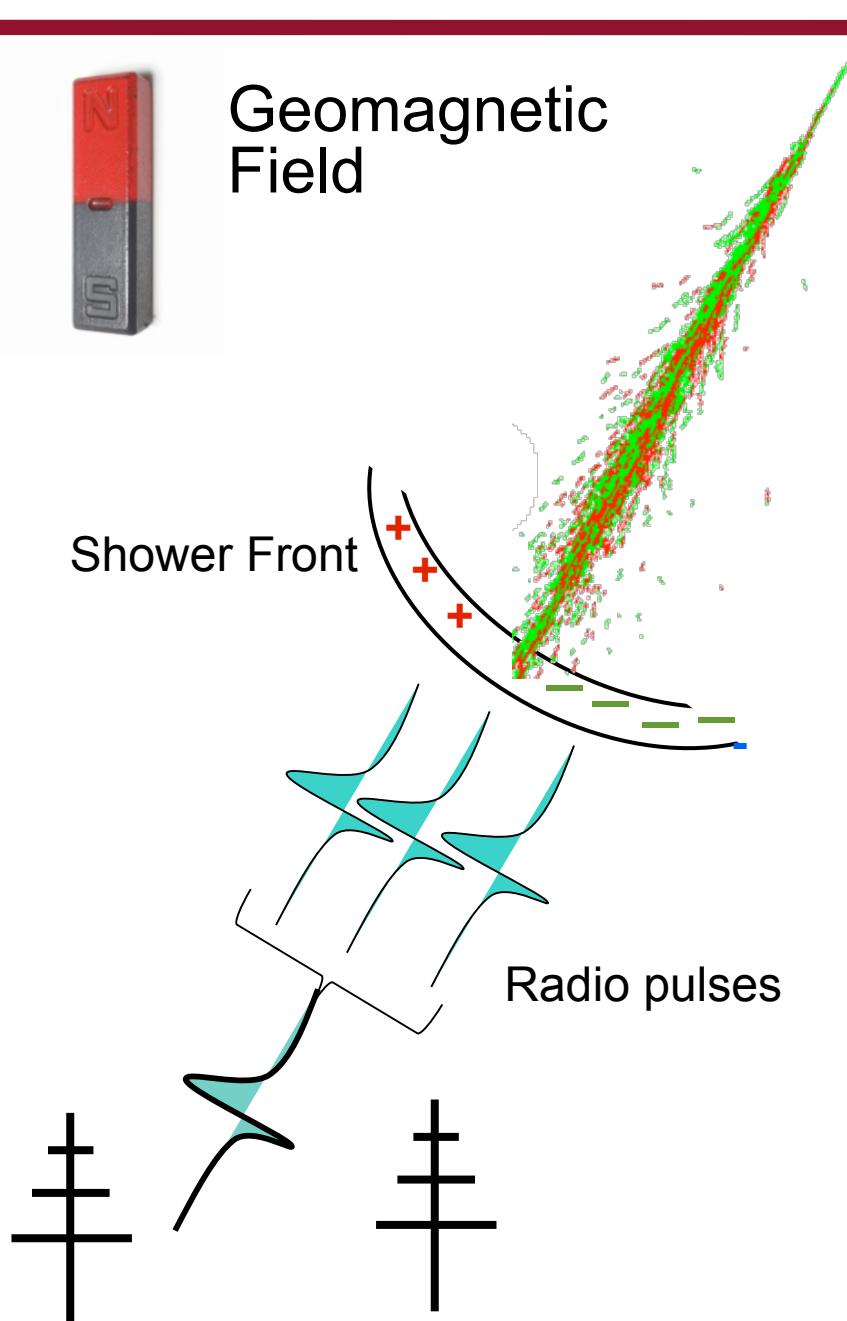
- sources of particles still unknown
- nature of particles not known at these energies

Pierre Auger Observatory



- province of Mendoza, Argentina, 3500 km^2
- 5 Fluorescence Detectors (FD)
 - fluorescent light from nitrogen molecules ($\sim 10\%$ duty cycle)
- 1600 Surface Detector Stations (SD)
 - particle content of shower at ground level ($\sim 100\%$ duty cycle)

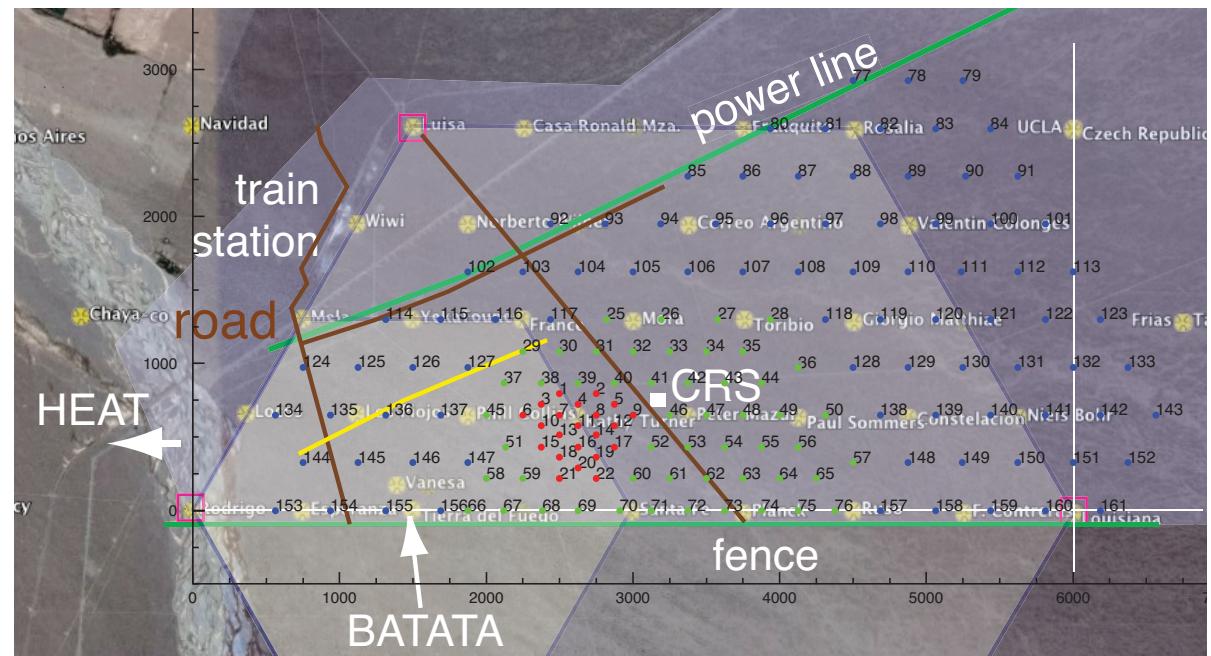
Radio Detection of Cosmic Rays



- electrons and positrons in shower emit coherent radiation
 - geomagnetic synchrotron
 - charge variation
 - ...
- frequencies in MHz range
- enables to observe the shower development
 - new method to study composition of air showers
 - about >90% duty cycle

AERA

- Collaboration of Groups from:
 - France, Germany and the Netherlands
- Colocation with Auger low energy extensions:
 - HEAT (FD), AMIGA (SD)
- Three stages:
- Stage 1
 - 24 antennas
- Stage 2
 - + 52 antennas
- Stage 3
 - + 85 antennas



AERA stations



**Logarithmic-Periodic Dipole Antenna
Bandwidth 30 - 80 MHz
with Low Noise Amplifier and Filter**

GPS Antenna

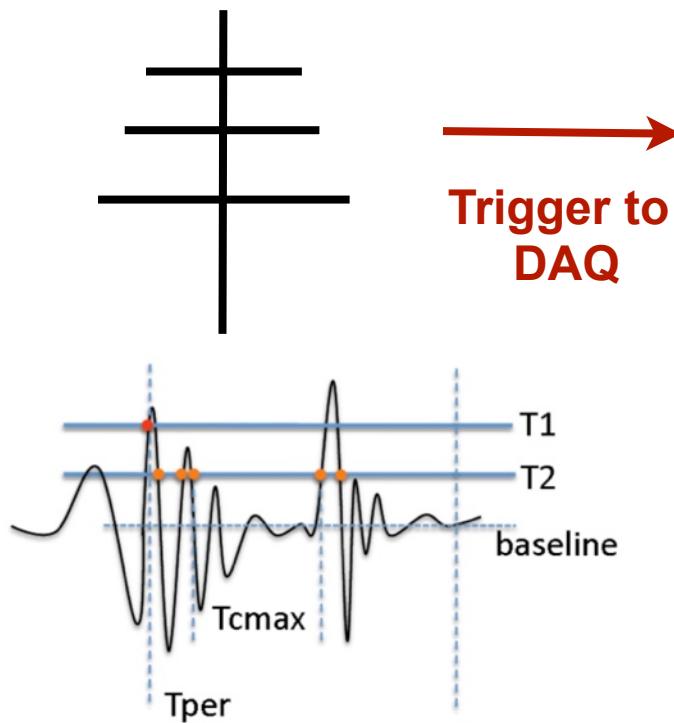
Solar Panels

**Fiber Connection
to Central Station**

**Electronics Box
Read-Out Electronics
Communication
Batteries**

AERA Data taking

On station level:

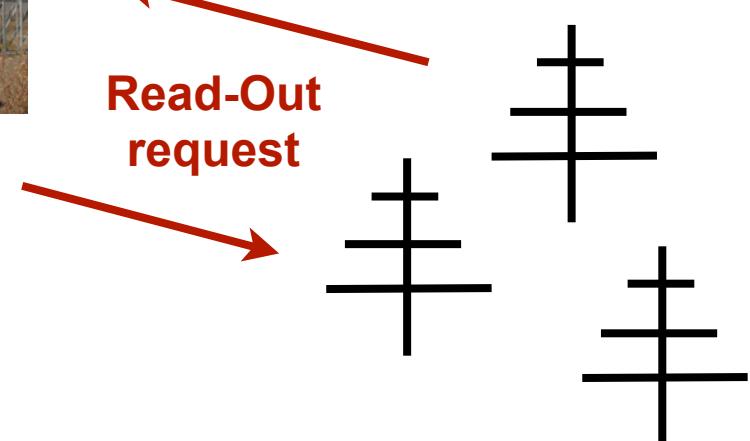


- dynamic thresholds
- noise thresholds
- trigger on either polarization

DAQ in Central Radio Station



Read-Out request

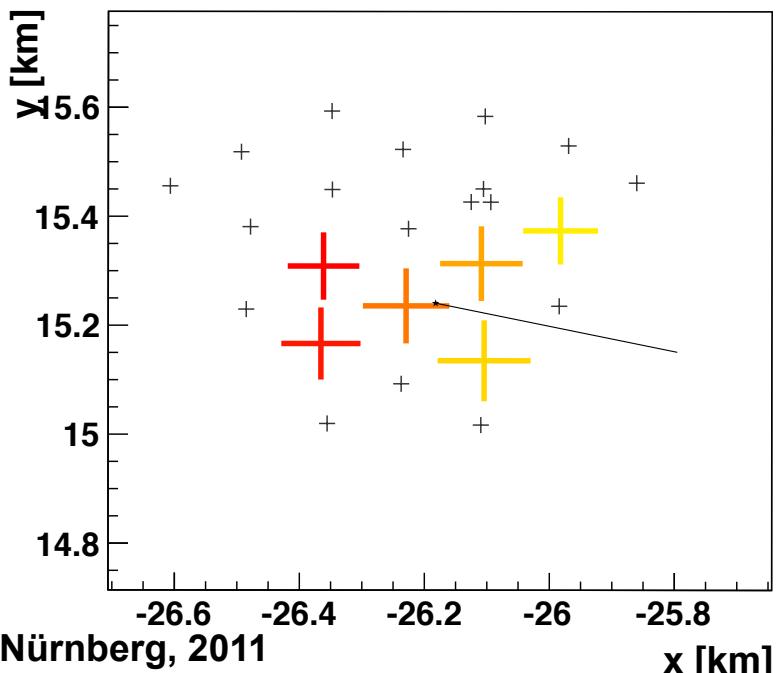
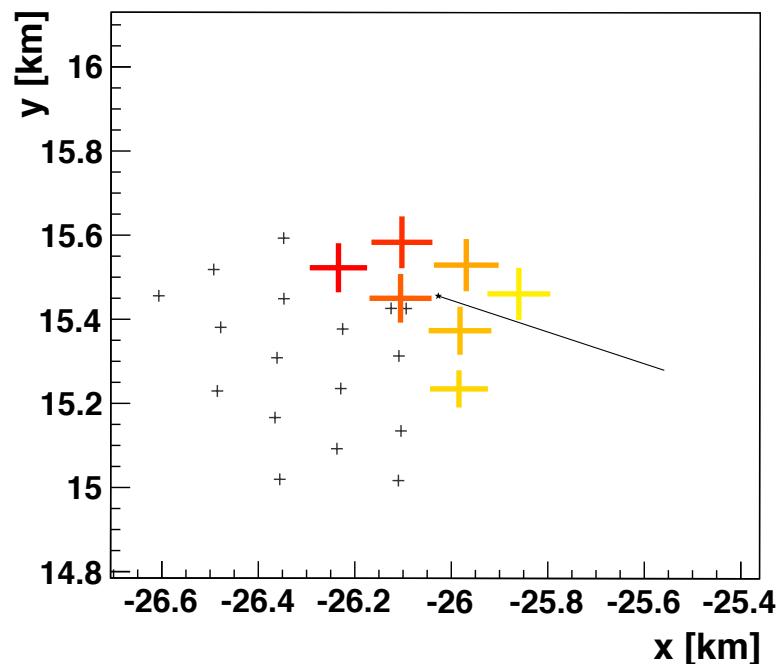
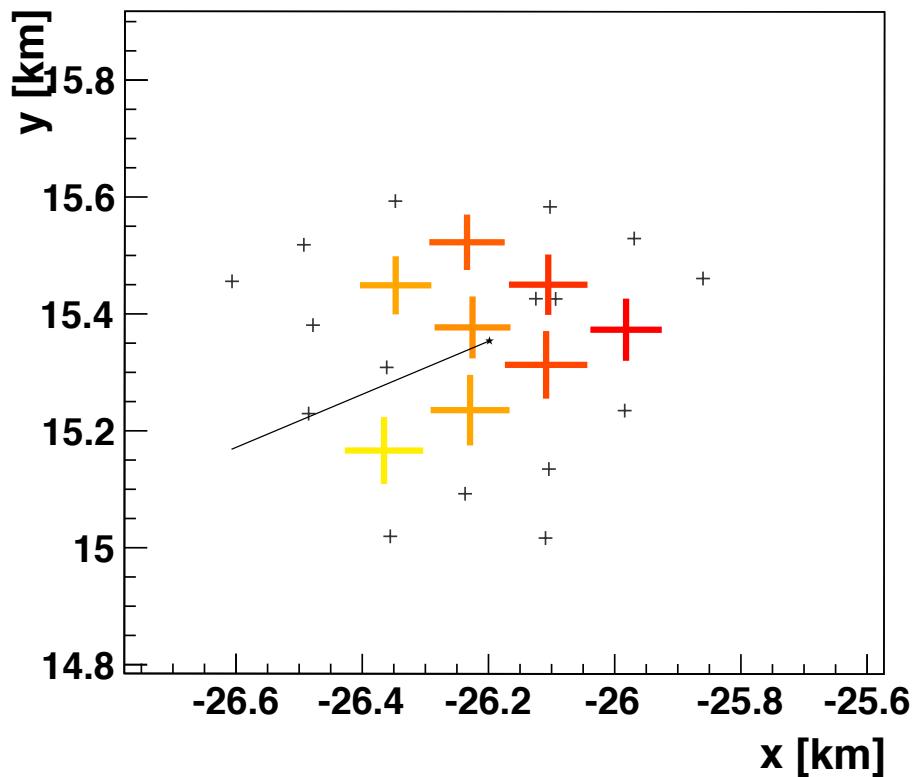


- checking for coincidences in time in > 1 station

- study for optimal time-window
- reduction of data

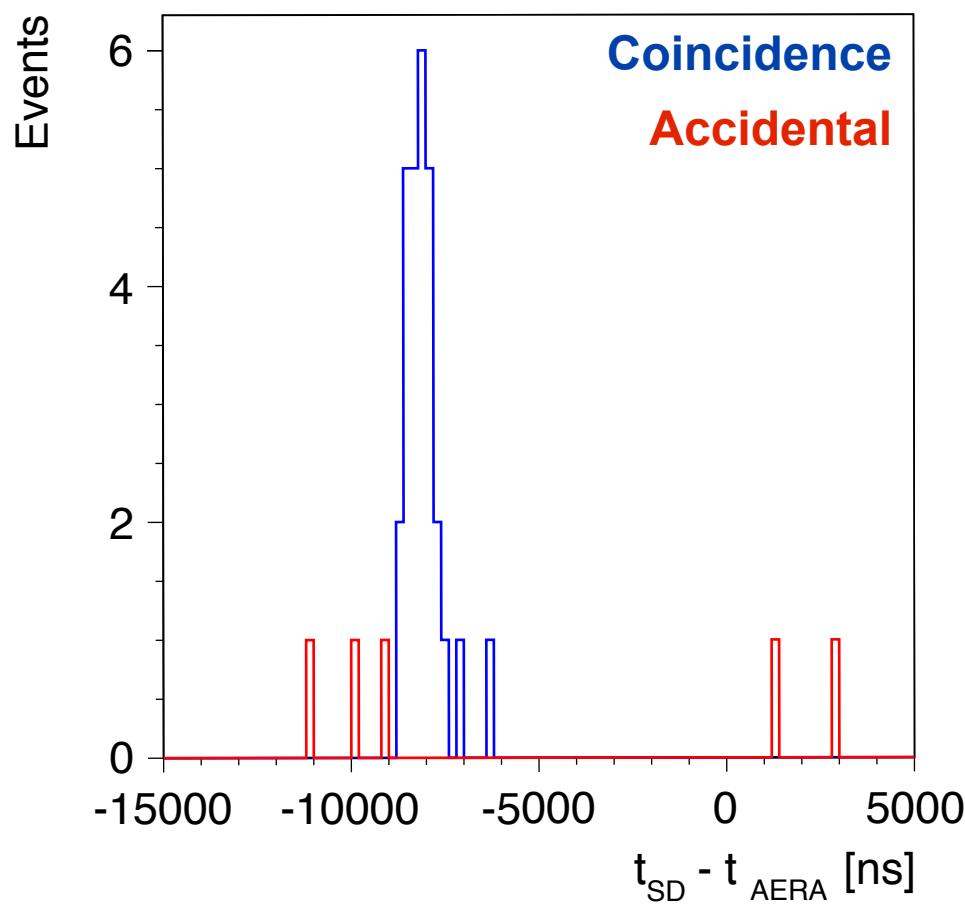
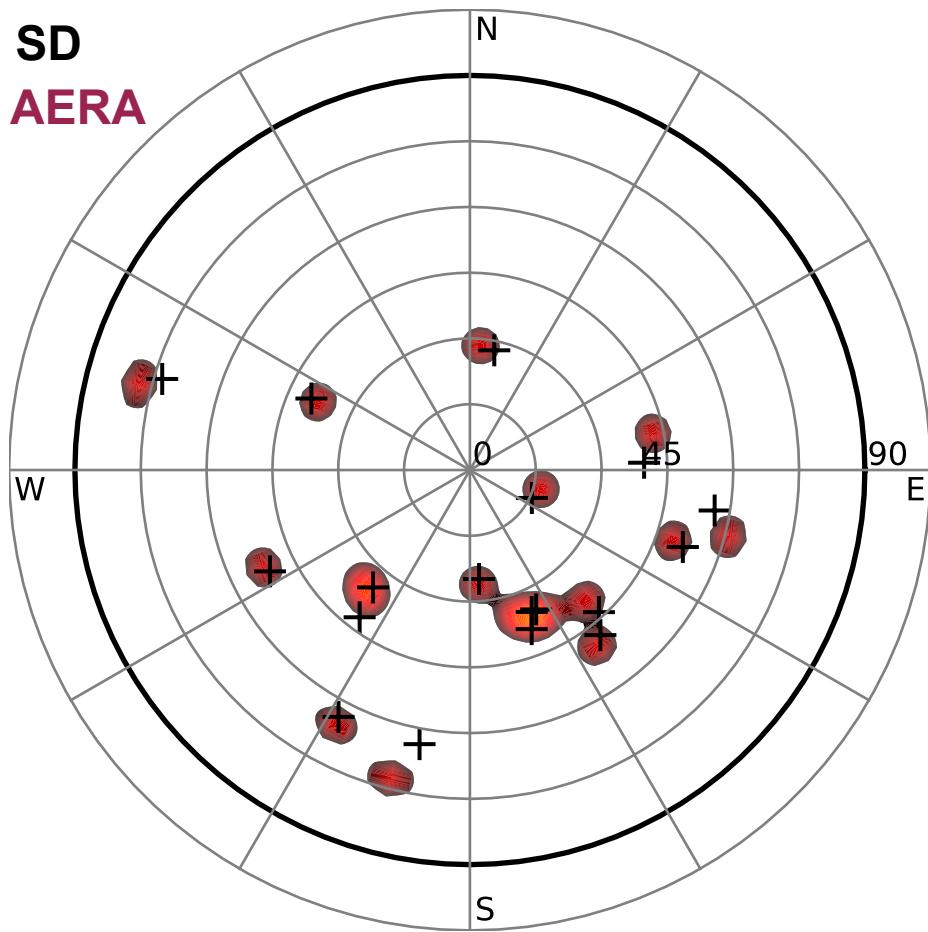
- Event building
- storage of raw data for analysis

Cosmic Ray Events



- Since April 15th:
 - self-triggered cosmic ray events in coincidence with SD
 - 37 events of which 18 at least 3-fold, one with FD

Cosmic Ray Events

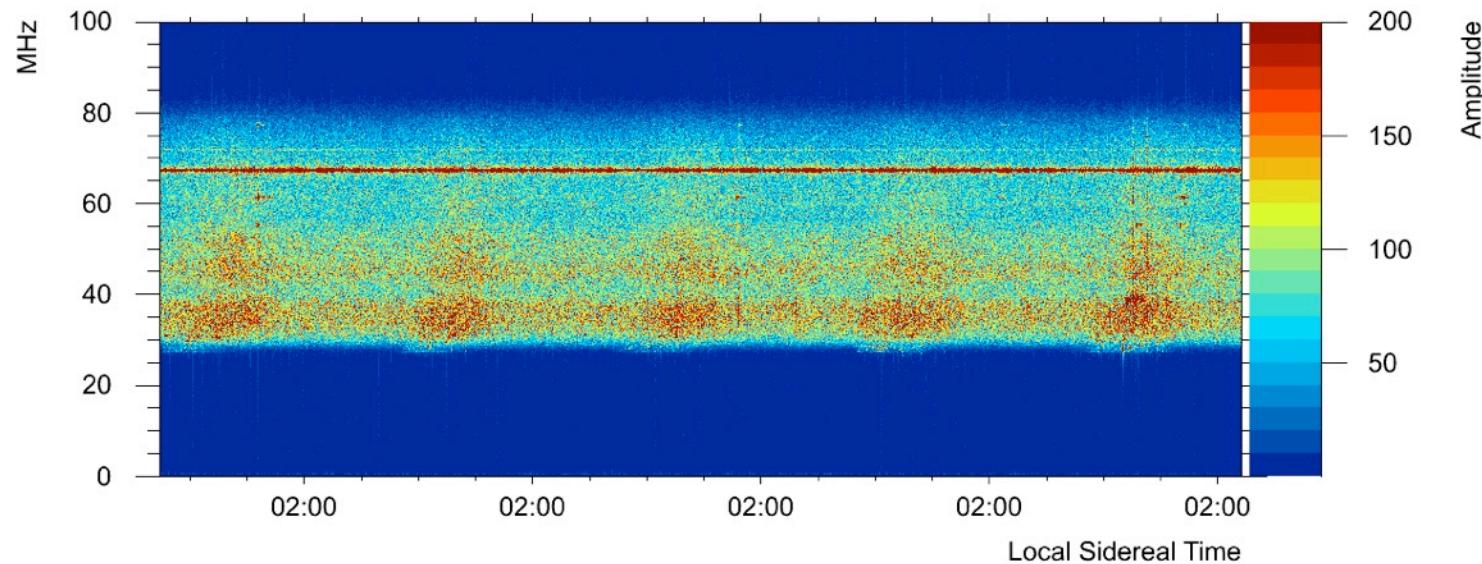
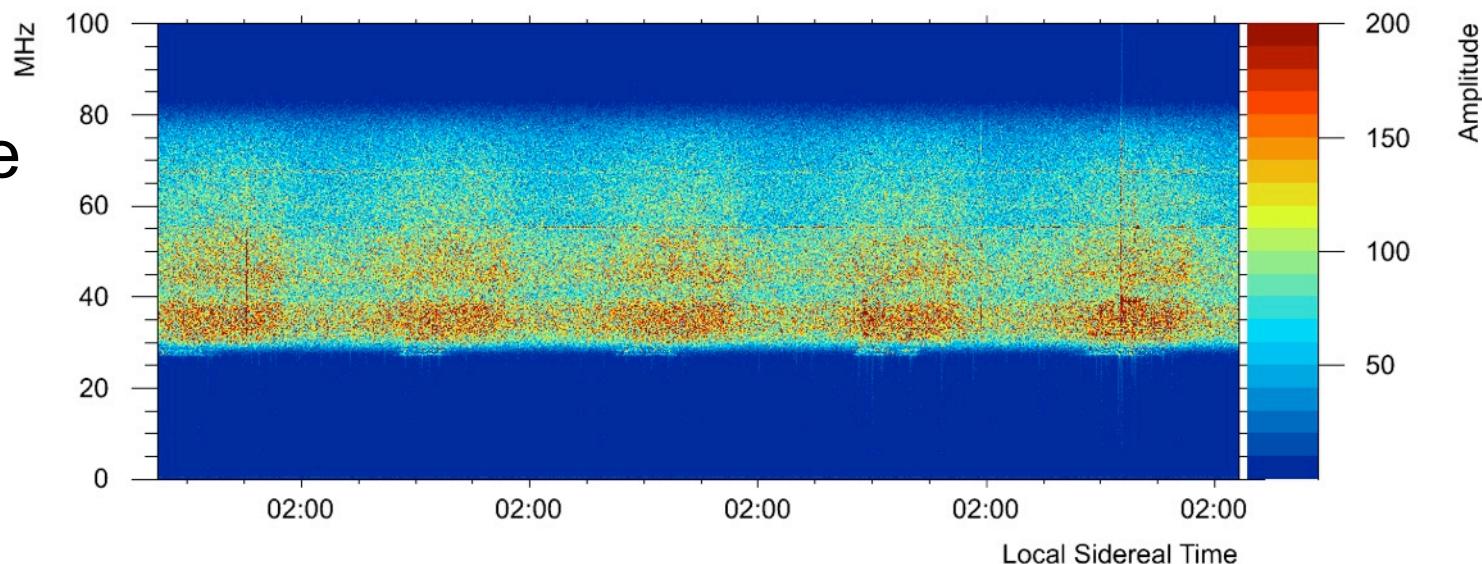


Agreement of directions to within 3 degrees space angle

Very clearly defined timing offset due to trigger conditions

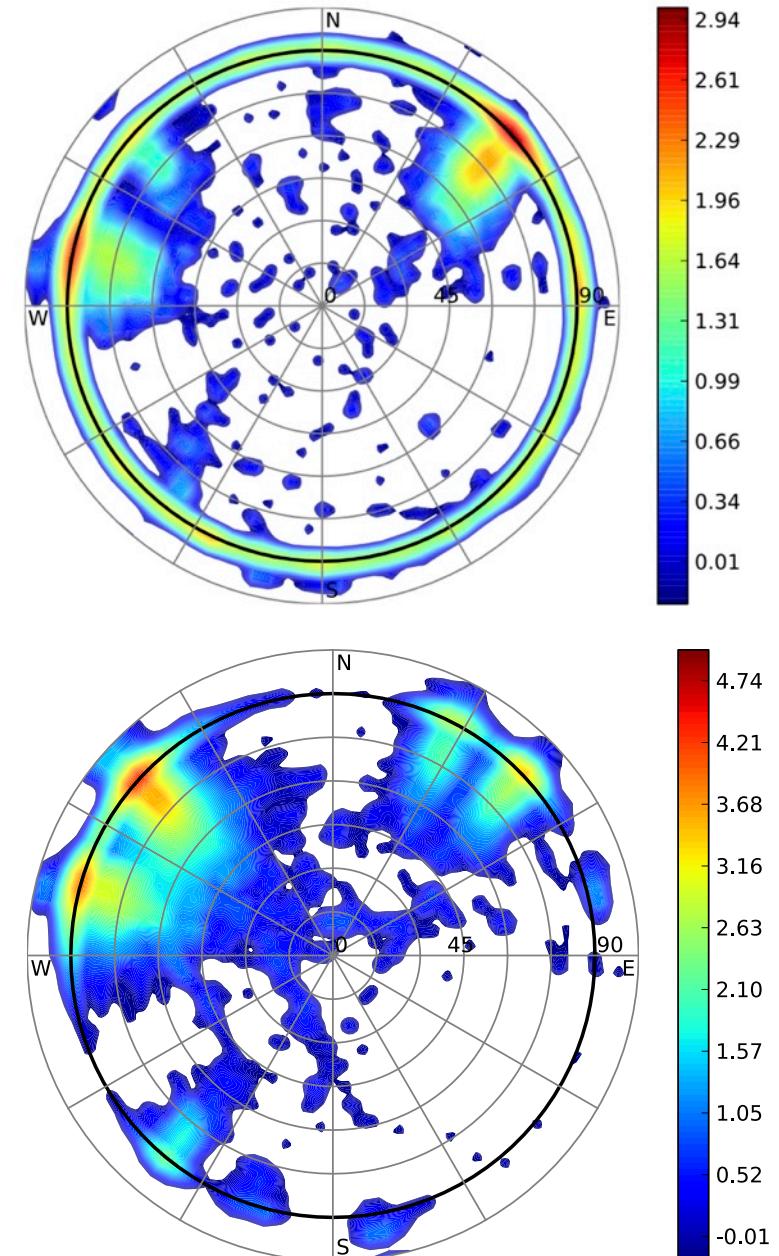
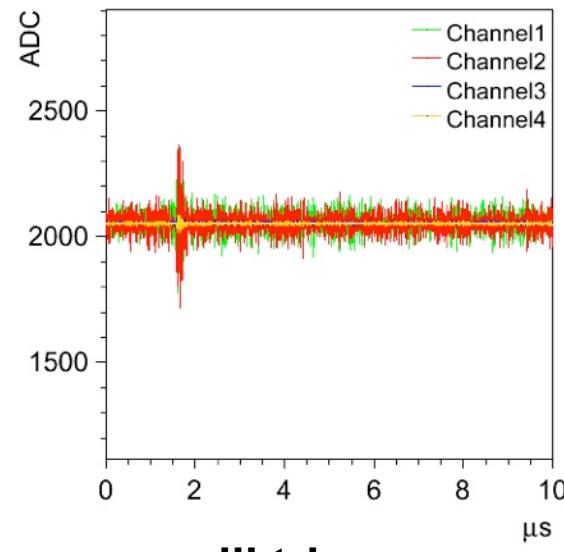
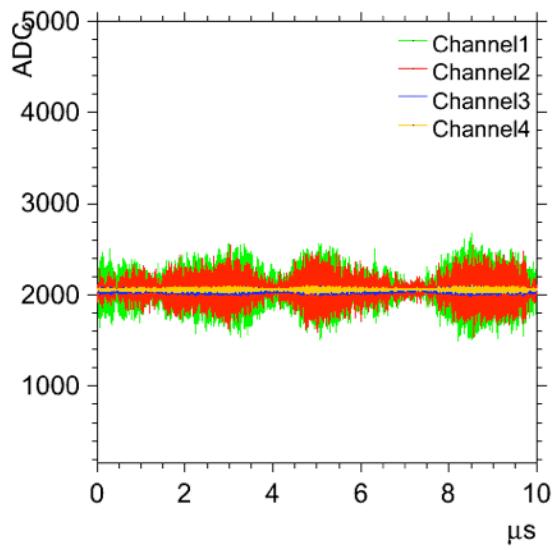
Longterm-Performance

- every 10 s a background trace is recorded
- stations are running stable
- very little narrow-band RFI
- electronics very quiet
- background spectrum dominated by galactic noise



Noise and Transients

- non-negligible amount of radio interference
 - five primary sources
 - probably human made
 - at large zenith angles less interference



Summary and Outlook

- 21 Stations deployed in Argentina
- Stations are operating within Engineering parameters
- Ongoing effort on suppression of interference and background monitoring
- Coincidences of radio self-triggered events with SD observed → first air-showers detected

