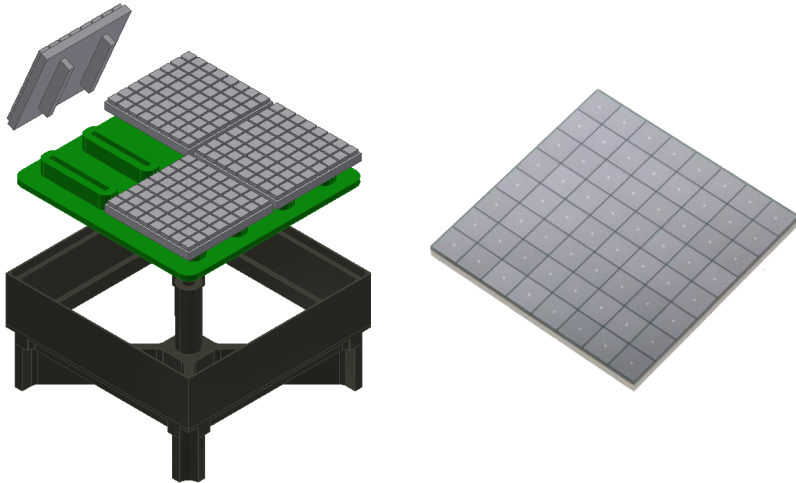


SiPMs for the space-based fluorescence telescope JEM-EUSO

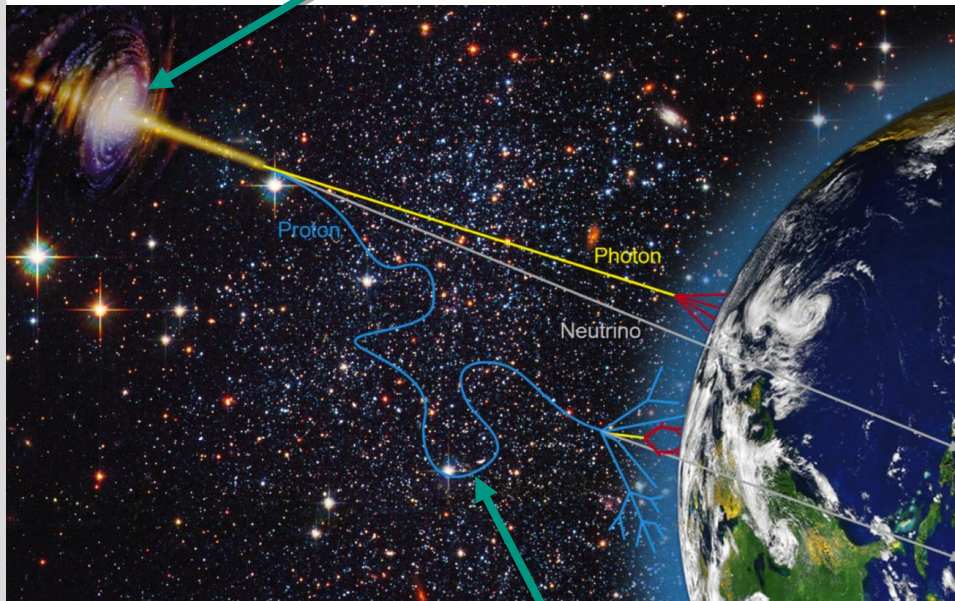
School for Astroparticle Physics | Obertrubach | October 2015

Thomas Huber | Karlsruhe Institute of Technology | Institute of Experimental Nuclear Physics (IEKP)

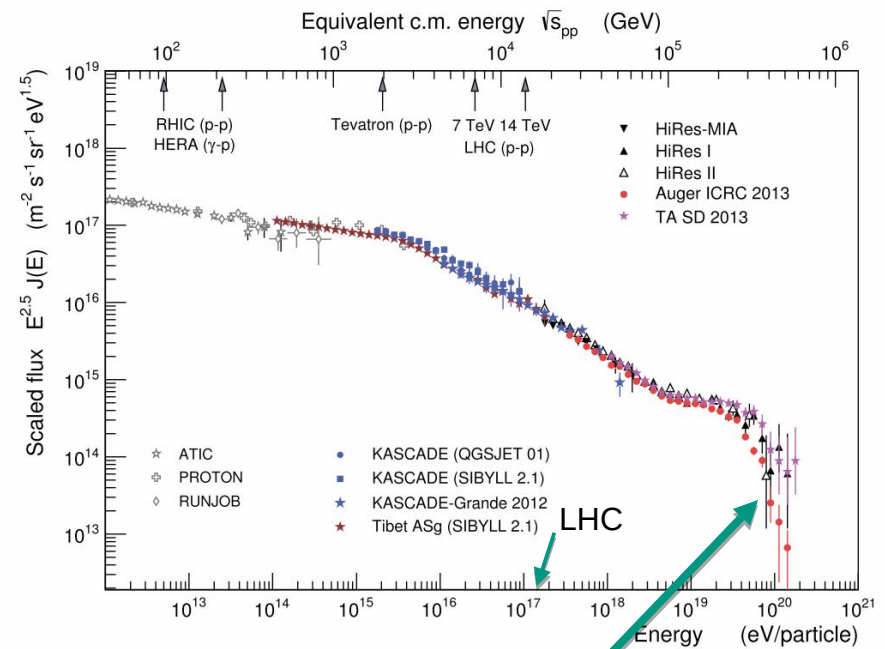


Cosmic Rays

Sources?
Accelerators?

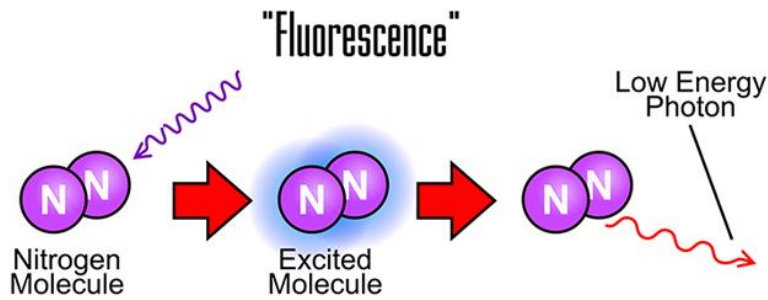
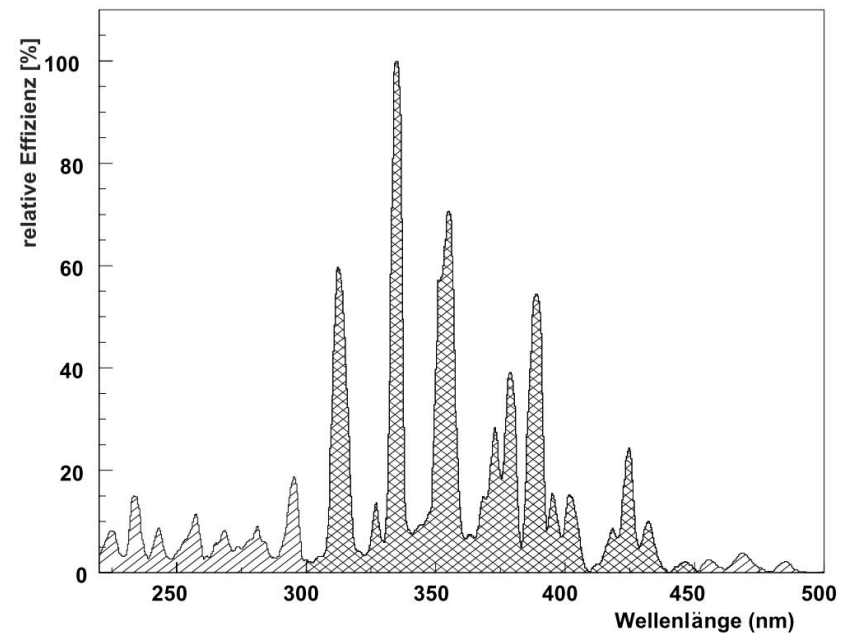
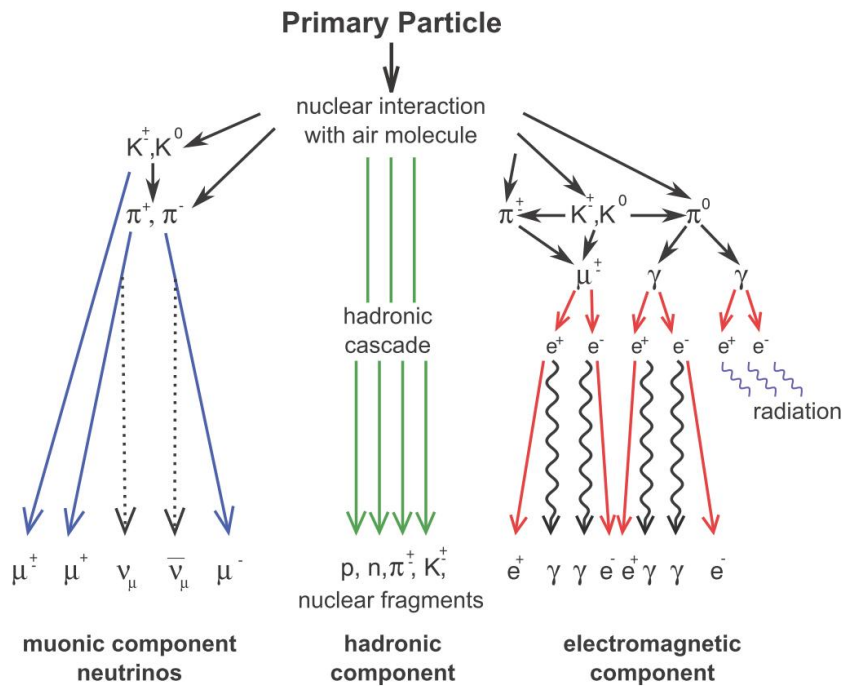


Protons? Nucleons?
Composition?



$$1 \cdot \frac{1}{\text{km}^2 \cdot \text{year}}$$

How to measure? Fluorescence light



Copyright 2014 Telescope Array

JEM-EUSO

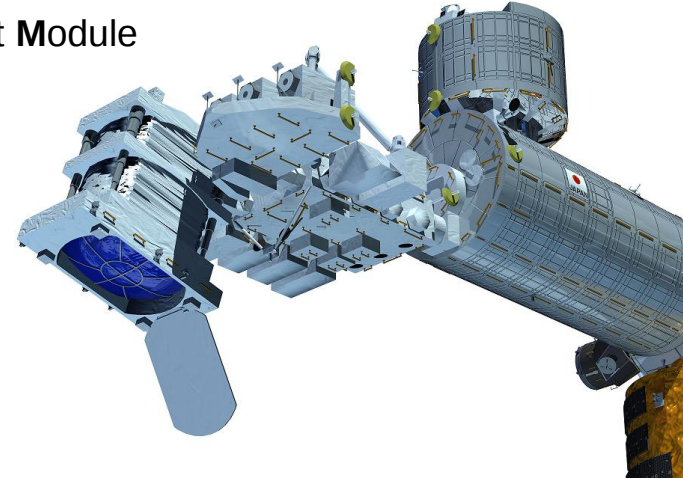
Extreme Universe Space Observatory onboard Japanese Experiment Module

Reconstruction:

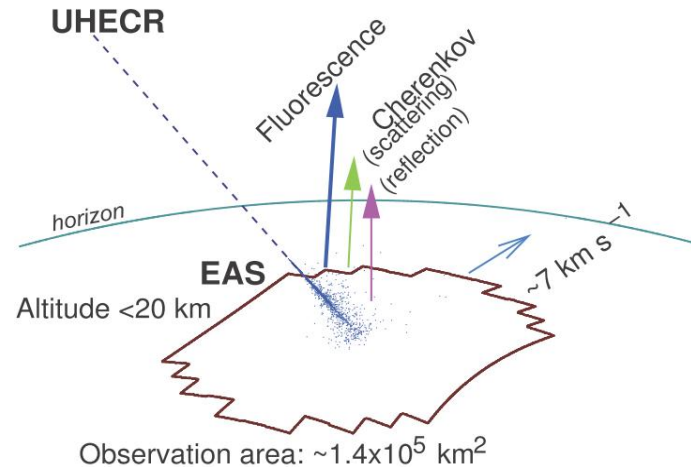
- Arrival directions better than 2.5 deg.
- Energy resolution better than 30%

Observation Area:

- TA, Utah (700 km²)
- AUGER, Malargüe (3.000 km²)
- JEM-EUSO, ISS (140.000 km²)

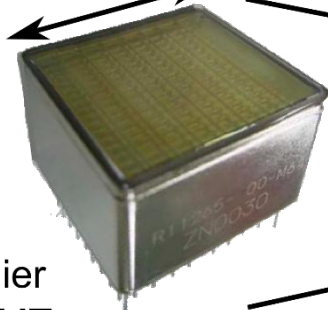


JEM-EUSO  Orbit altitude:
~400km



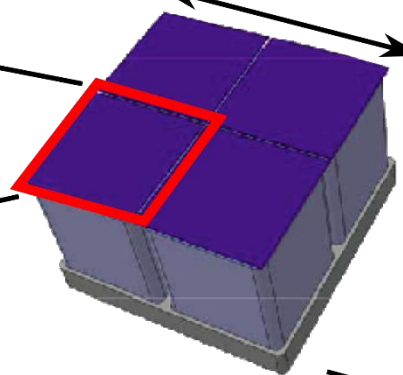
JEM-EUSO – baseline design (MAPMTs)

26 mm



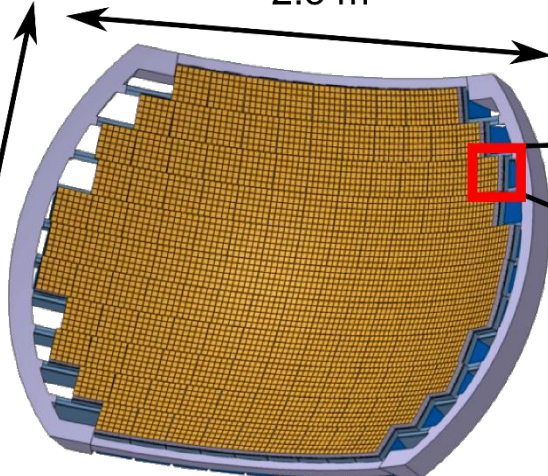
Multi-Anode
Photomultiplier
Tube - MAPMT
8x8 pixel

55 mm



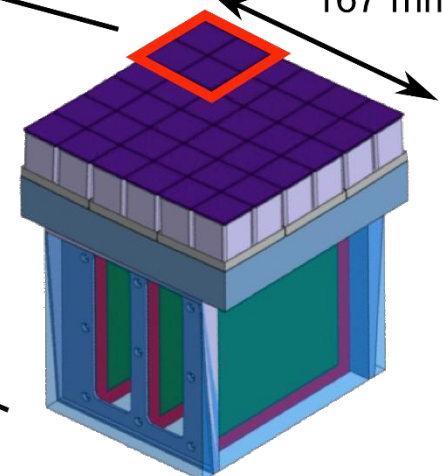
Elementary Cell - EC
& UV-Filter
2x2 MAPMTs = 256 pixel

2.3 m

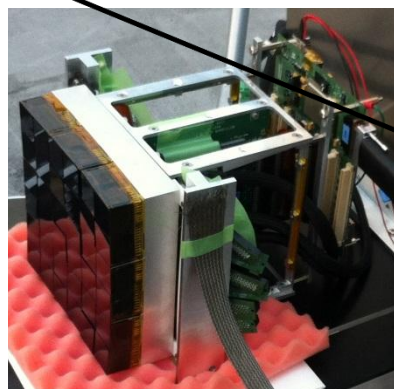


Focal Surface Detector
137 PDMs = 0.3 Mpixel

167 mm

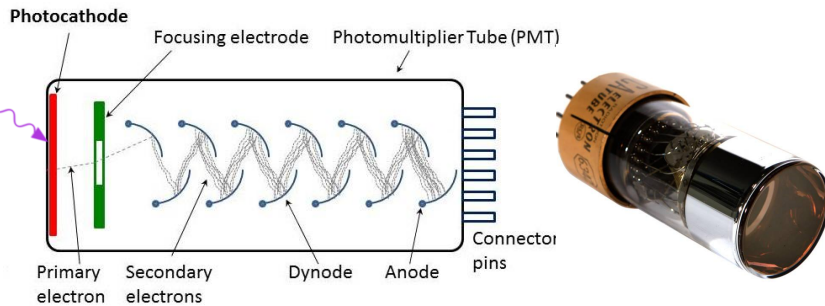


Photodetector Module - PDM
3x3 ECs = 2,304 pixel

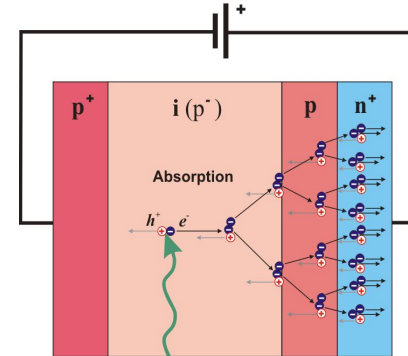


Silicon Photomultiplier (SiPM)

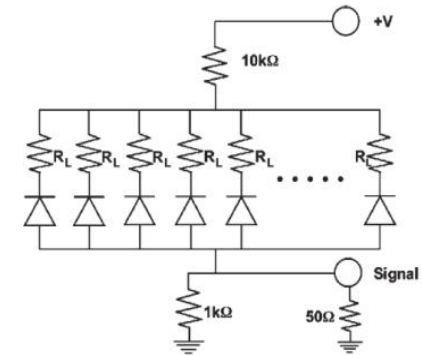
PMT



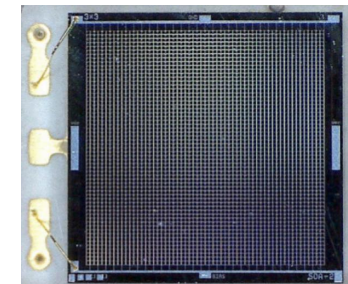
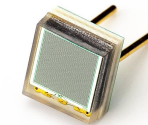
APD



SiPM



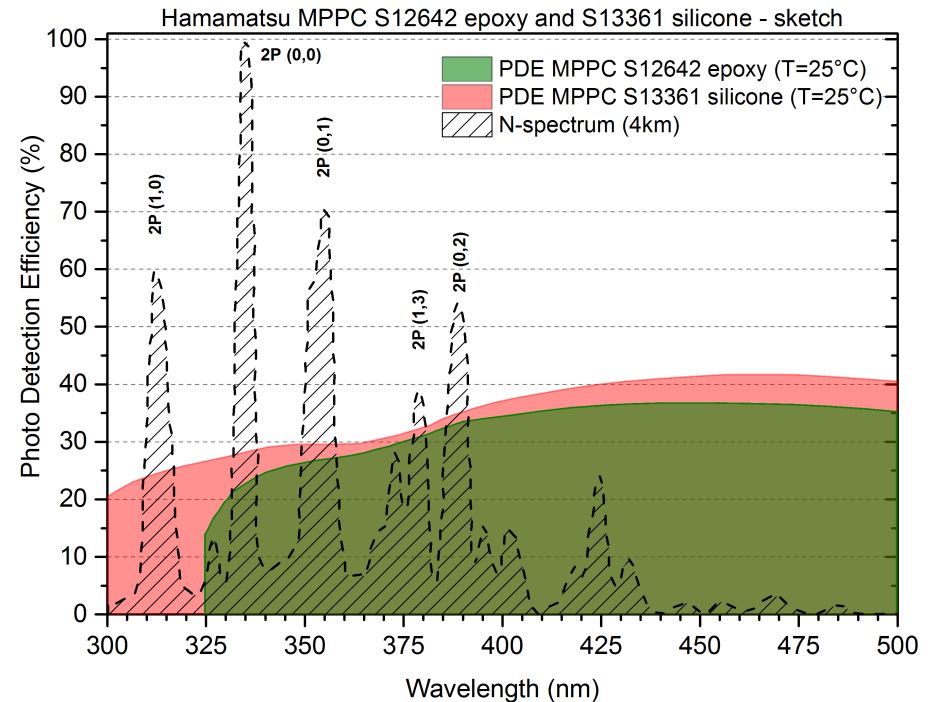
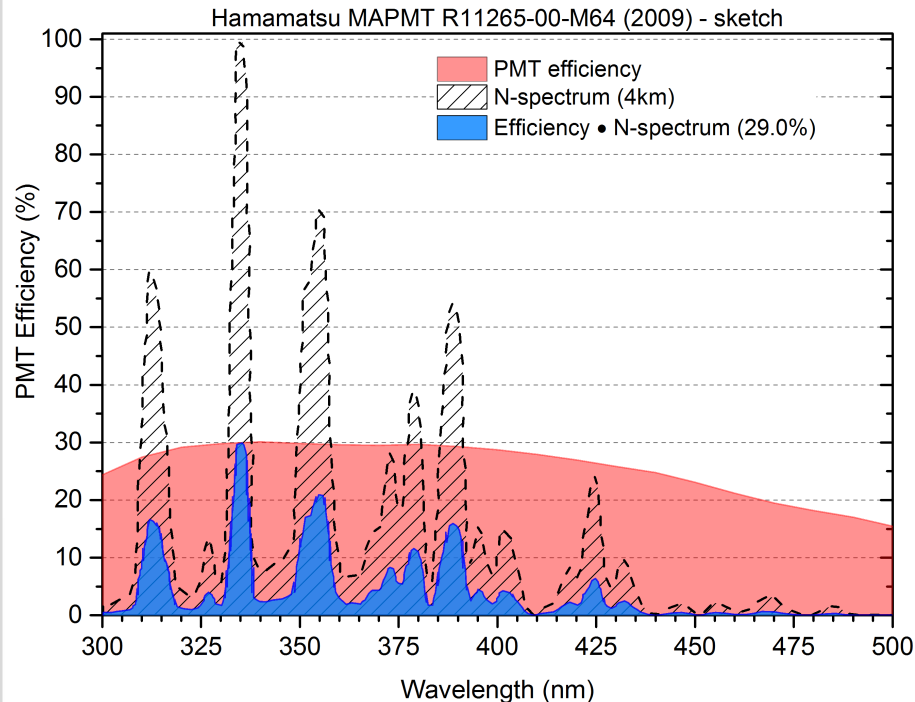
	PMT	SiPM
Photo Detection Efficiency PDE	20-40%	20-60%
Gain	10^6	10^6
TTS (Transit Time Spread)	~1 ns	~1 ns
Dynamic range	10^6	10^3
Dark noise rate	~Hz 😊	~MHz 😞
Behavior in magnetic fields	😞	😊
Operation Voltage	1000+ V 😞	50-70 V 😊
Temperature sensitivity	😊	😞
Robustness and compactness	😞	😊



← 3 mm →

$$\text{PDE} = \frac{\text{Number of detected photons}}{\text{Number of incident photons}}$$

Motivation: MAPMT vs. SiPM



Ratio of MAPMT/SiPM area and nitrogen spectrum area:

MAPMT:
~29%

SiPM:
~31%

Spectrum from AUGER Design Report 1997

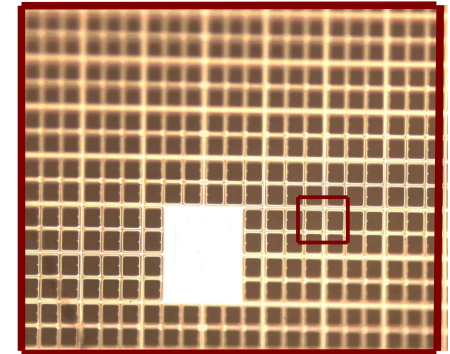
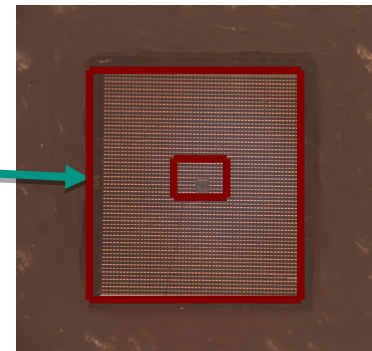
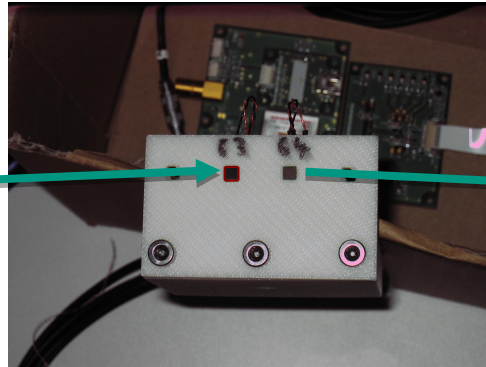
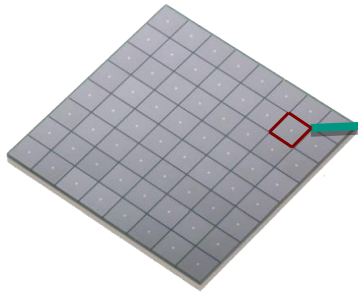
Our Candidate:

Candidate:
 Hamamatsu
 64 Pixel
 SiPM TSV-Array

1 Pixel from the
 TSV Array

2.5x Zoom into 1 Pixel

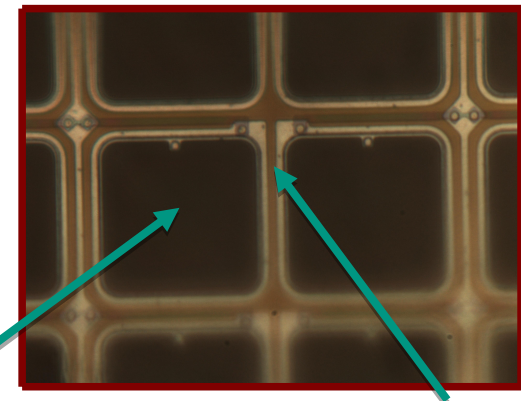
10x Zoom



Hamamatsu S13361-3050

Gain:	$\sim 10^6$
Low Bias Voltage:	$\sim 53V$
Number of APDs:	~ 3600
Darkcountrate:	~ 1 Mcps
Crosstalk	$\sim 3\%$

100x Zoom



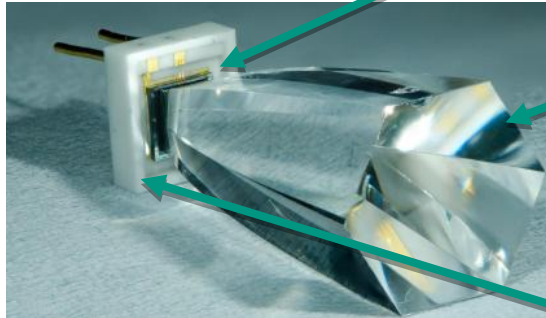
1 Avalanche Photodiode
 (of est. 3600 @ 1 pixel)

Crosstalk-reducing
 Isolator

Our Candidate:

FACT / FAMOUS:

3 mm Gap
between 2 pixels



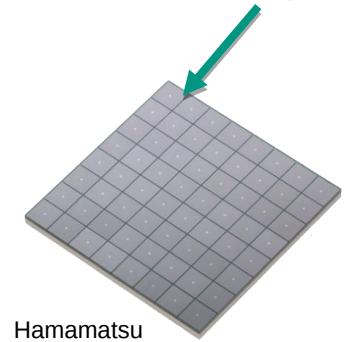
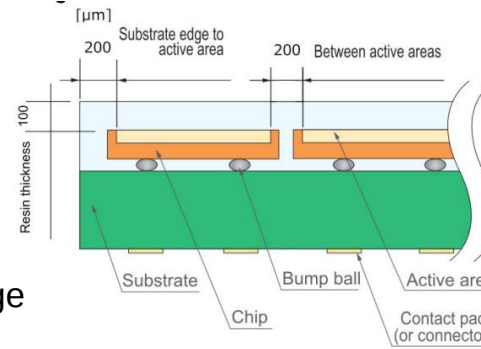
Winston cone

Daniela Dorner HAP Workshop Zeuthen 2014

Hamamatsu SiPM
MPPC ceramic package

New TSV-MPPCs SiPM :
TSV = Trough Silicon Via

0.2 mm Gap
between 2 pixels



Hamamatsu

Absolute maximum ratings

Parameter	Symbol	S12		
		-025C	-050C	-100C
Operating temperature*1	Topr		-20 to +40	
Storage temperature*1	Tstg		-20 to +60	

Hamamatsu Datasheet



...for operating outdoor

Check for thermal induced changes in

- PDE
- Bias-Voltage
- Darkcounts
- Structural changes

Investigation with AFM
before/after cooling

Absolute maximum ratings

Parameters	Symbol	S12642	S12642	S12642
		-0404PA-50	-0808PA-50	-1616PA-50
Operating temperature	Topr		0 to +40	
Storage temperature	Tstg		-20 to +60	

Hamamatsu Datasheet

???

- structural problems?
- fear of condensed water?
- loss of performance?



TSV- SiPM

Experimental Setup for Bias-Voltage / Darkcounts

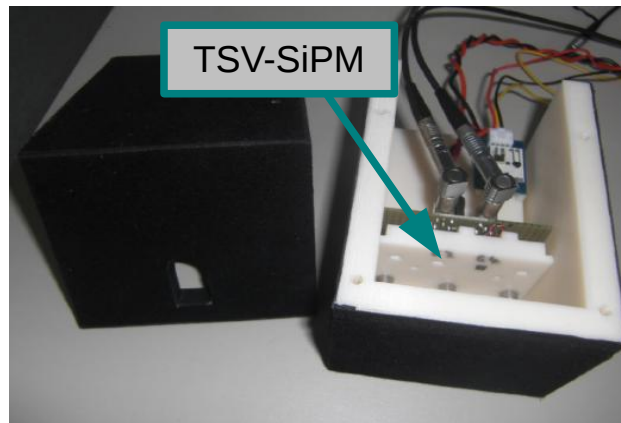
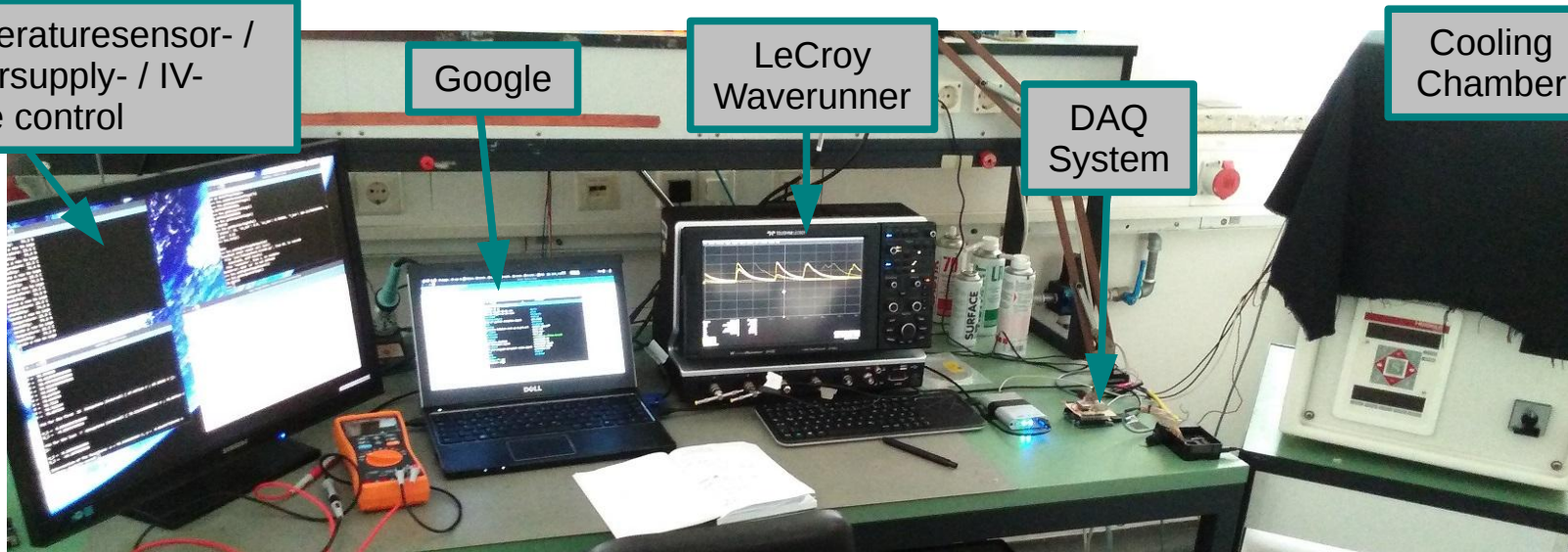
Temperaturesensor- /
Powersupply- / IV-
Curve control

Google

LeCroy
Waverunner

DAQ
System

Cooling
Chamber

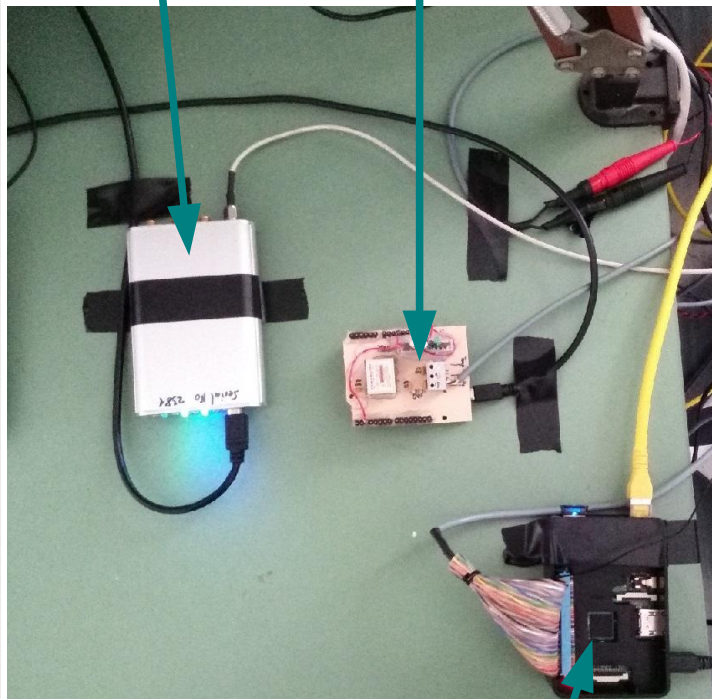


- TSV-SiPM testing device with attached temperature/humidity sensor and one part of the readout electronics
- 2x 3mm², 1x 2mm², 1x 6mm² TSV SiPM sockets to compare them simultaneously
- Photon shielding is overlapping the corners.

Experimental Setup for Bias-Voltage / Darkcounts

DRS4 Evaluation Board

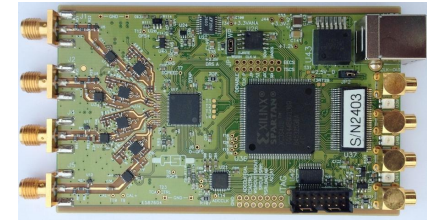
Arduino UNO Rev 3



Raspberry Pi B+

DRS4 Evaluation Board:

- Up to 5 GS/s
- DAQ SiPM Signals



4 Channel ADC

Arduino UNO Rev 3:

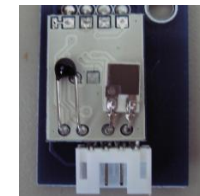
- Mounted with Hamamatsu SiPM Power Supply
- Control of Bias-Voltage via Python interface
- Current Monitor



C11204-01
Output Voltage:
50V to 90V

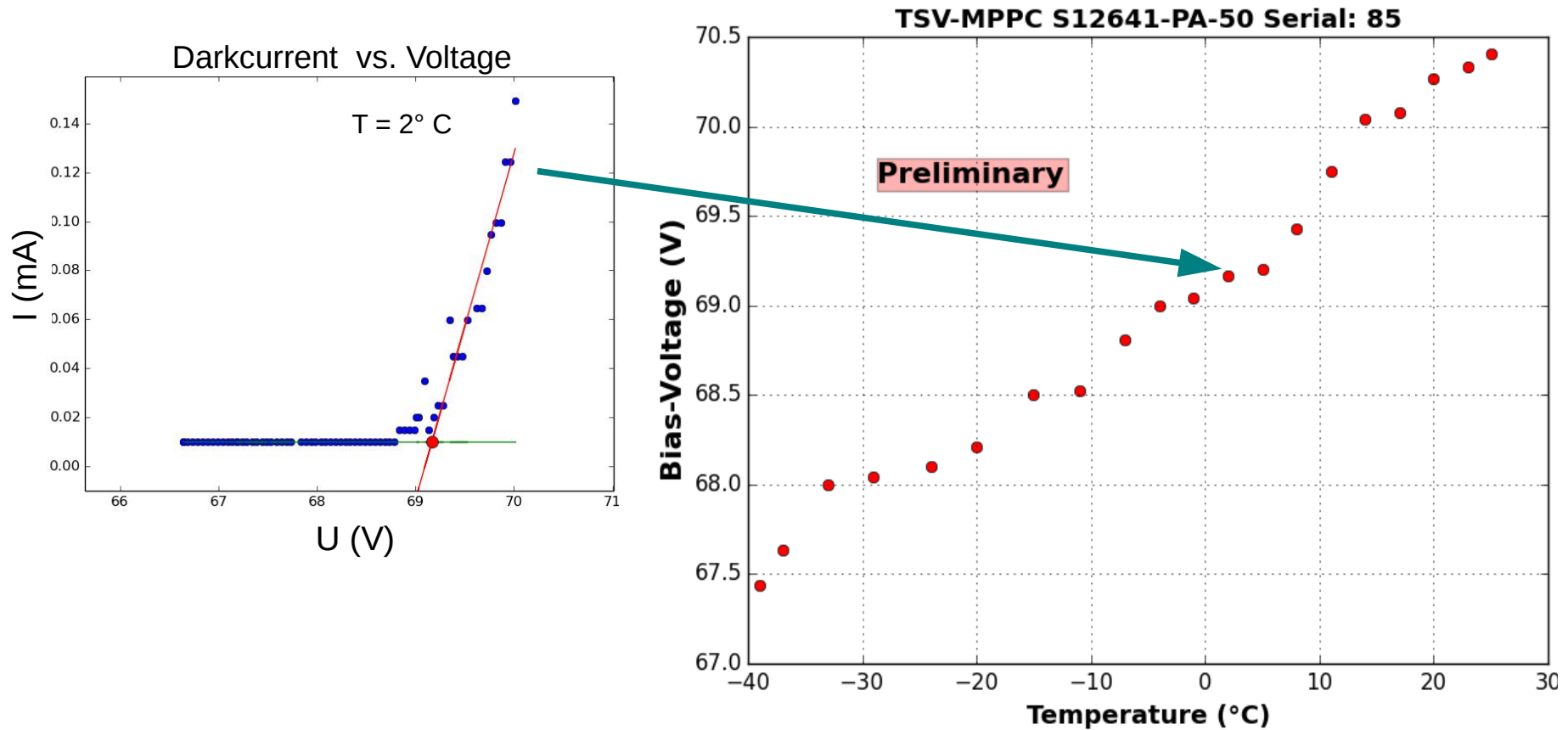
Raspberry Pi B+

- Connected via GPIO to temperature sensor
- Monitoring temperature. Written in Python; Controlled via SSH; Saving data via SQL to a local webserver



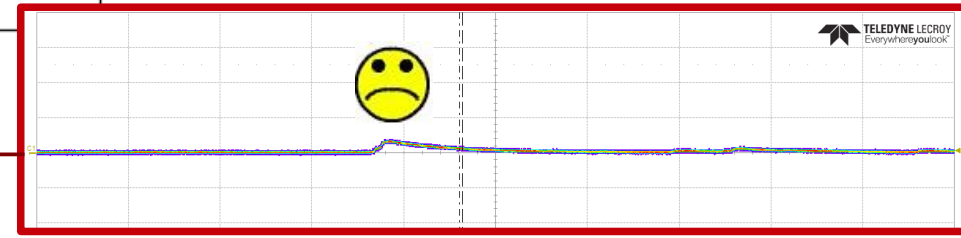
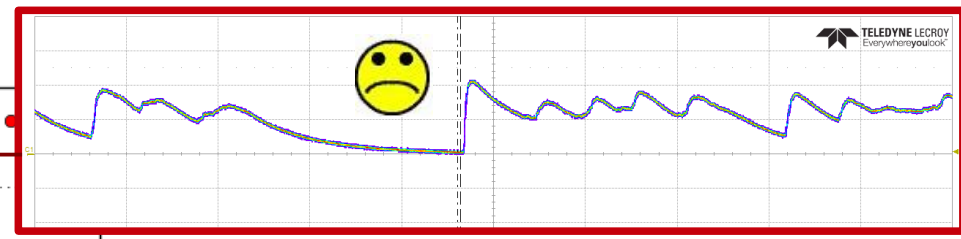
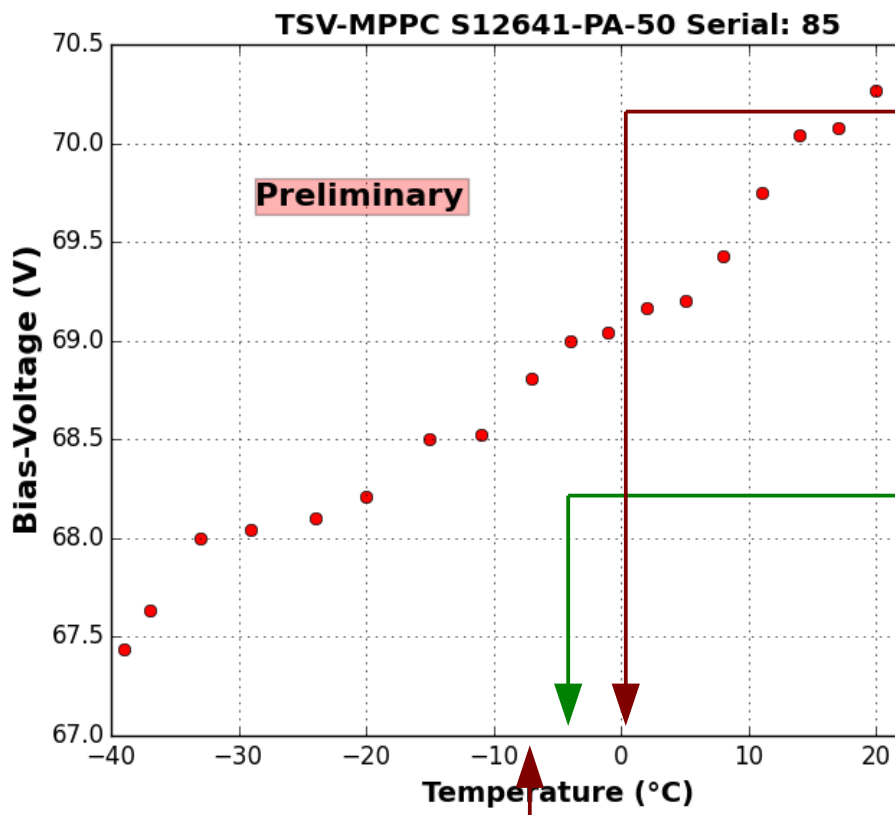
AM2303
 $\pm 2\%$ of humidity level
 $\pm 0.3^\circ\text{C}$ temperature

Temperature dependency of Bias-Voltage



Consequence of not adjusting the Bias-Voltage

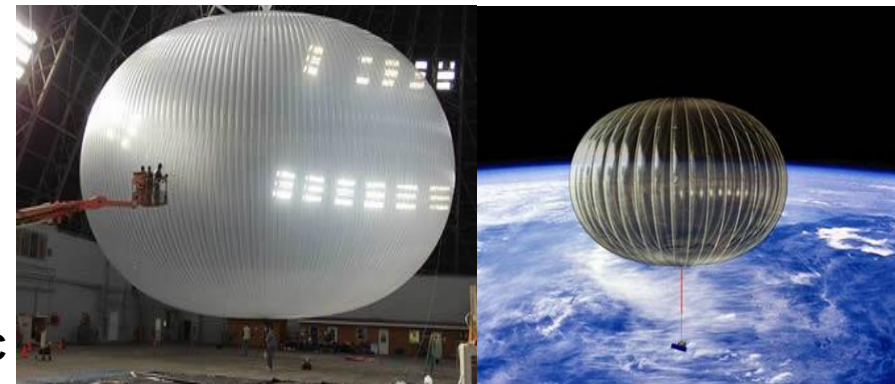
With a constant Bias-Voltage:



V: 50mV/div Timebase: 500ns/div

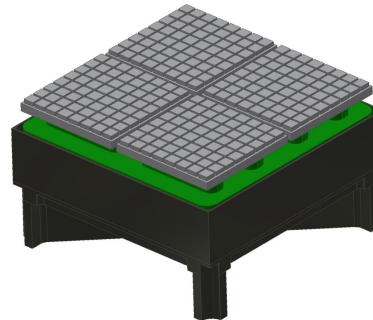
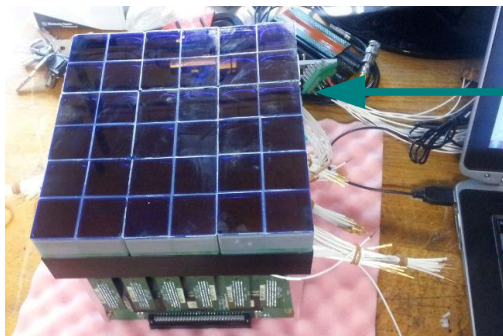
Pathfinder Mission: NASA SPB Flight 2017

- Payload 1000kg of science instruments
- Sustained flight altitude: ~33km
- ~ 1,5km altitude variation during flight
- Flight duration ~ 100 days



MAPMT PDM

SiPM EC



Delivery to NASA: July 2016

Possibly flight in spring 2017 from Wanaka, NZ

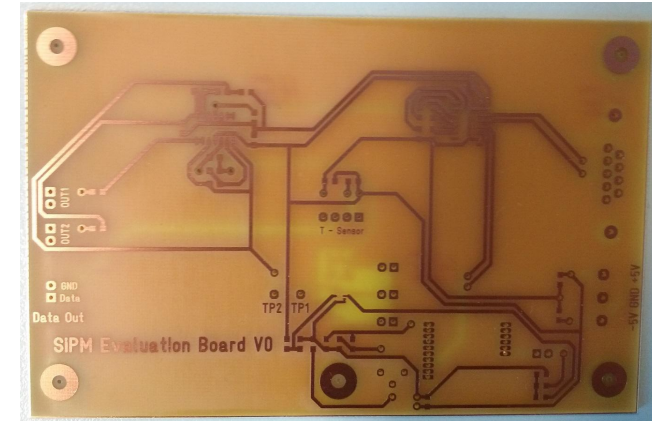
NASA's first Super Pressure Balloon flight, March 2015, Wanaka, NZ:

- Flight duration 32 Days

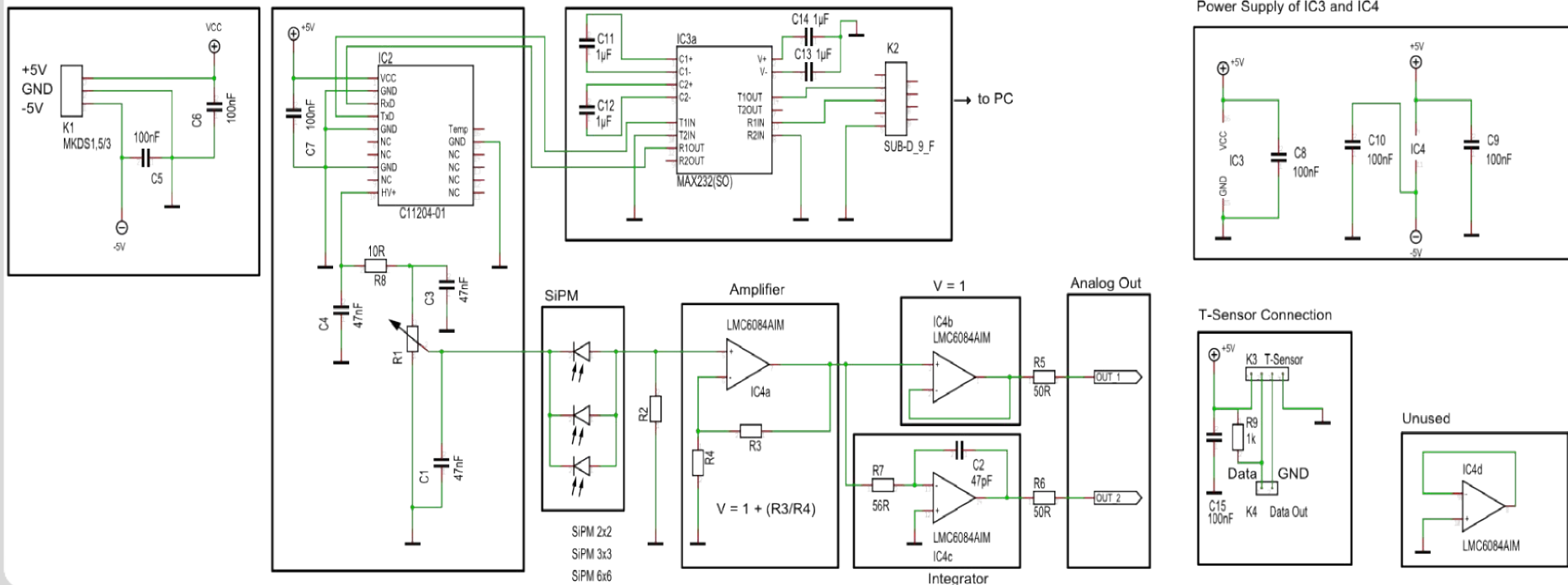


First step: TSV-SiPM Readout board

- For PDE / Gain Measurement
- Integrated amplifier (via OP)
- Temperature sensor
- Voltage adjustment
- Fast readout for investigate time resolution
- Integrated readout for charge measurements



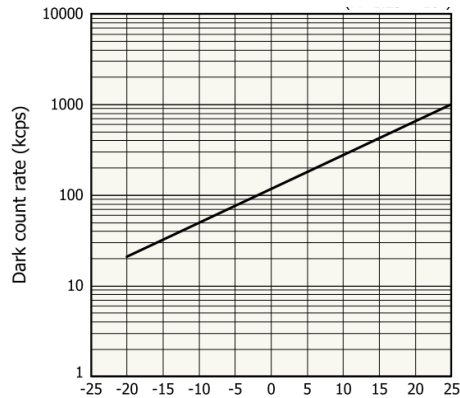
Power Supply of IC3 and IC4



Outlook

Darkcounts of TSV-SiPMs:

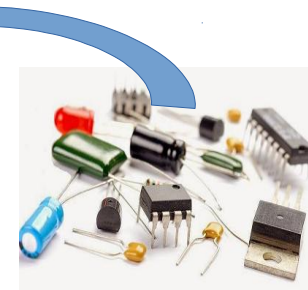
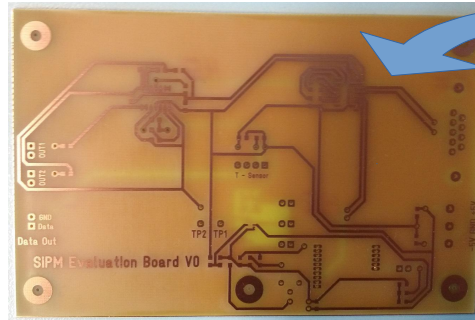
3mm² Area, pitch 50μm, typical example



Ambient temperature (°C)

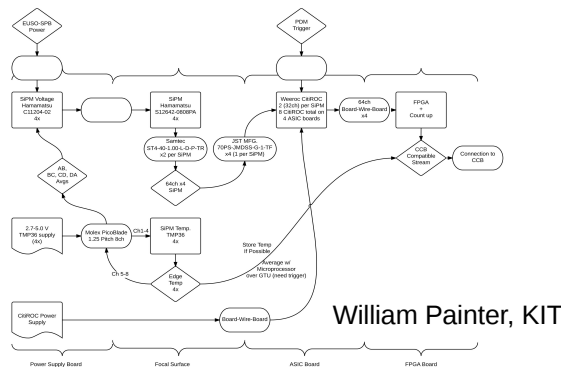
$$N_{DC} \approx T^{\frac{3}{2}} \cdot e^{\frac{E_g}{2kT}}$$

Assembling the readout-board:

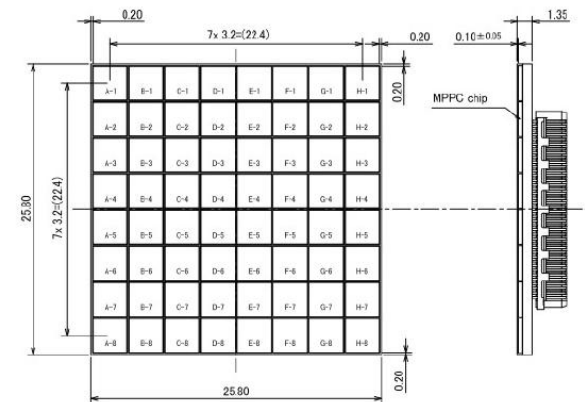


Is Hamamatsu able to produce a uniform array?

DAQ-System of the SiPM-EC (256 Channels):



William Painter, KIT



Thanks for listening.

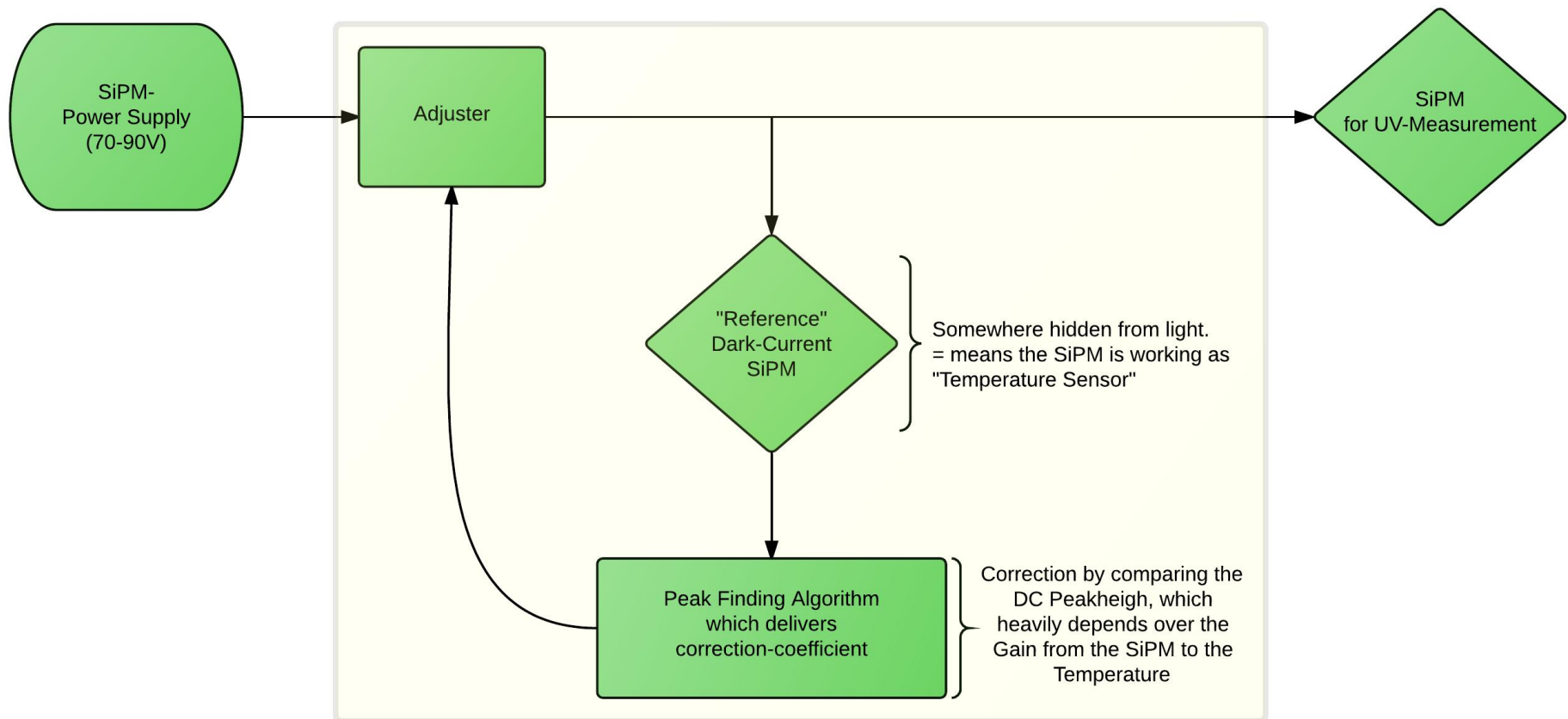
BACKUP

Possible Way to handle this?: Temperature-Correction-Coefficient α

$$\text{Gain} = \frac{C_{Pixel} \cdot \Delta V}{e}$$

$$V_{Bias} = \alpha \cdot T$$

$$\alpha \approx \frac{60 \text{ mV}}{^{\circ}\text{C}} \longrightarrow \Delta T = 2\text{K} \longrightarrow \Delta \text{Gain} = 33\% !!$$



Working principle Avalanche Photodiode

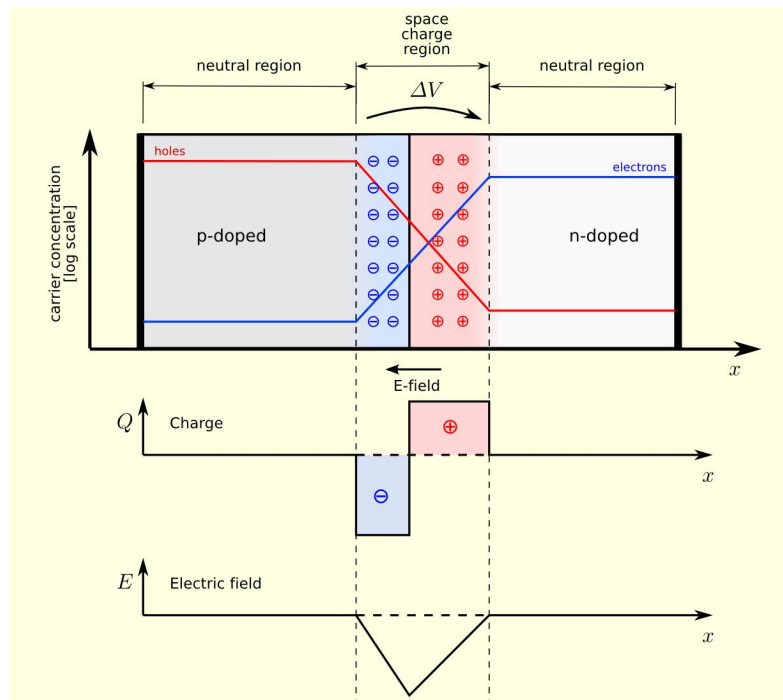
They are using:

- Photoelectric effect
- Avalanche multiplication

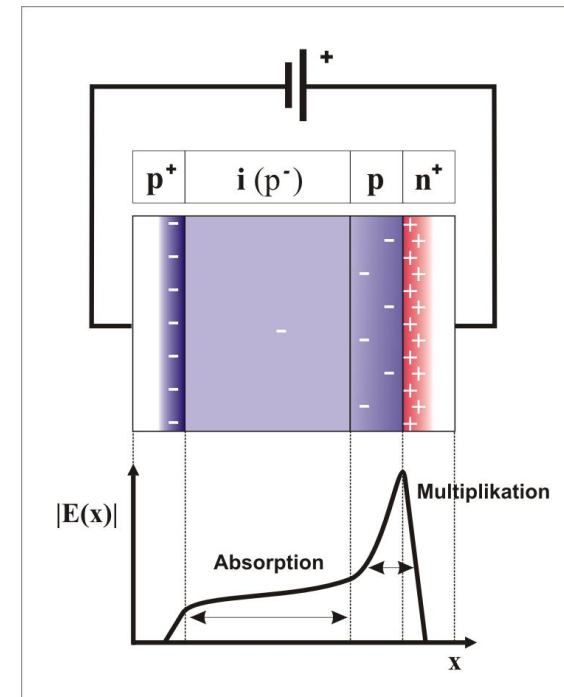
$$G = \frac{Q_{pixel}}{e} = \frac{C_{pixel} \cdot (U_{bias} - U_{breakdown})}{e}$$

$$C_{pixel} \approx 100 \text{ fF}$$

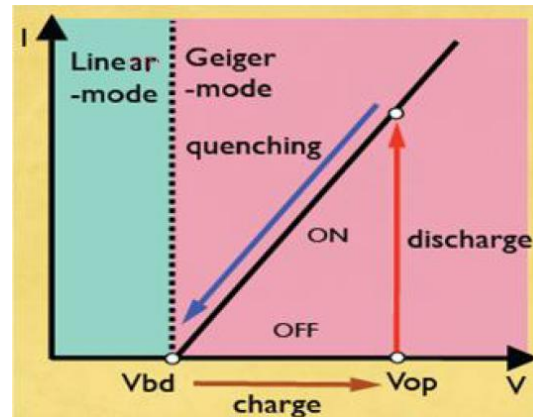
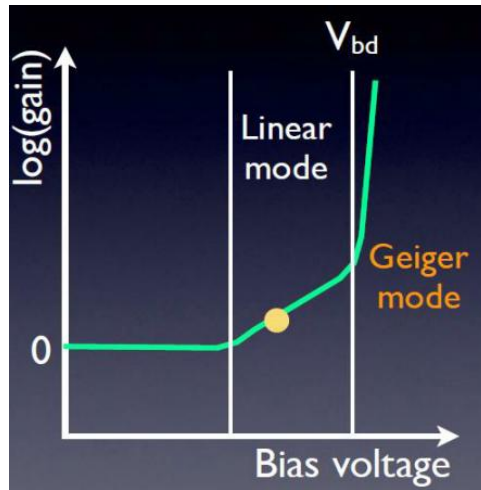
Normal diode



APD



Geiger-Mode



Darkcounts

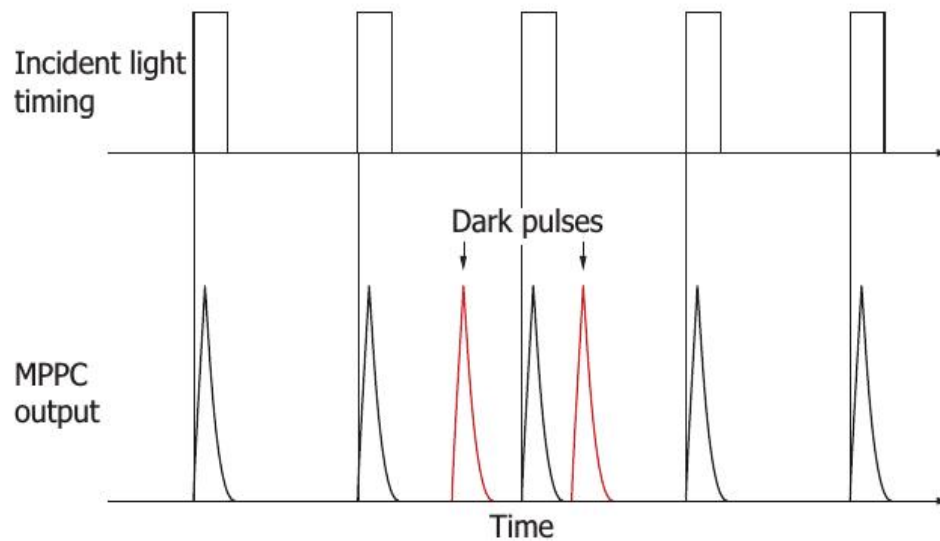
$$N_{0.5 \text{ p.e.}}(T) \approx AT^{\frac{3}{2}} \exp\left[\frac{E_g}{2kT}\right] \dots\dots\dots (4)$$

T : absolute temperature [K]

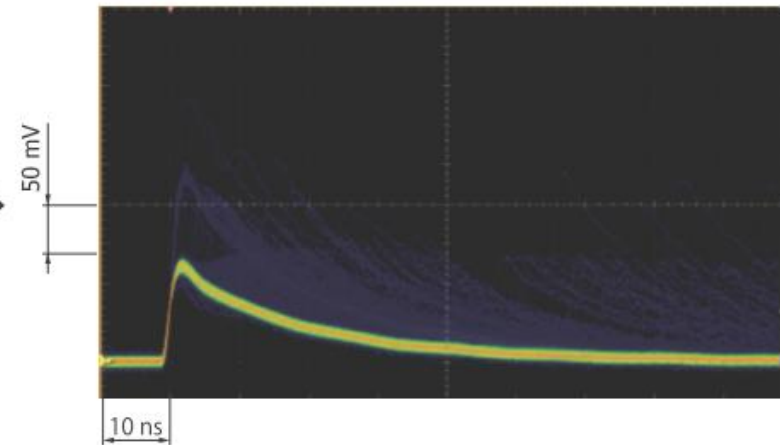
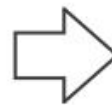
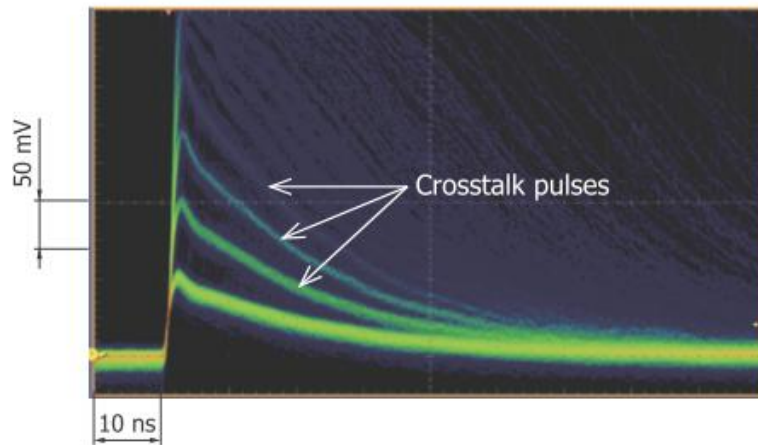
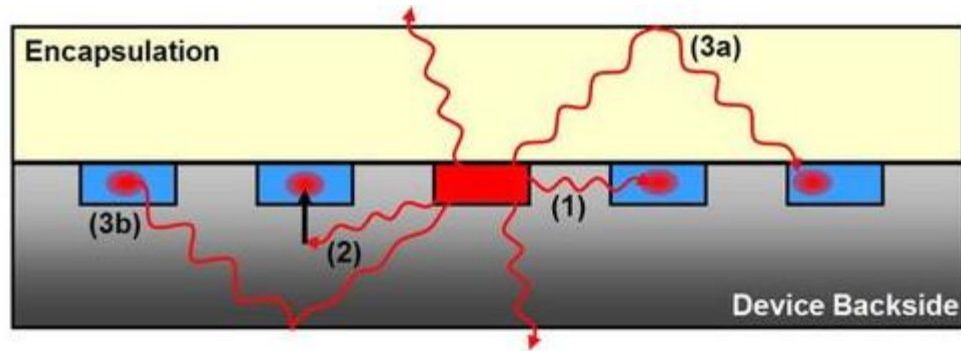
A : arbitrary constant

E_g: band gap energy [eV]

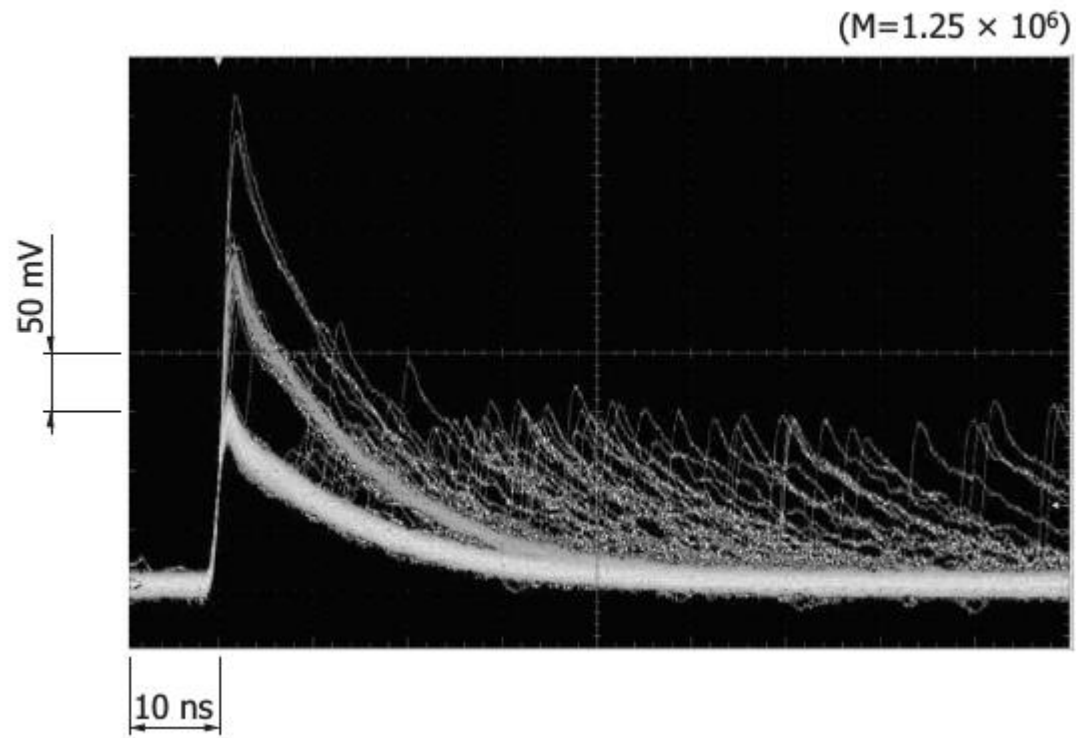
k : Boltzmann's constant [eV/K]



Crosstalk



Afterpulse



SiPM specialized peak finding algorithm

- Time Calibration via samplerate

Find increasing flank:

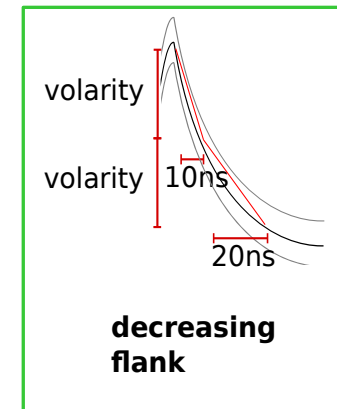
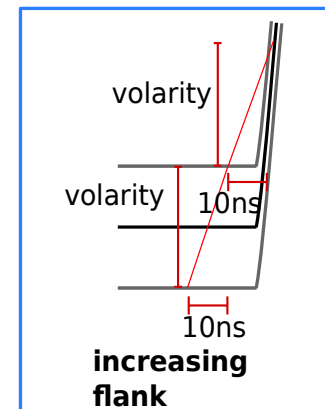
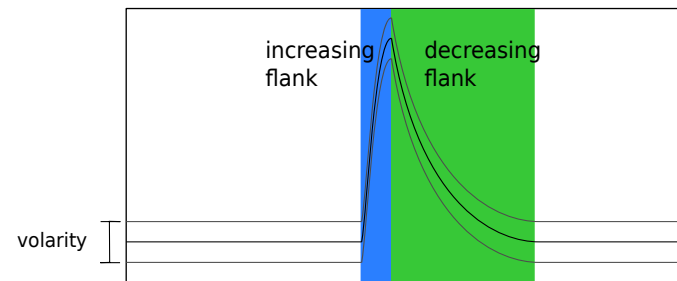
- difference between 10ns must be bigger than volarity (default 0.01, uncertainty of measurment)
- difference between 10ns must be bigger than volarity and direct afer first condition

Find decreasing flank:

- for a peak, a decreasing flank must be after an increasing flank
- between 20ns the difference must be bigger than volarity (first gradient is higher than second because of exponential decrease)
- direct after 2. condition: between 10ns the difference must be bigger than volarity
- no other decreasing flank found since a increasing flank (no double counting)

```

tom@tom:~/code/peakfinding$ ./amp_peak_counter -h
help:
  --file          inputfile
  --threshold     set threshold in (V)
  -o             --output   set file for maximum value of peak (default: output.txt)
  --volarity      set volarity in (V)
  --10GHz        set sample rate to 10GHz (default)
  --1GHz         set sample rate to 1GHz
  
```



DRS4 Data Acquisition: „get_data“

```

tom@tom-Vostro-V131:~$ get_data -h
get_data - DRS4 acquisition Rev. 1393

Usage: get_data [OPTIONS]

Record settings
-c CH1[,CH2,...] Set one or more readout channel numbers (default = 1)
-n NUM_FRAMES  Number of frames to record
-a             Free-running mode (no trigger)
-T [CH_NUM|ext] Trigger on channel CH_NUM or 'ext' for external trigger
-t TrigTrheshV Set the trigger threshold in Volts. Default = -0.05V
-P           Trigger on positive edge [default NEGATIVE]
-D delay      Trigger Delay in percent
-F f_SAMPLE   Sampling frequency in GSp/s, range ~0.68-5, default 0.68GSp/s
-p           Short for '-r 0.5', sets the input range to [0 .. 1V]
-r CenterVolt Set the center input voltage. Quantisation is done in
              the CenterVolt-0.5 .. CenterVolt+0.5 voltage range.
              Default = 0

Data Format
-f FORMAT      Set the format of the recorded data.
              FORMAT is one of MULTIFILE, MULTIFILE_BIN, TEXT, BIN, YAML or ROOT.
-d            Output directory for MULTIFILE output (will create one file per frame!)
-o            Name of the output file(s). The correct file extension will be
              appended automatically, so there is no need to specify it. If the
              wrong extension is specified, the correct one is appended, too!
-C            Enable zlib compression (only works with single text file).
-l LVL        Set compression level (default 9). Only used if -c is set
-H user_header Add a line to the user header

Temperature Control
-U socket      UNIX domain socket for detector control.
              Default: /tmp/detector_control.unix
-s T_soll      Enable temperature stabilized measurement.
              Value in Kelvin if suffixed by K, other wise
              it is interpreted as degree Celsius

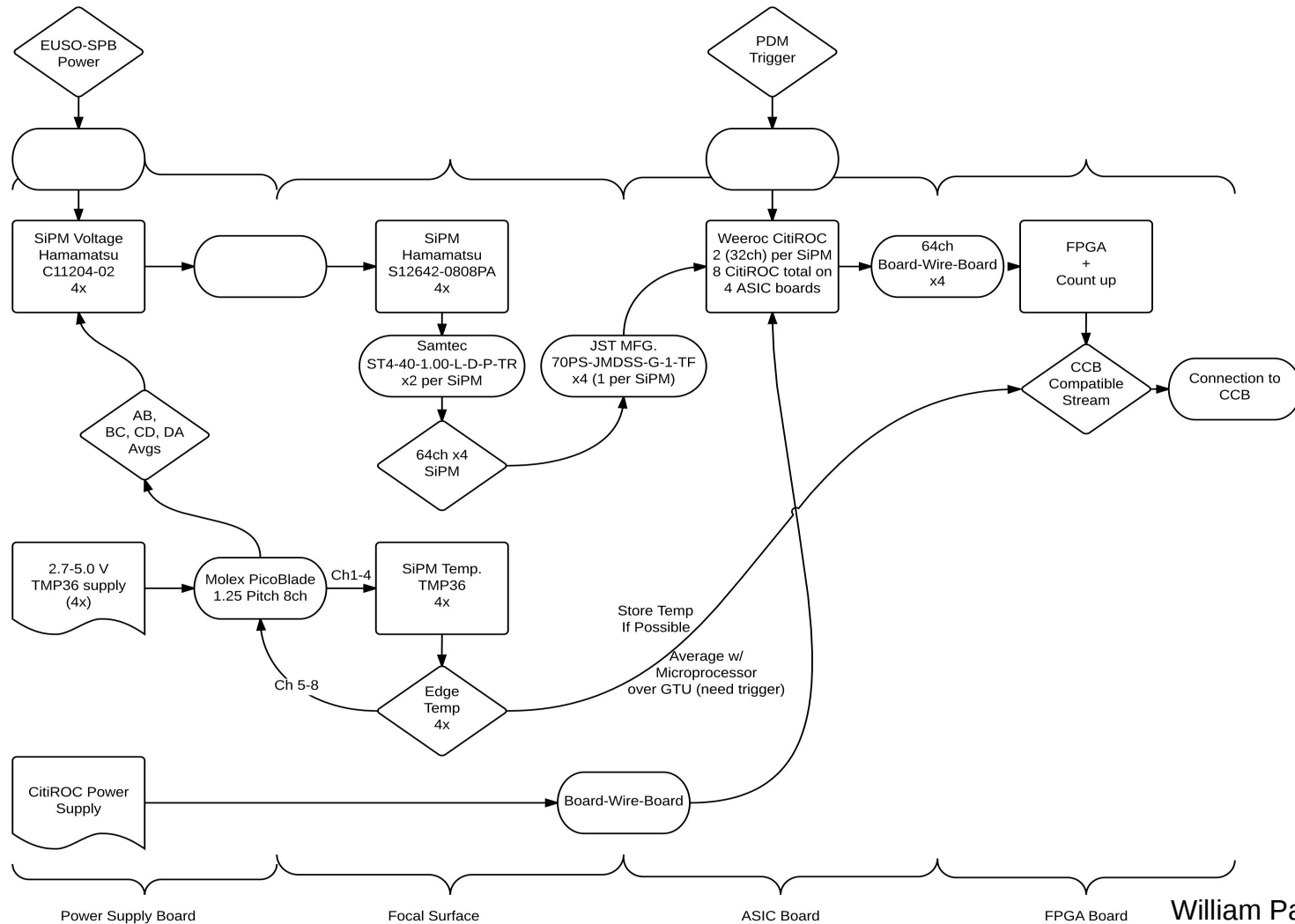
Miscellaneous Options
-v            Show version information
-X            Stand-alone, embedded mode.
              Outputs certain information in a machine-readable JSON. Each piece
              of information is written line by line to stdout in JSON format; errors
              are written to stderr in verbose format.
              The advantage of this output mode is simple processing,
              e.g. for GUI embedding with progrssbars, while maintaining
              full flexibility and extensibility.
  
```

- Very creative name :-D
- Alpha Status
 - Working, but a lot of work to do!
 - Already a lot of functions (see below)
- Functions:
 - adjustable trigger threshold
 - Multiple Channel Readout at once
 - Important to do: boolean algebra
- Binary Output (buggy) and also ROOT Format (working)
- Zlib compression
- First attempts to implement temperature stabilized data acquisition

Licensing:
 get_data is distributed under the GPL3 license

FindROOT.make by F.Uhlig@gsi.de <http://fairroot.gsi.de>
 FindLibUSB.cmake by Michal Cihar, michal@cihar.com

DAQ-System of the SiPM-EC (256 Channels)



William Painter, KIT

The near future of JEM-EUSO

2014

SPB-EUSO

2017

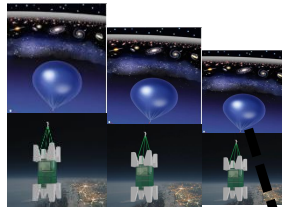


2024

EUSO



TA-EUSO



EUSO-Balloon

JEM-EUSO



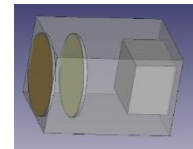
2020

2024

AO from NASA in 2014?

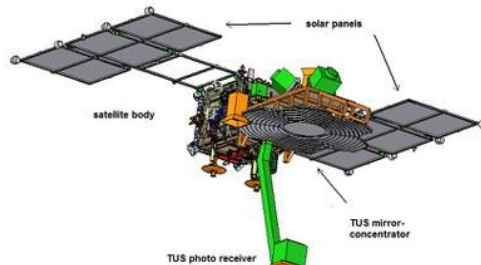
SiPM-PDM?

M4 mission
Free flyer?

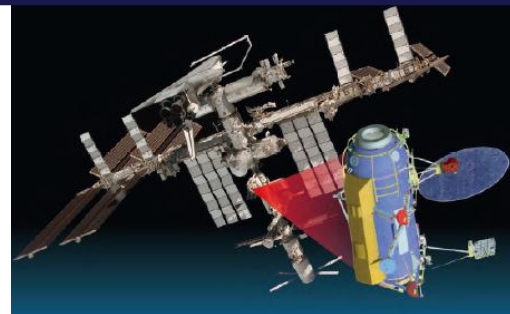


Mini-EUSO

Collaboration with a Russian project, KLYPVE

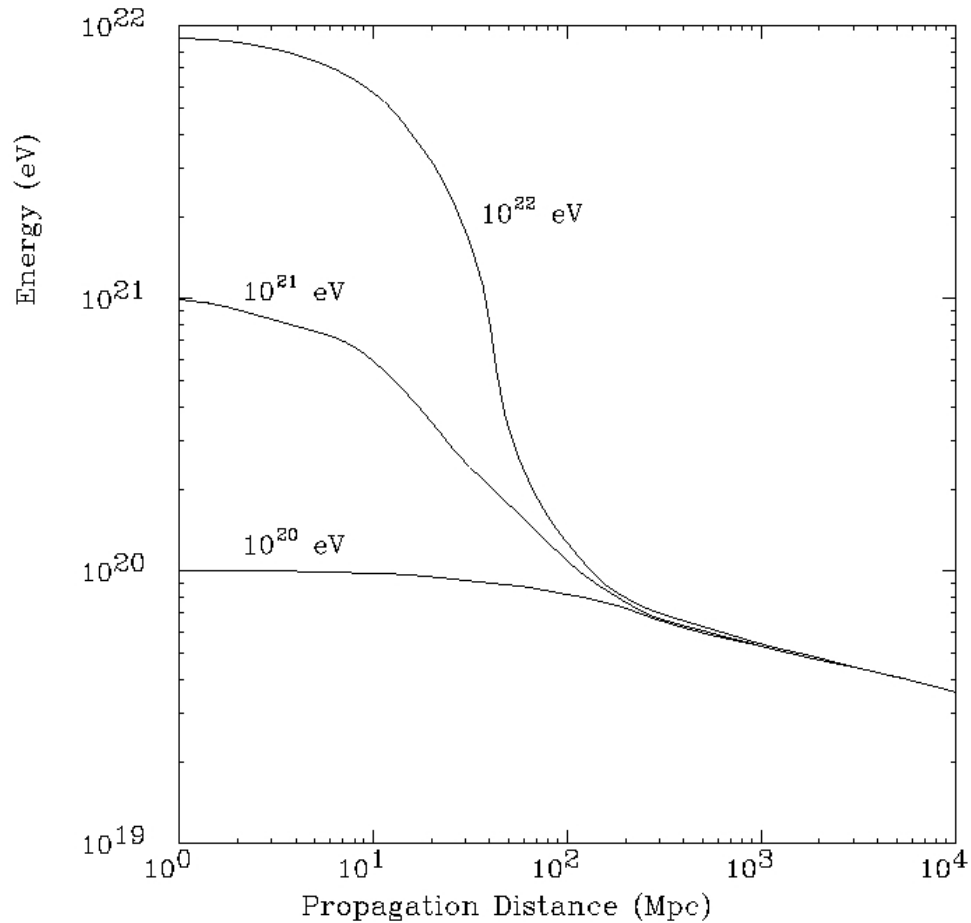


TUS



“K-EUSO”

GZK-Cutoff



PDE from PMTs and SiPMs

SiPM

$$PDE(\lambda, V) = QE(\lambda) \cdot \varepsilon(V) \cdot GE$$

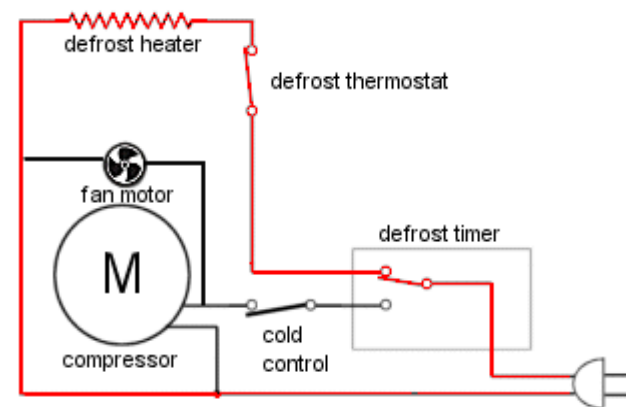
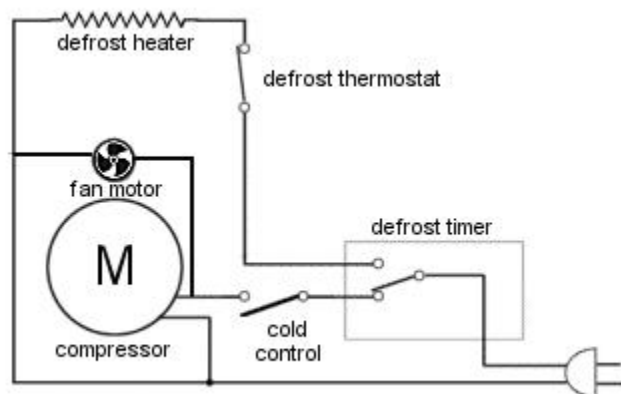
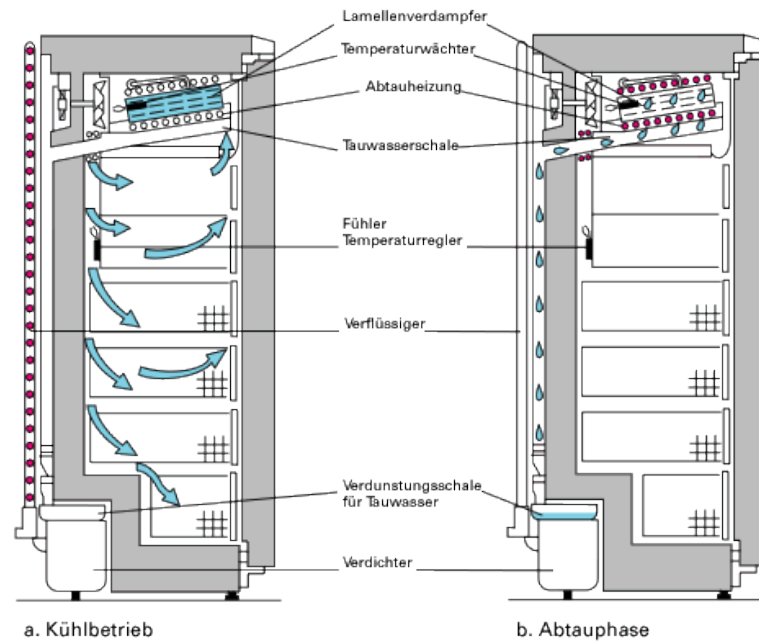
<i>QE</i>	Quantum Efficiency
ε	Geiger Efficiency
<i>GE</i>	Geometrical Efficiency
λ	Wavelength
<i>V</i>	Bias

PMT

$$PDE = QE(\lambda) \cdot CE$$

<i>QE</i>	Quantum Efficiency
<i>CE</i>	Collection Efficiency

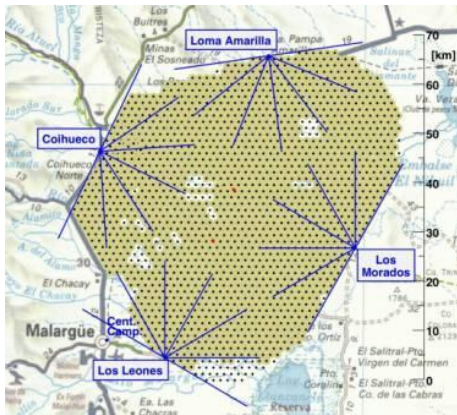
Defrost System



JEM-EUSO Performance

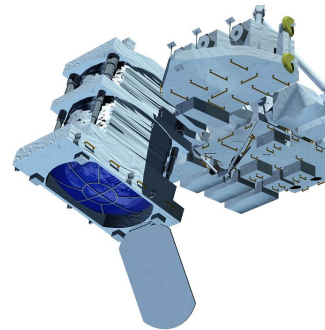
How many UHECR with $E > 6 \cdot 10^{19} eV$?

AUGER, Malargüe (3.000 km²)



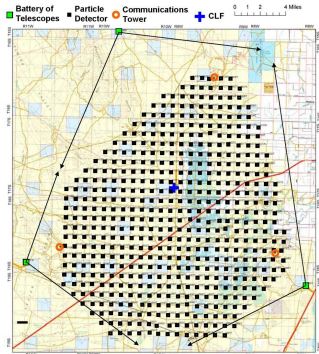
~ 20 Events
each Year

JEM-EUSO, in Space (140.000 km²)



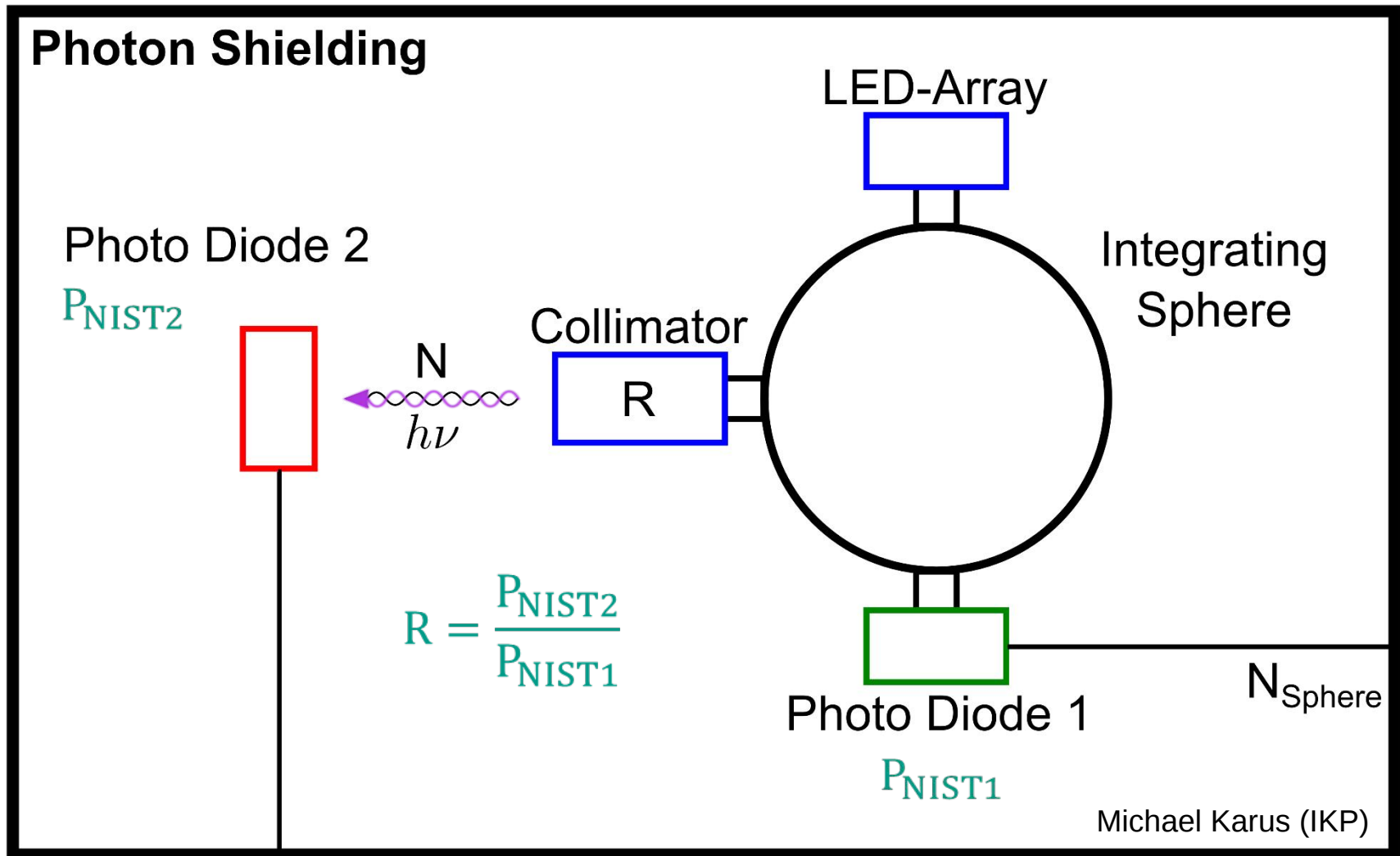
~ 200 Events
each Year

TA, Utah (700 km²)



~ 5 Events
each Year

Calibration set-up



Calibration set-up

