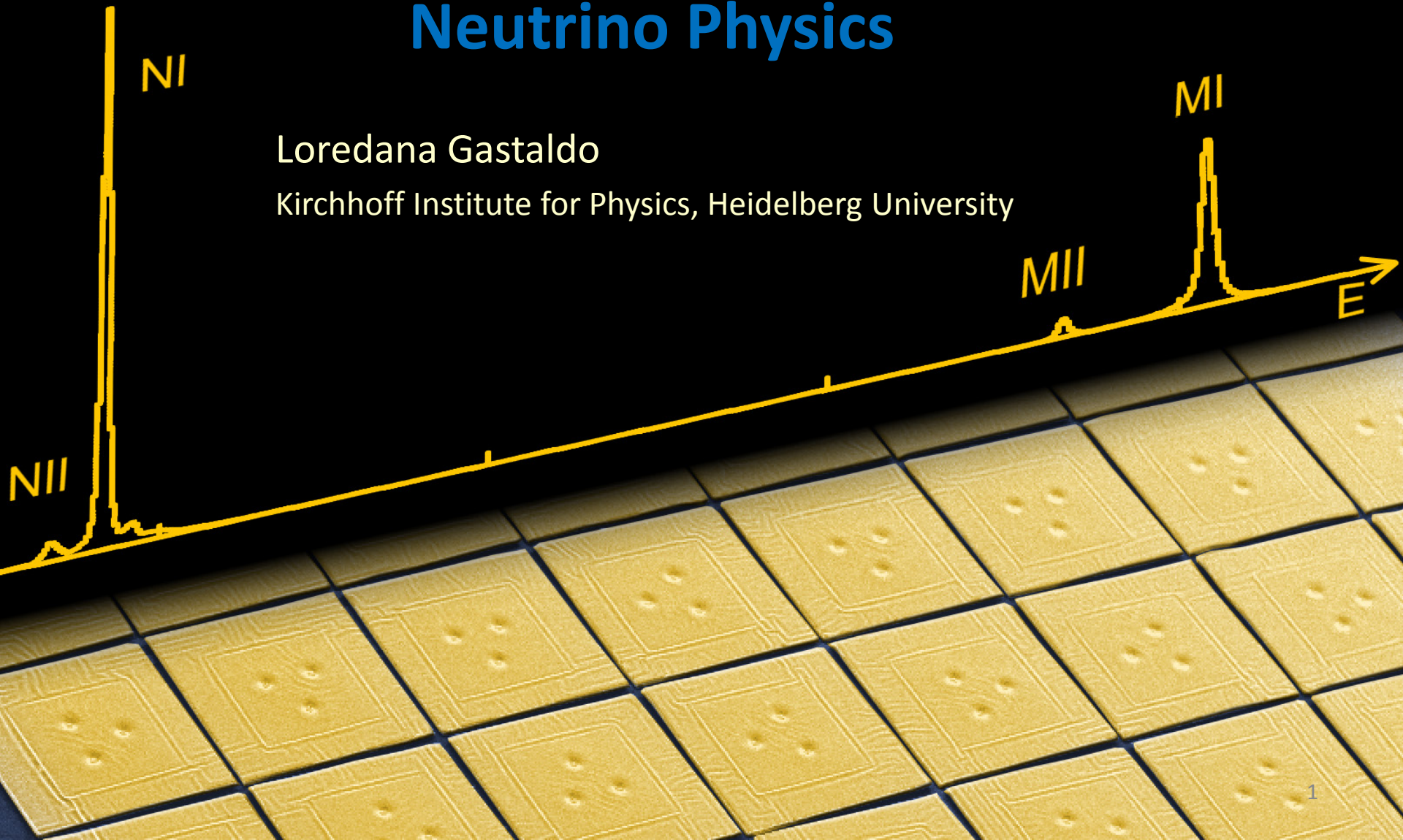


Magnetic Calorimeters for Neutrino Physics

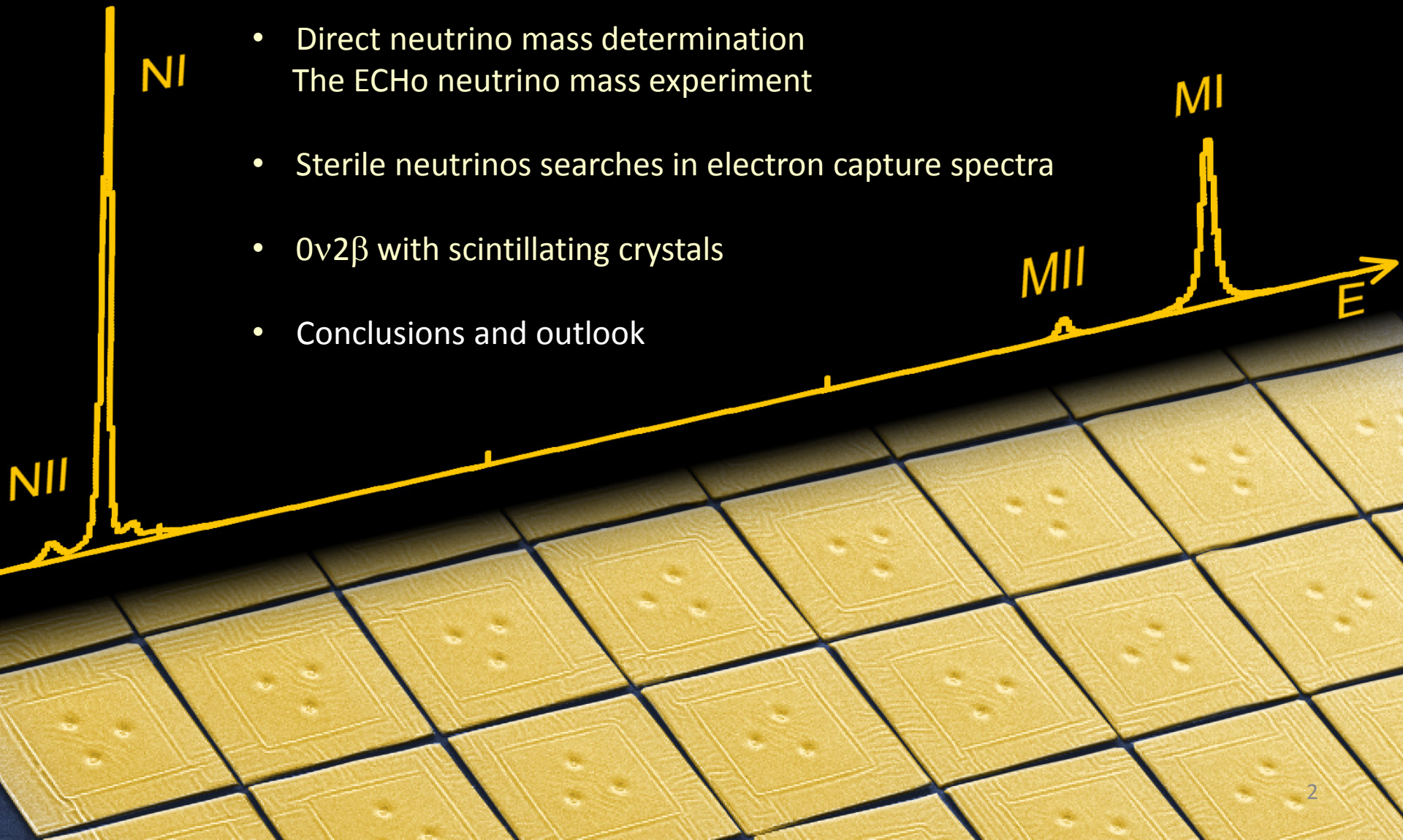
Loredana Gastaldo

Kirchhoff Institute for Physics, Heidelberg University



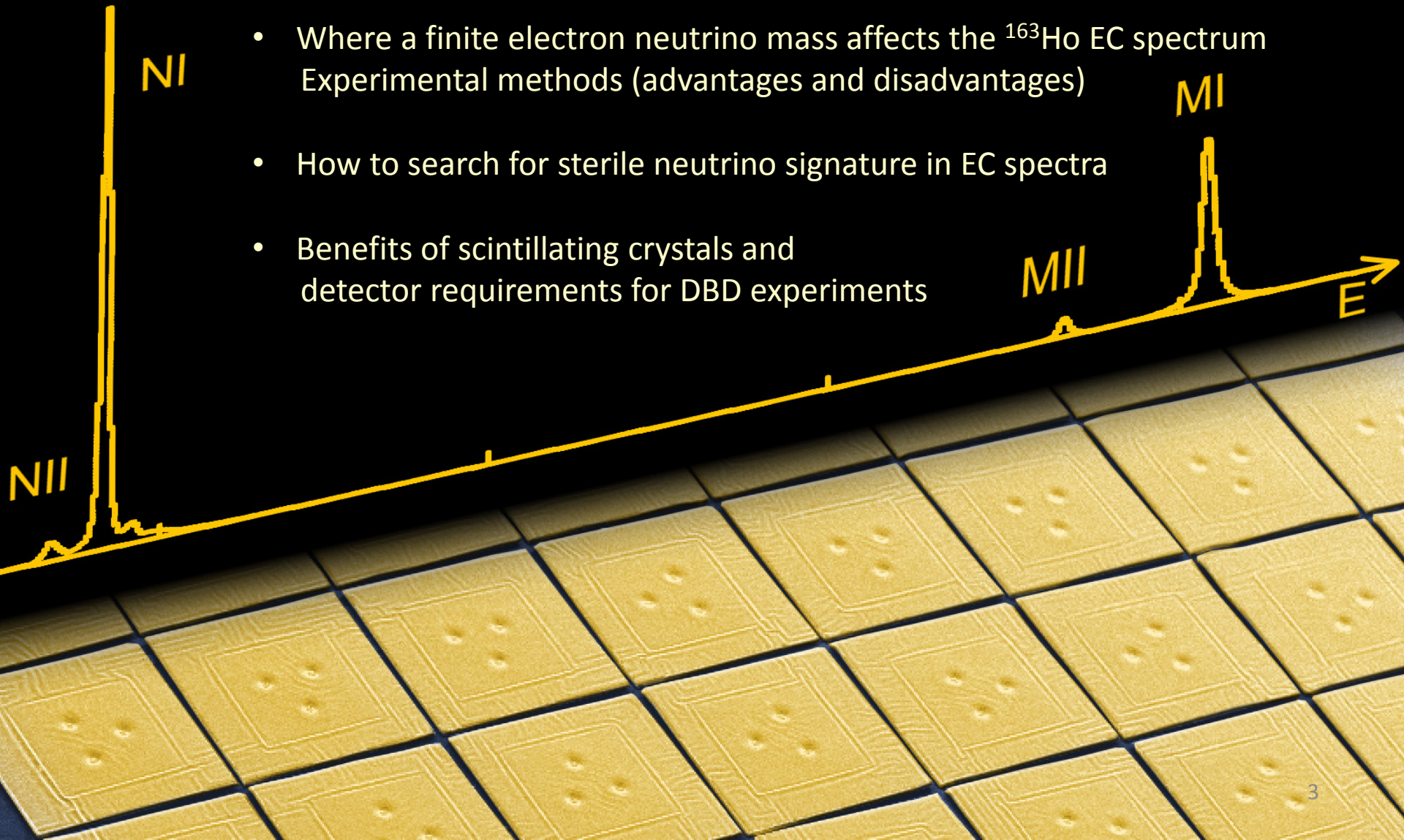
Contents

- **Metallic Magnetic Calorimeters**
- Direct neutrino mass determination
The ECHO neutrino mass experiment
- Sterile neutrinos searches in electron capture spectra
- $0\nu 2\beta$ with scintillating crystals
- Conclusions and outlook

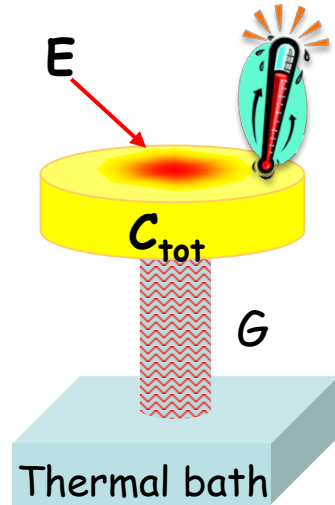


Take-home messages

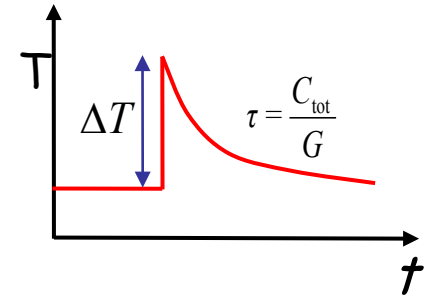
- Working principle of MMCs and their performance
- Where a finite electron neutrino mass affects the ^{163}Ho EC spectrum
Experimental methods (advantages and disadvantages)
- How to search for sterile neutrino signature in EC spectra
- Benefits of scintillating crystals and detector requirements for DBD experiments



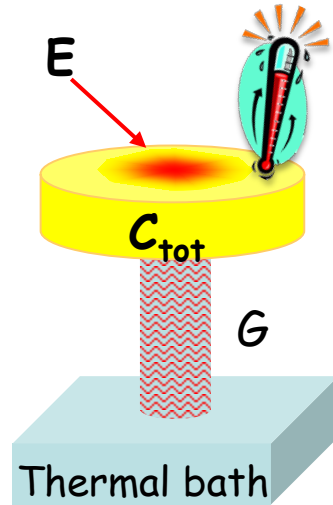
Low temperature micro-calorimeters



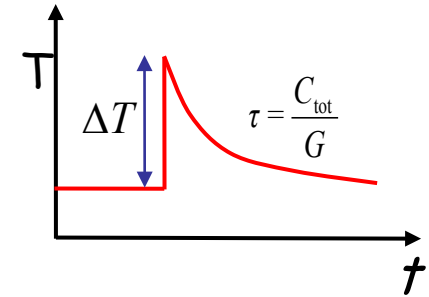
$$\Delta T \cong \frac{E}{C_{\text{tot}}}$$



Low temperature micro-calorimeters

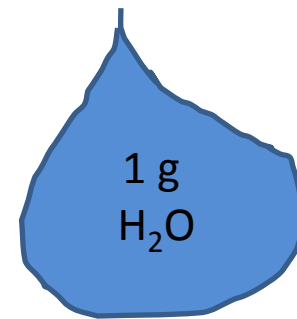


$$\Delta T \cong \frac{E}{C_{\text{tot}}}$$

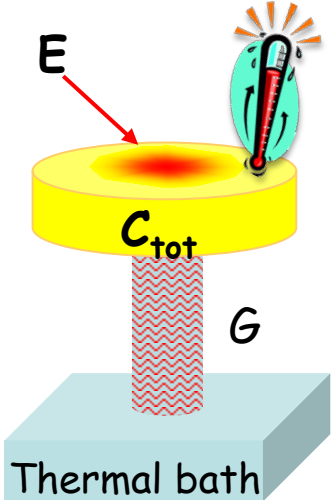


$E = 10 \text{ keV}$
 $C_{\text{tot}} = 4.18 \text{ J/K}$

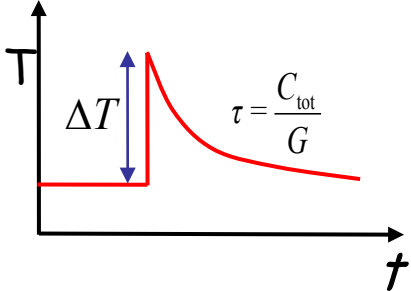
$\rightarrow \sim 3.8 \cdot 10^{-16} \text{ K}$



Low temperature micro-calorimeters



$$\Delta T \cong \frac{E}{C_{\text{tot}}}$$

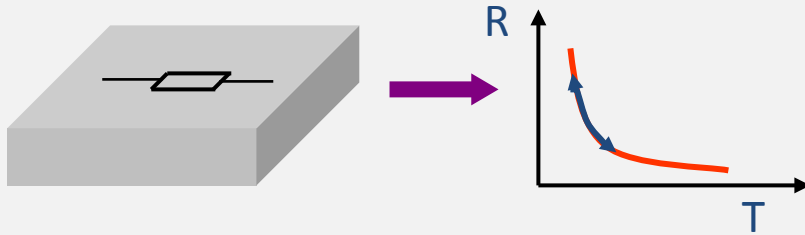


$E = 10 \text{ keV}$
 $C_{\text{tot}} = 1 \text{ pJ/K}$ } $\rightarrow \sim 1 \text{ mK}$

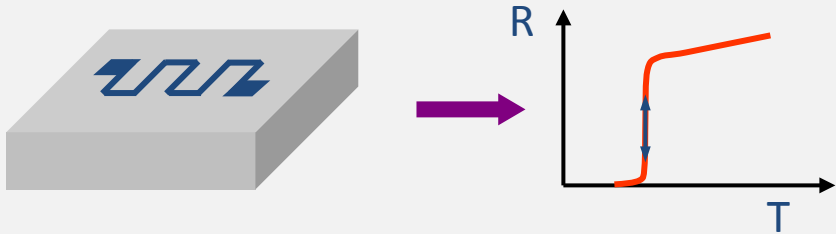
- Very small volume
- Working temperature below 100 mK
small specific heat
small thermal noise
- **Very sensitive temperature sensor**

Temperature sensors

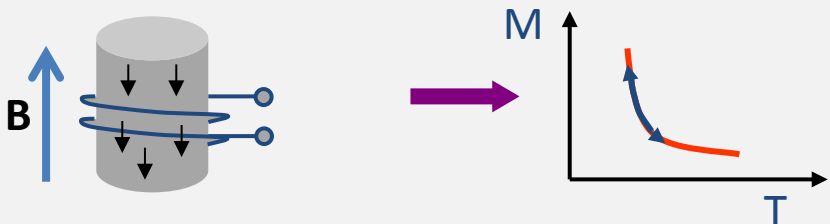
Resistance of highly doped semiconductors, NTD-Ge



Resistance at superconducting transition, TES

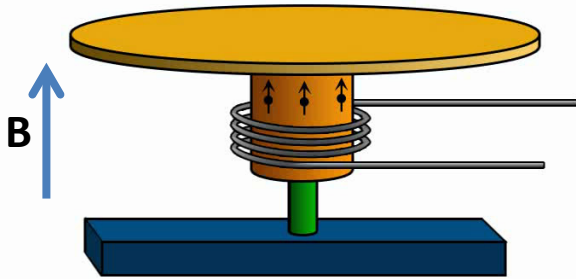


Magnetization of paramagnetic material, MMC

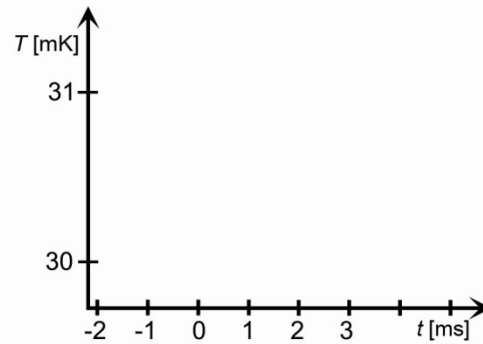


Metallic Magnetic Calorimeters - MMC

Paramagnetic sensor: Au:Er
Ag:Er

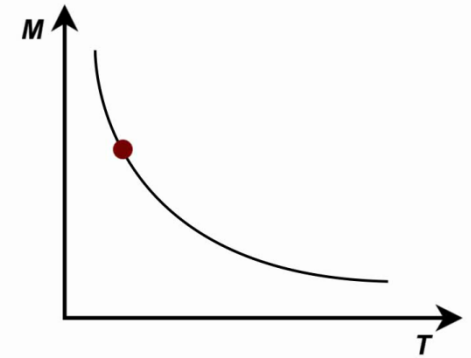


Sensor temperature



Pulse: 6 keV photon in
 $C_{abs} \sim 1$ pJ/K

Sensor magnetization



$$\delta M = \frac{\partial M}{\partial T} \delta T = \frac{\partial M}{\partial T} \frac{E_\gamma}{C_{tot}}$$

main differences to calorimeters with resistive thermometers

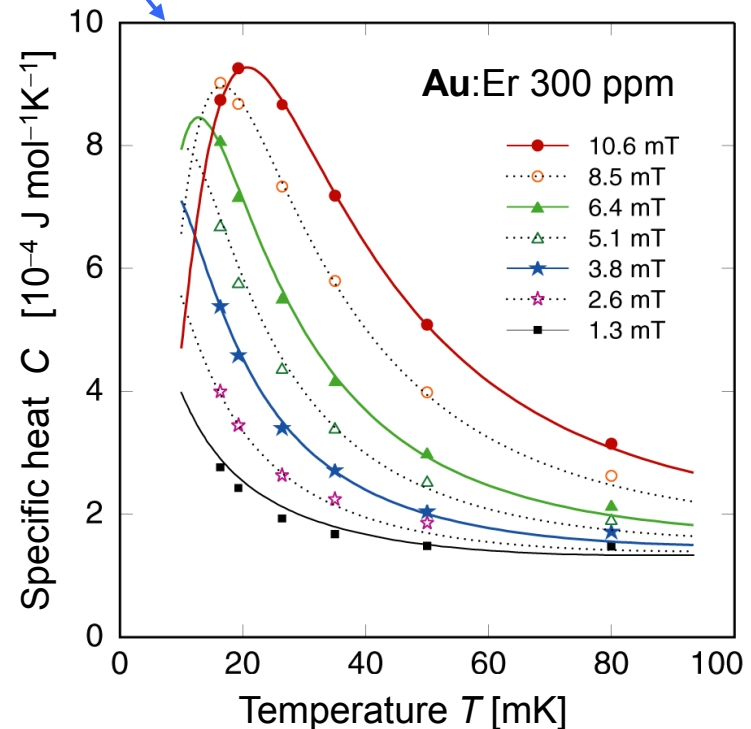
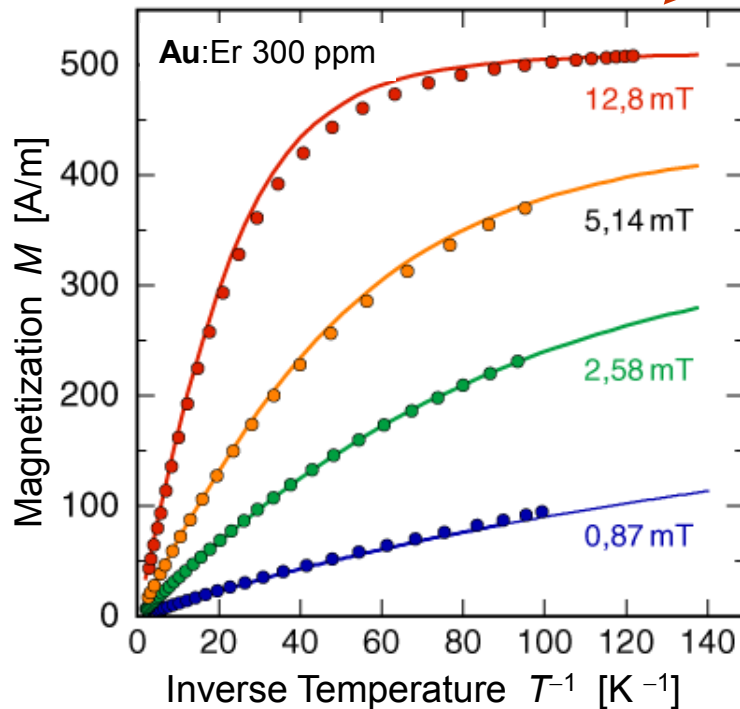
no dissipation in the sensor

no galvanic contact to the sensor

MMC: signal size

Numerical calculations based on mean field approximation are used to describe thermodynamical properties of interacting spins (RKKY)

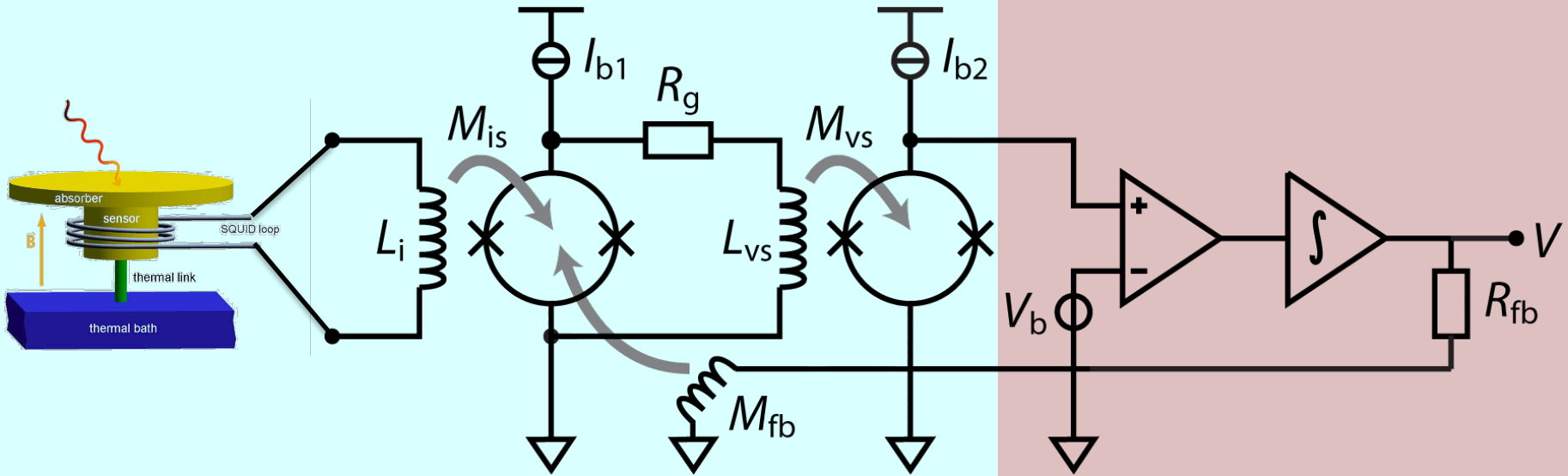
$$\delta\Phi_S \propto \frac{\partial M}{\partial T} \frac{1}{C_{\text{tot}}} \delta E$$



MMCs: Readout

$T \sim 30 \text{ mK}$

$T \sim 300 \text{ mK}$

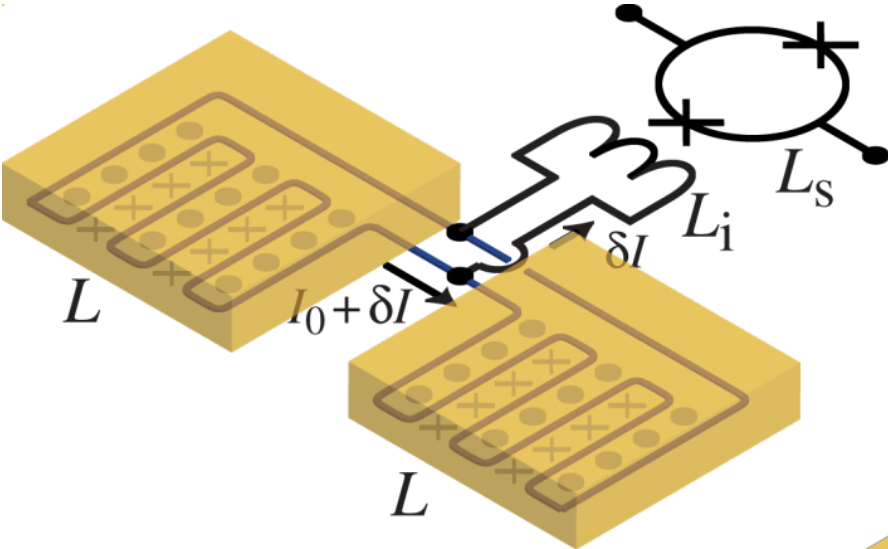


Two-stage SQUID setup with flux locked loop allows for:

- low noise
- large bandwidth / slewrate
- small power dissipation on detector SQUID chip (voltage bias)

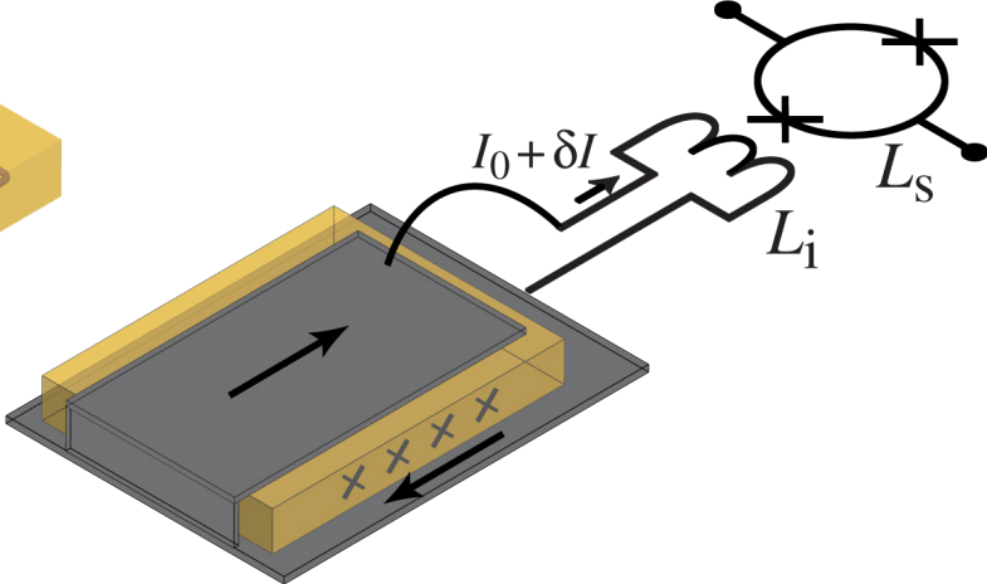
MMCs: Planar Geometries

- Planar temperature sensor
- B-field generated by persistent current
- transformer coupled to SQUID

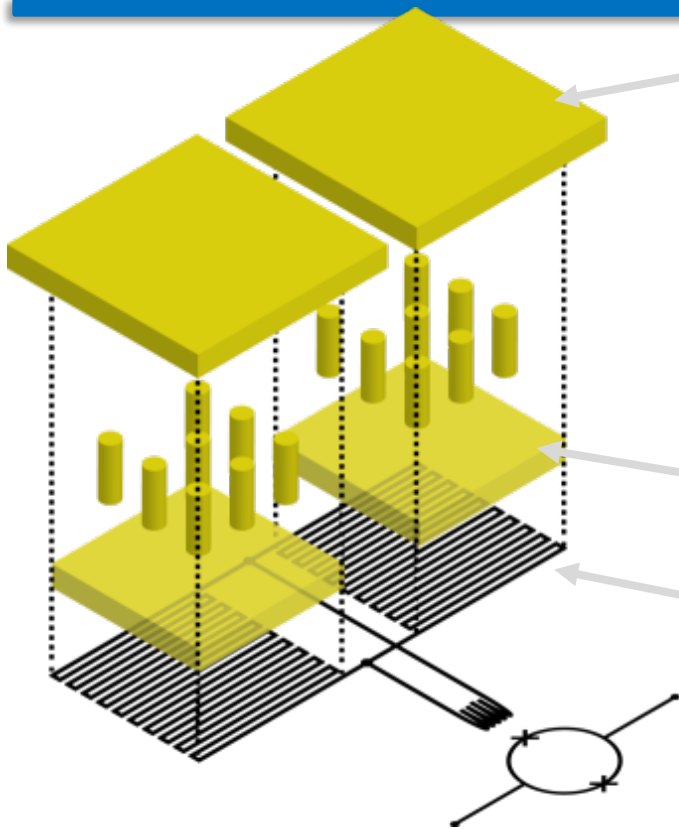


Meander-shaped pick-up coil:
Filling factor ~ 0.5

Sandwich design:
Filling factor ~ 1



maXs20: 1d-array for soft x-rays



- **1×8 x-ray absorbers**

- 250μm×250μm gold, 5μm thick
- 98% Qu.-Eff. @ 6 keV
- electroplated into photoresist mold (RRR>15)
- mech/therm contact to sensor by stems to prevent loss of initially hot phonons

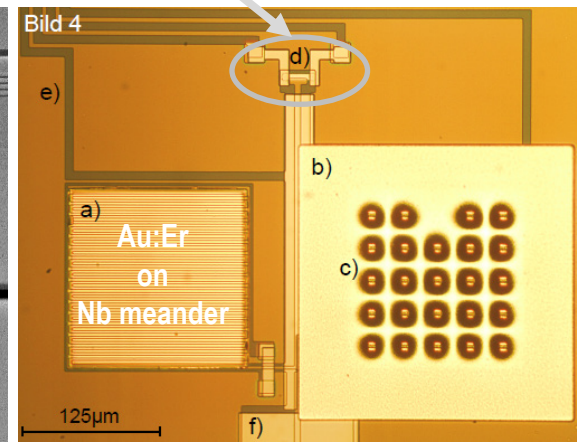
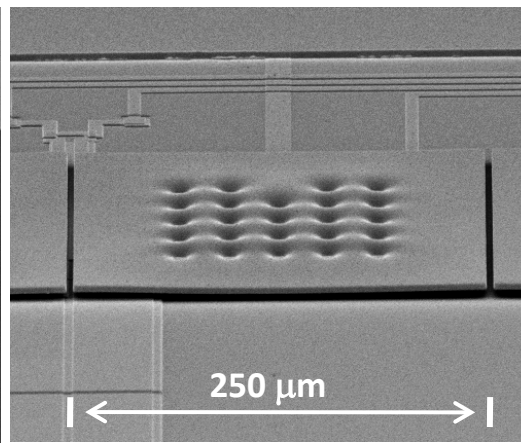
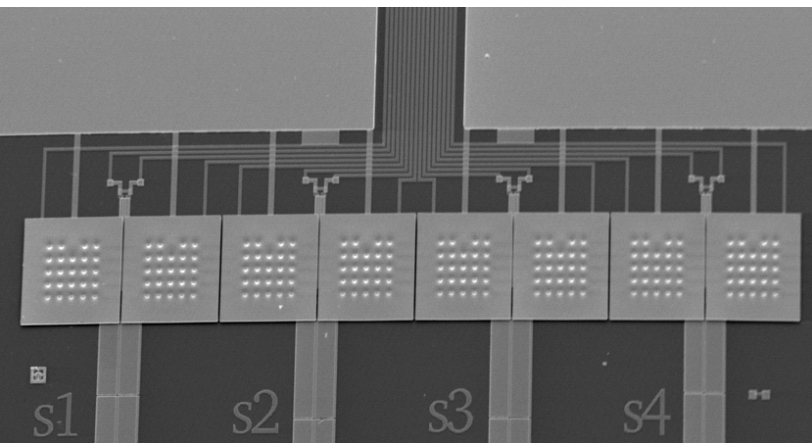
- **Au:¹⁶⁸Er_{300ppm} temperature sensors**

- co-sputtered from pure Au and high conc. AuEr target

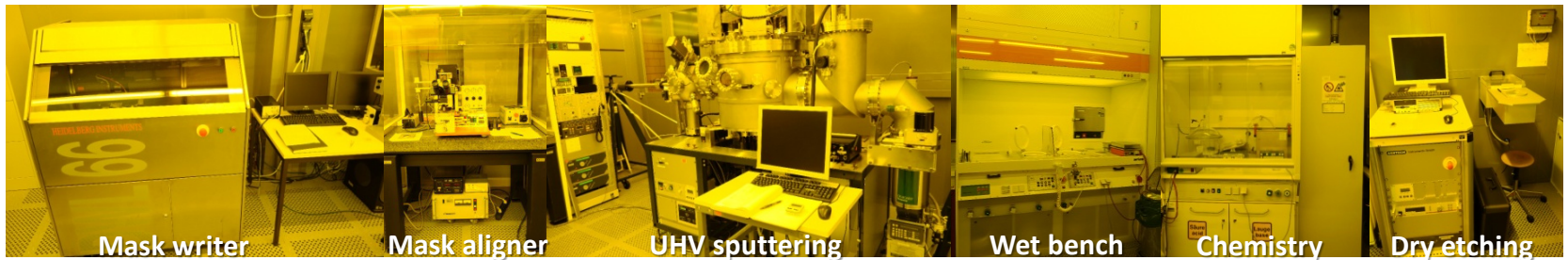
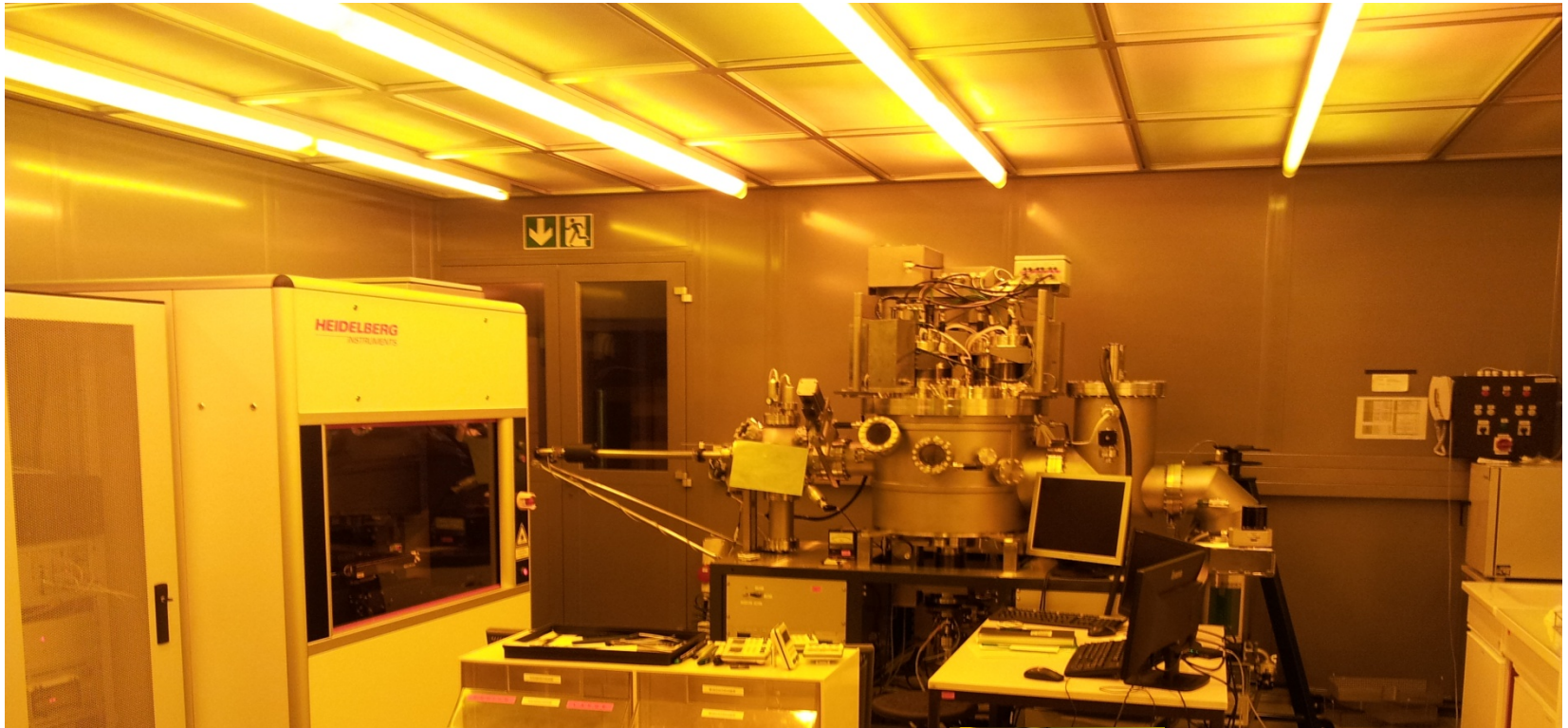
- **Meander shaped pickup coils**

- 2.5 μm wide Nb lines
- $I_c \approx 100\text{mA}$

- **On-chip persistent current switch (AuPd)**



MMCs: Microfabrication



Mask writer

Mask aligner

UHV sputtering

Wet bench

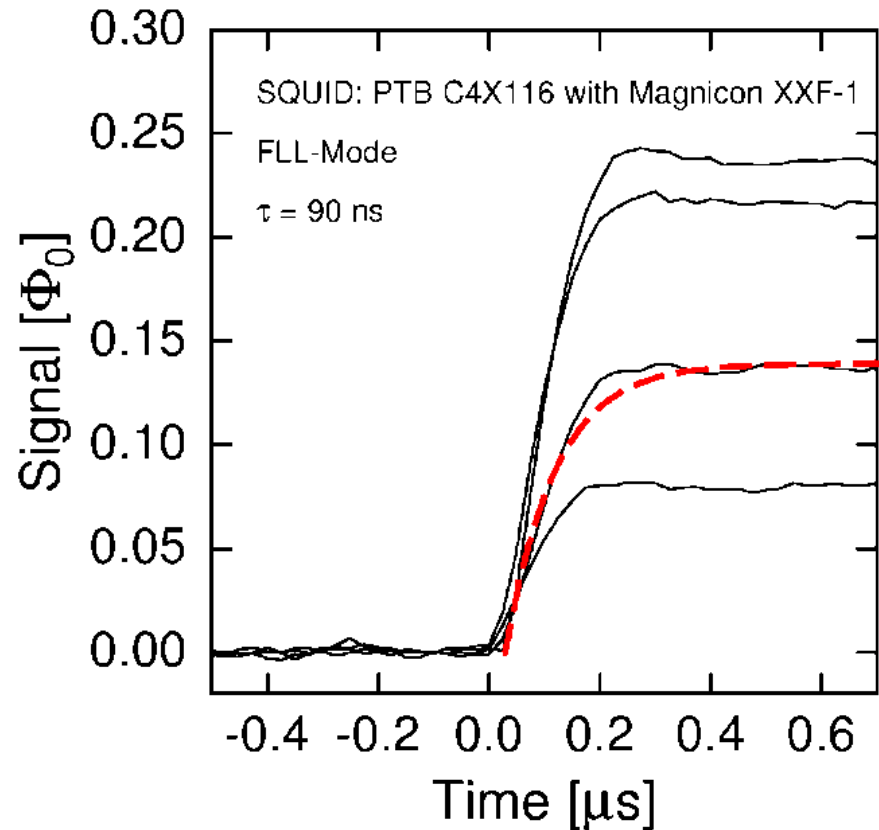
Chemistry

Dry etching

maXs20: 1d-array for soft x-rays

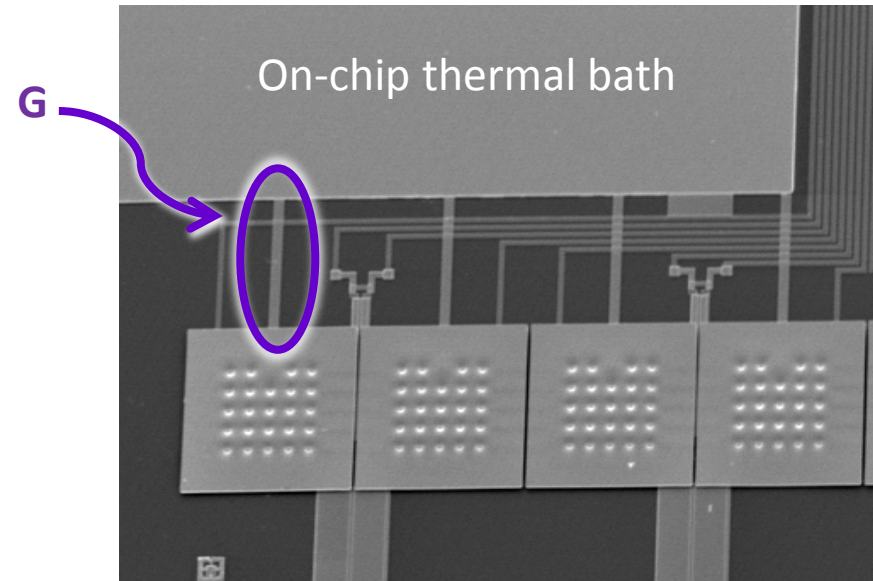
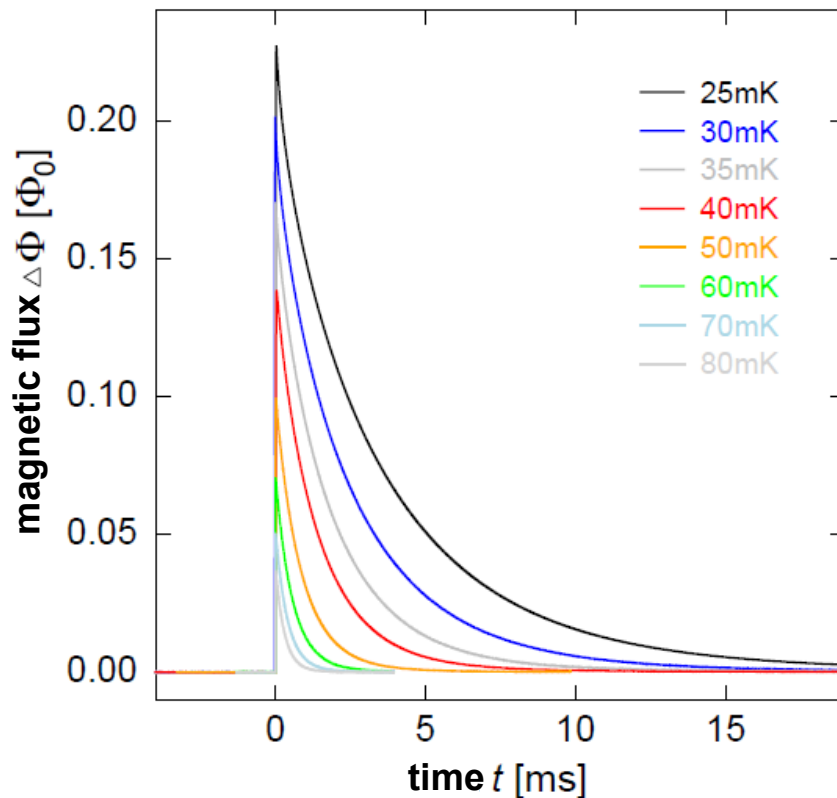
- **rise time: 90 ns @ 30 mK,**
as expected for the **spin-electron-relaxation**
from Korringa-constant of Er in Au

Fastest rise-time among
 μ -cal for x-ray detection



maXs20: 1d-array for soft x-rays

- **decay time** here: **3 ms @ 30 mK**
- nearly single exponential decay

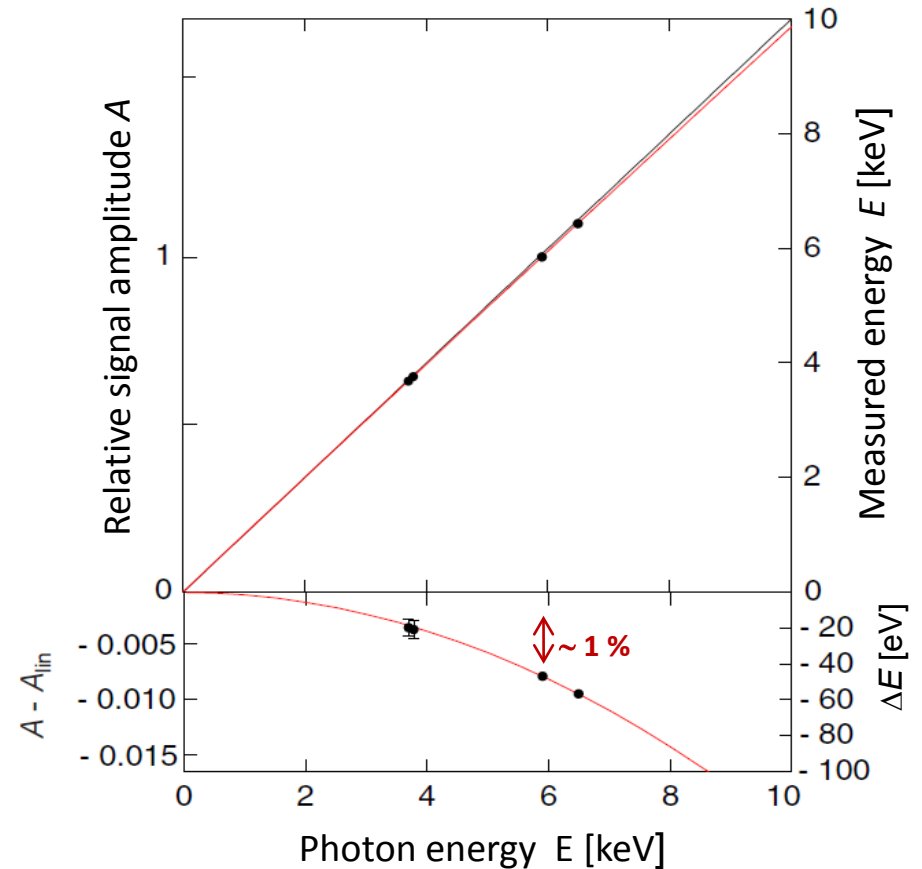


adjusted by sputtered thermal link (Au)

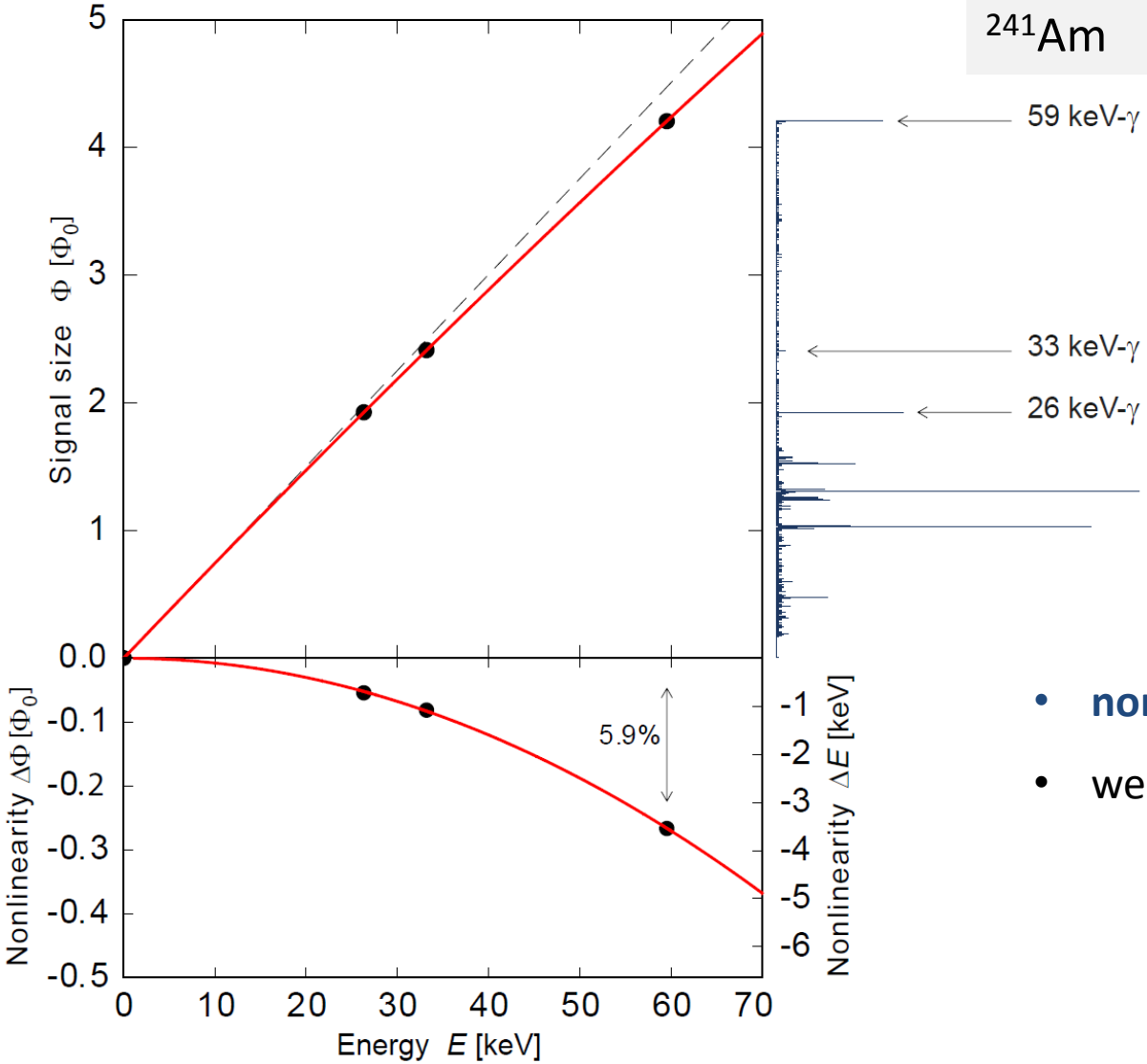
maXs20: 1d-array for soft x-rays

- **non-linearity: 1% at 6 keV**
- well described by quadratic term

The energy scale is defined with high precision



maXs20: 1d-array for soft x-rays

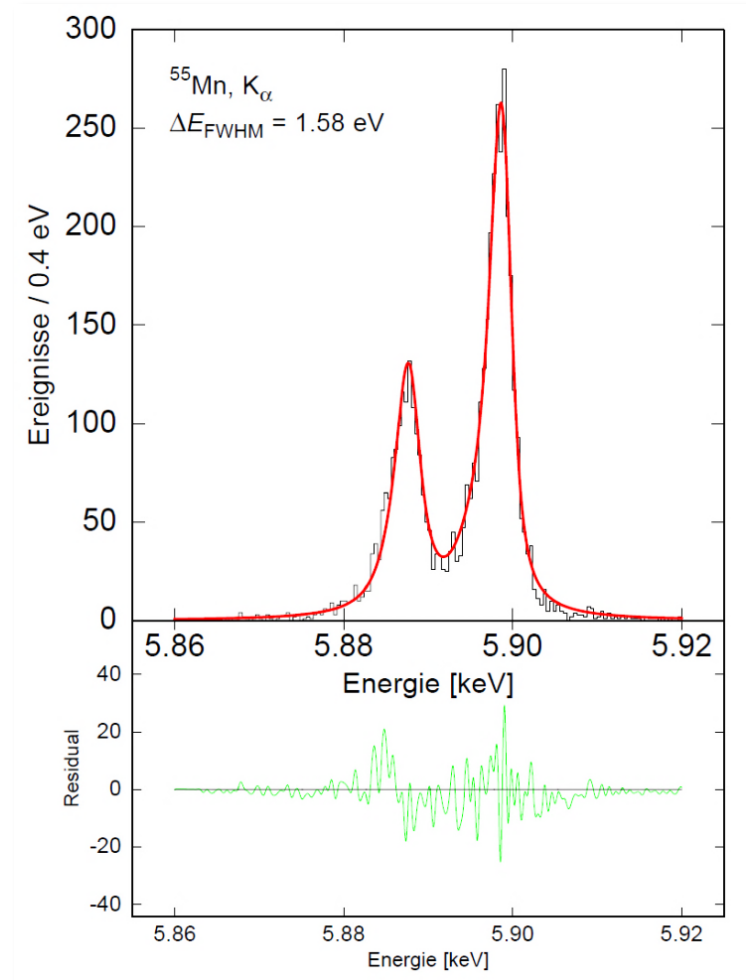


- **non-linearity: 6% at 60 keV**
- well described by quadratic term

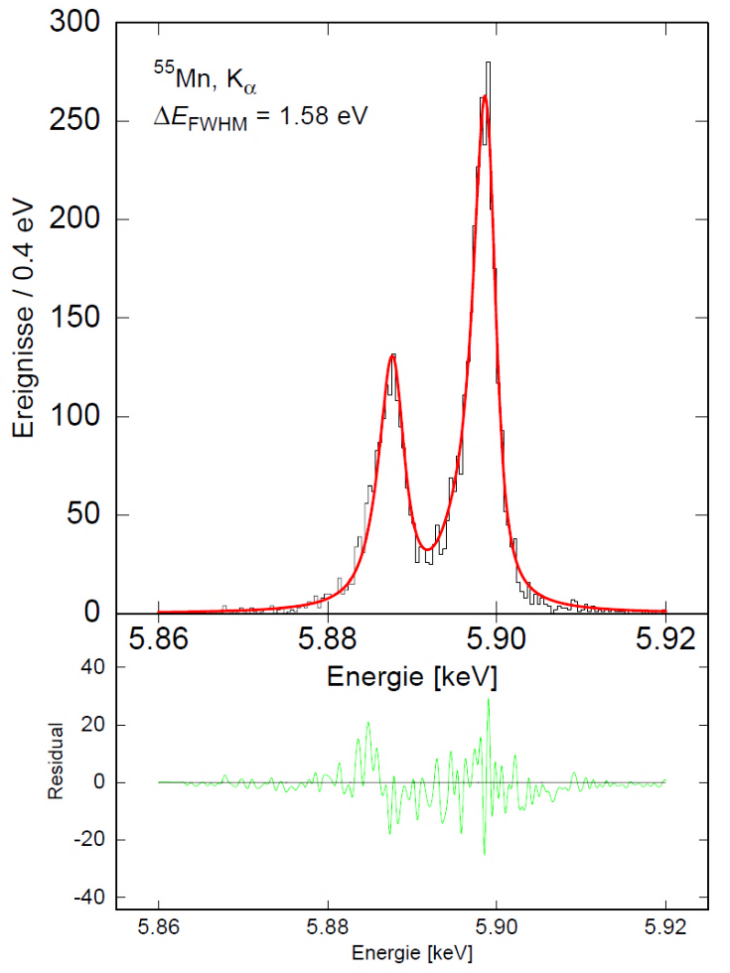
maXs20: 1d-array for soft x-rays ($T=20$ mK)

- Very good energy resolution

$$\Delta E_{\text{FWHM}} = 1.6 \text{ eV @ } 6 \text{ keV}$$

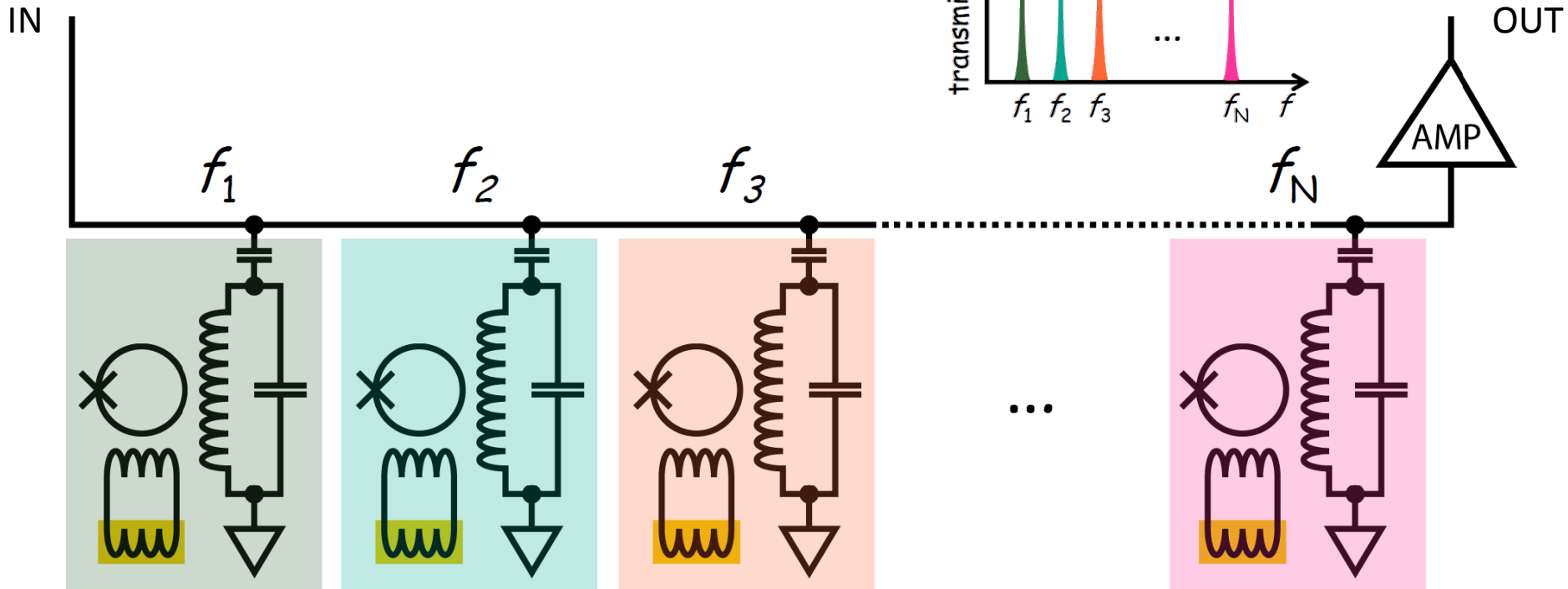
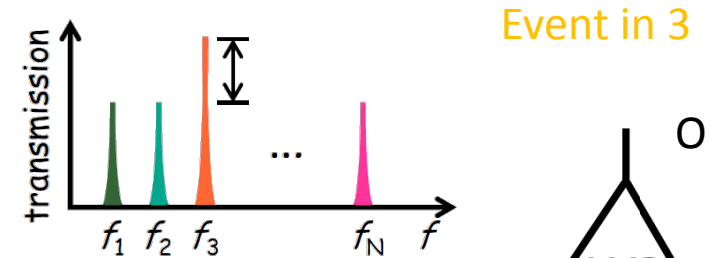
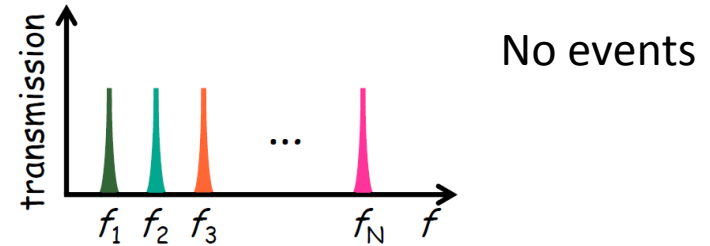


maXs20: 1d-array for soft x-rays ($T=20$ mK)

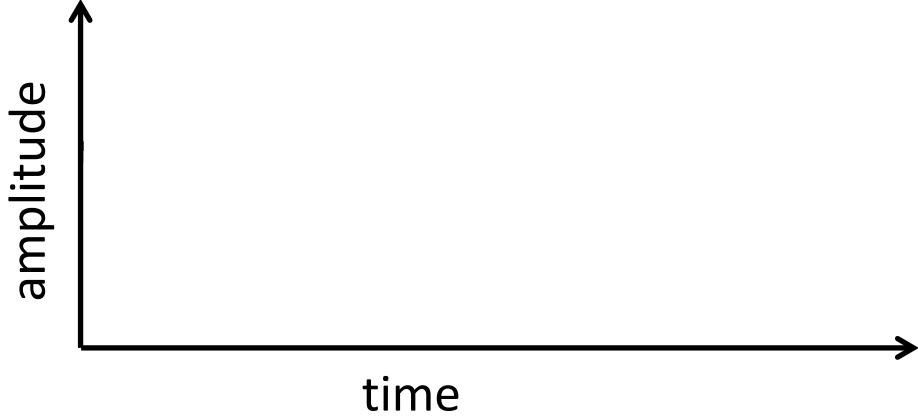
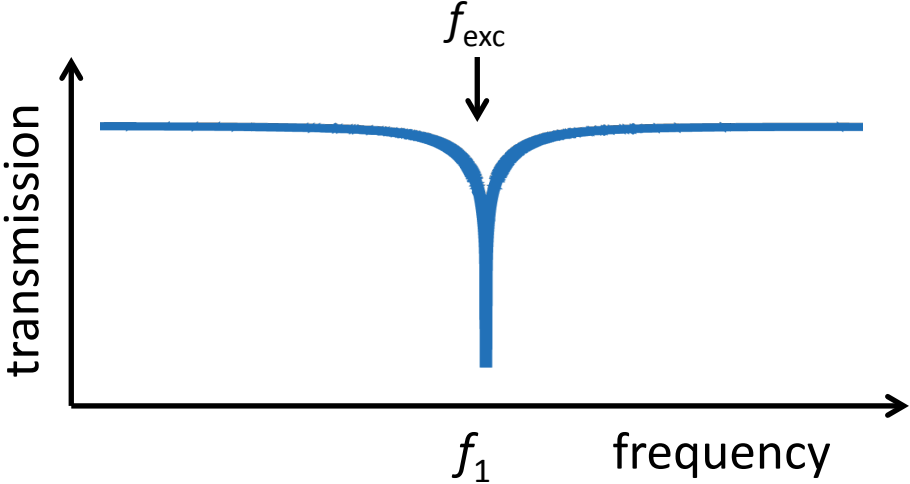
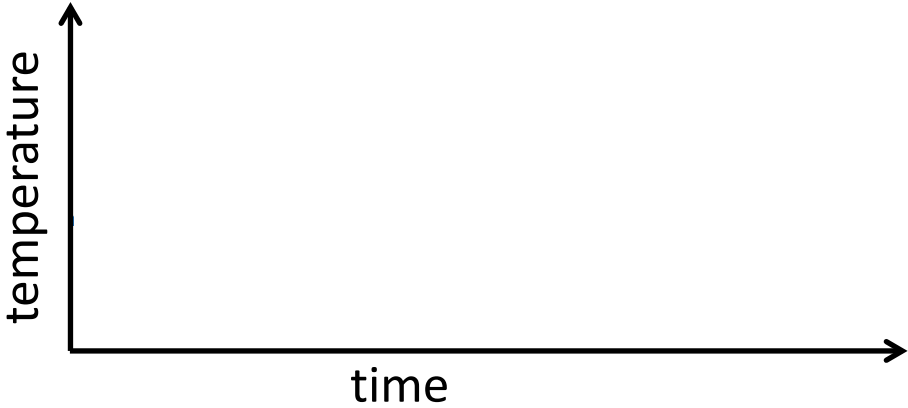
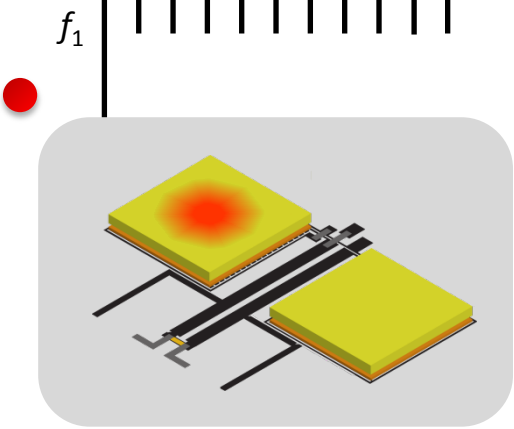


MMCs: Microwave SQUID multiplexing

Single HEMT amplifier and 2 coaxes to read out 100 - 1000 detectors



MMCs: Microwave SQUID multiplexing



Massive neutrinos

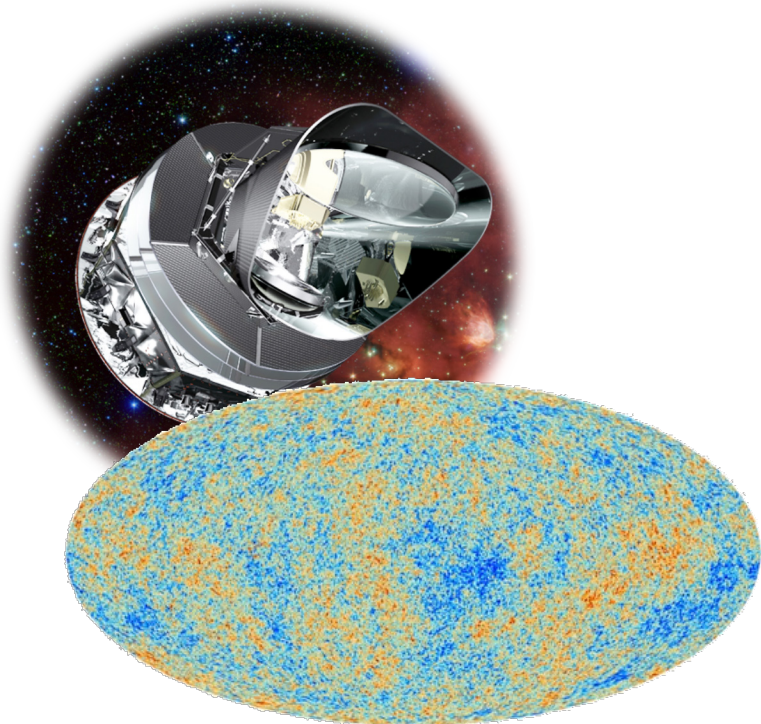


Neutrino mass determination

Cosmology

$$M_\nu = \sum m_i$$

- Model dependent
- Need of satellites
- Present limit 0.12 – 1 eV
- Next future 15-50 meV



Neutrinoless Double beta decay

$$m_{\beta\beta} = \left| \sum_i U_{ei}^2 m_i \right|$$

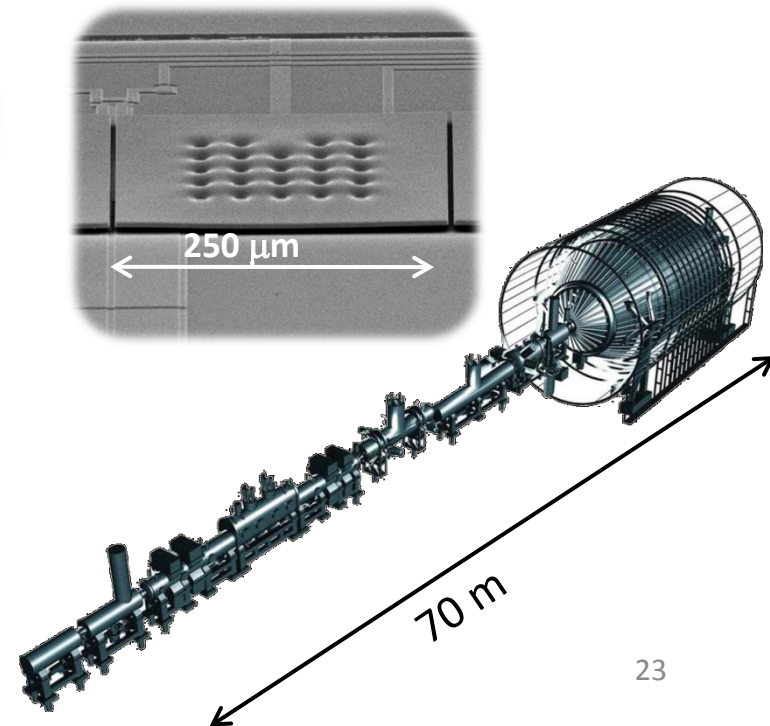
- Model dependent
- Laboratory experiments
- Present limit 0.1 – 0.4 eV
- Next future 15-50 meV



Kinematics of β -decay and electron capture

$$m^2(\nu_e) = \sum_i |U_{ei}|^2 m_i^2$$

- Model independent
- Laboratory experiments
- Present limit 2 eV
- Next future 200 meV



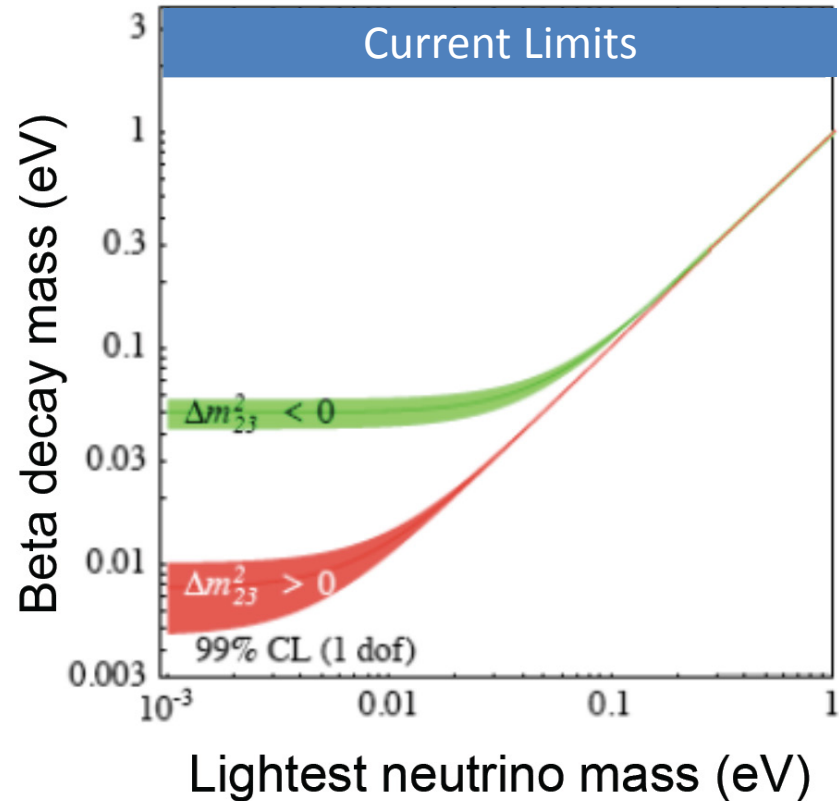
Direct neutrino mass determination

Kinematics of beta decay

$$m^2(\nu_e) = \sum_i |U_{ei}|^2 m_i^2$$

- Model independent
- Laboratory experiments

$$m(\bar{\nu}_e) < 2 \text{ eV} \quad {}^3\text{H} \quad (1)$$



(1) Ch. Kraus *et al.*, Eur. Phys. J. C **40** (2005) 447
N. Aseev *et al.*, Phys. Rev D **84** (2011) 112003

Direct neutrino mass determination

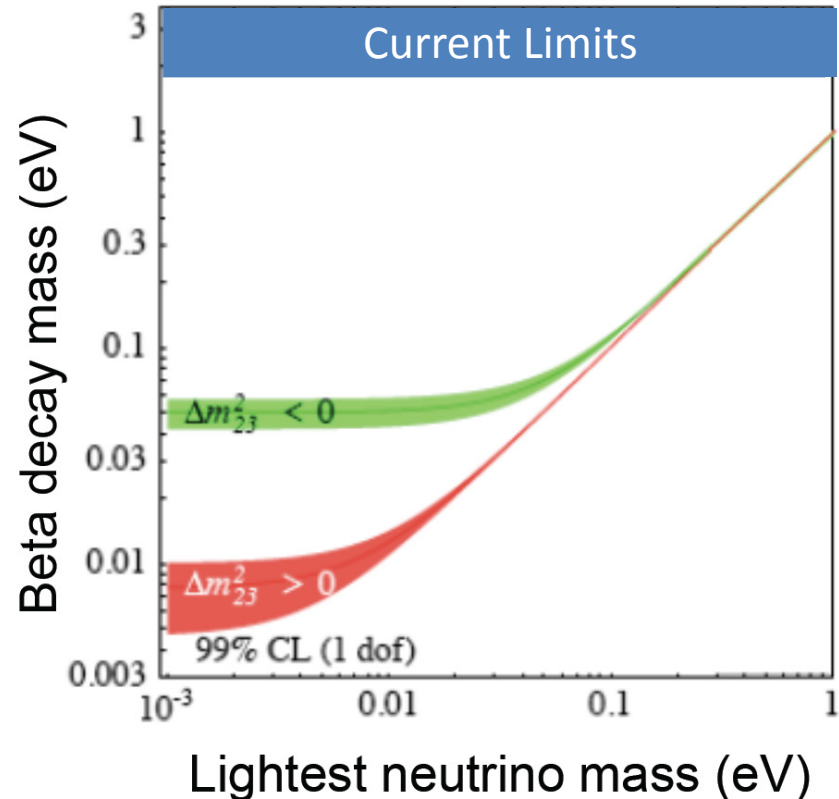
Kinematics of beta decay

$$m^2(\nu_e) = \sum_i |U_{ei}|^2 m_i^2$$

- Model independent
- Laboratory experiments

$$m(\bar{\nu}_e) < 2 \text{ eV} \quad {}^3\text{H} \quad (1)$$

$$m(\nu_e) < 225 \text{ eV} \quad {}^{163}\text{Ho} \quad (2)$$



(1) Ch. Kraus *et al.*, Eur. Phys. J. C **40** (2005) 447
N. Aseev *et al.*, Phys. Rev D **84** (2011) 112003

(2) P. T. Springer, C. L. Bennett, and P. A. Baisden Phys. Rev. A 35 (1987) 679

Direct neutrino mass determination

Kinematics of beta decay

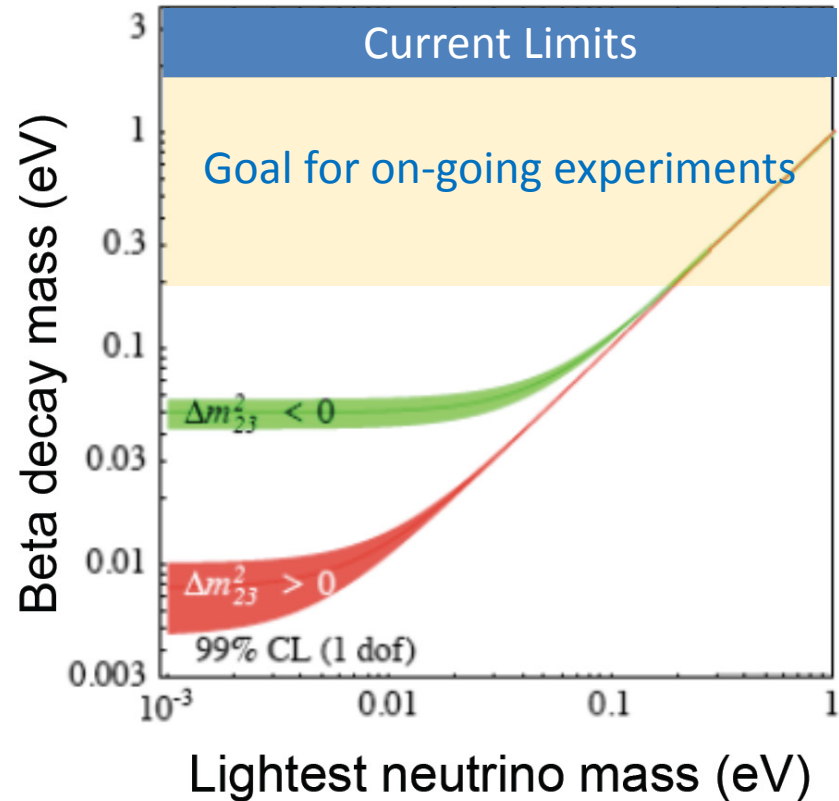
$$m^2(\nu_e) = \sum_i |U_{ei}|^2 m_i^2$$

- Model independent
- Laboratory experiments

$$m(\bar{\nu}_e) < 2 \text{ eV} \quad {}^3\text{H} \quad (1)$$

$$m(\nu_e) < 225 \text{ eV} \quad {}^{163}\text{Ho} \quad (2)$$

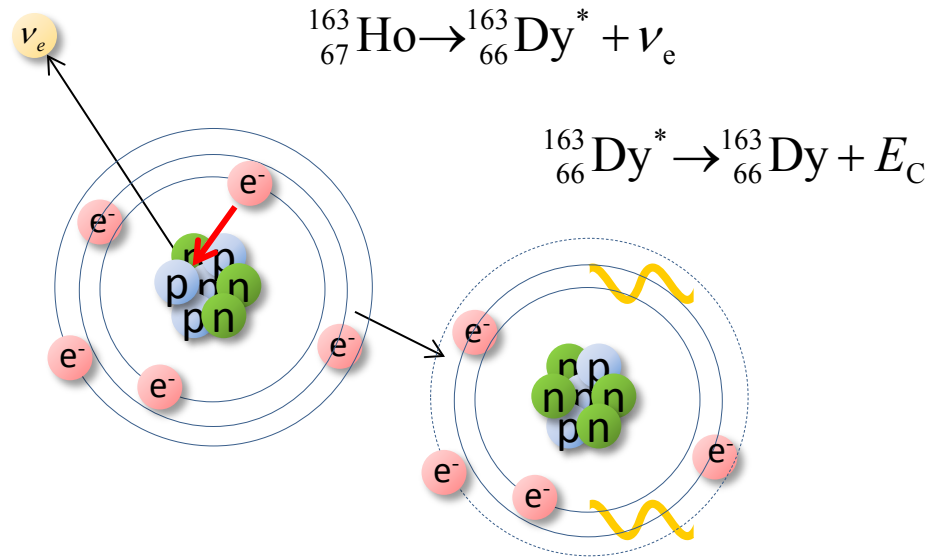
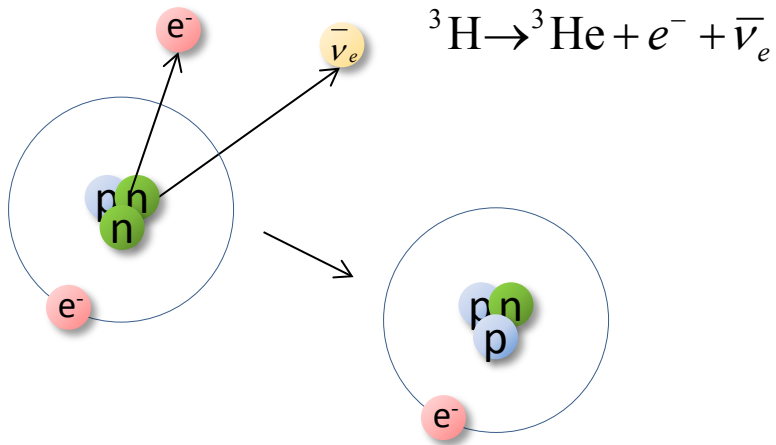
- Next future 200 meV



(1) Ch. Kraus *et al.*, Eur. Phys. J. C **40** (2005) 447
N. Aseev *et al.*, Phys. Rev D **84** (2011) 112003

(2) P. T. Springer, C. L. Bennett, and P. A. Baisden Phys. Rev. A 35 (1987) 679

Beta decay and electron capture



- $\tau_{1/2} \cong 12.3 \text{ years}$ ($4 \cdot 10^8$ atoms for 1 Bq)

- $Q_\beta = 18\,592.01(7) \text{ eV}$

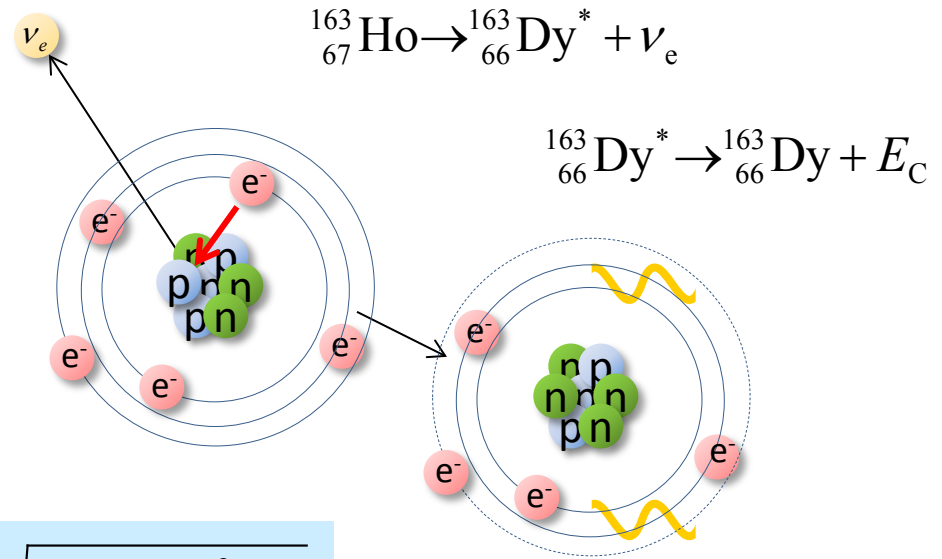
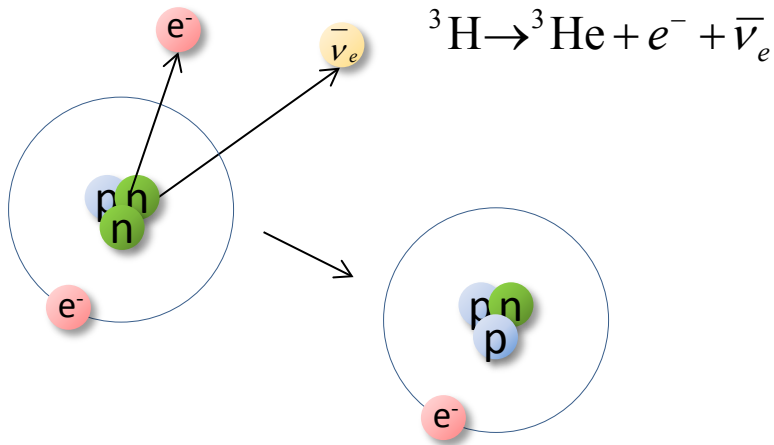
E.G. Myers et al., *Phys. Rev. Lett.* **114** (2015) 013003

- $\tau_{1/2} \cong 4570 \text{ years}$ ($2 \cdot 10^{11}$ atoms for 1 Bq)

- $Q_{EC} = (2.833 \pm 0.030^{\text{stat}} \pm 0.015^{\text{syst}}) \text{ keV}$

S. Eliseev et al., *Phys. Rev. Lett.* **115** (2015) 062501

Beta decay and electron capture



$$\frac{dW}{dE} \propto (Q - E)^2 \sqrt{1 - \frac{m_\nu^2}{(Q - E)^2}}$$

- $\tau_{1/2} \cong 12.3 \text{ years}$ ($4 \cdot 10^8$ atoms for 1 Bq)

- $Q_\beta = 18\,592.01(7) \text{ eV}$

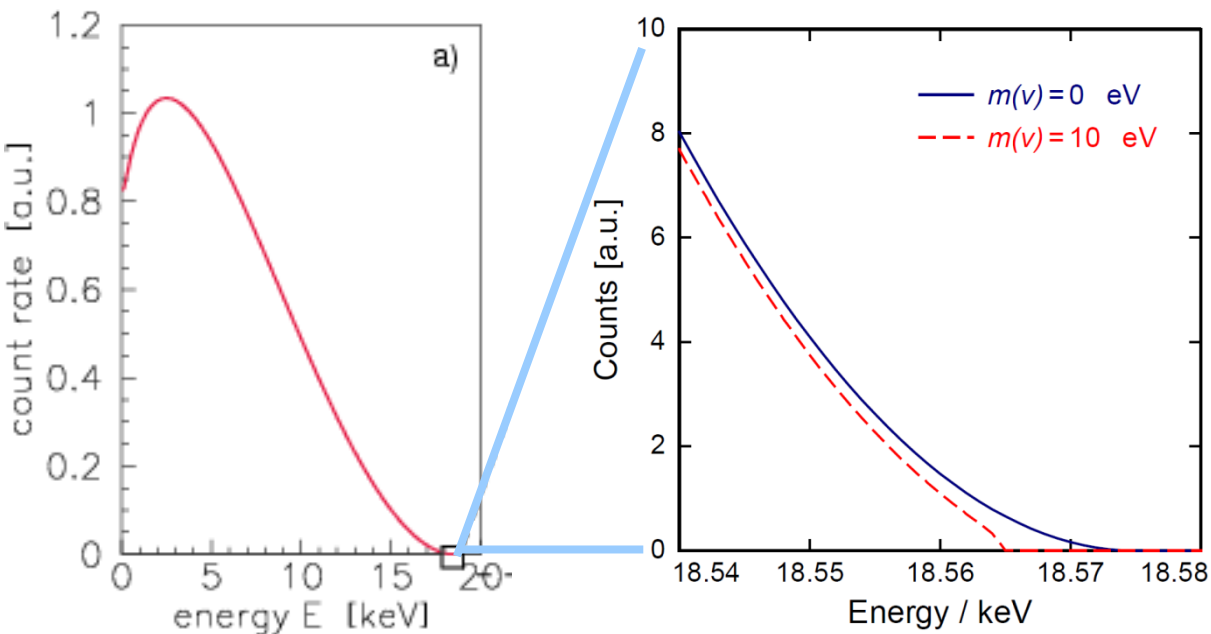
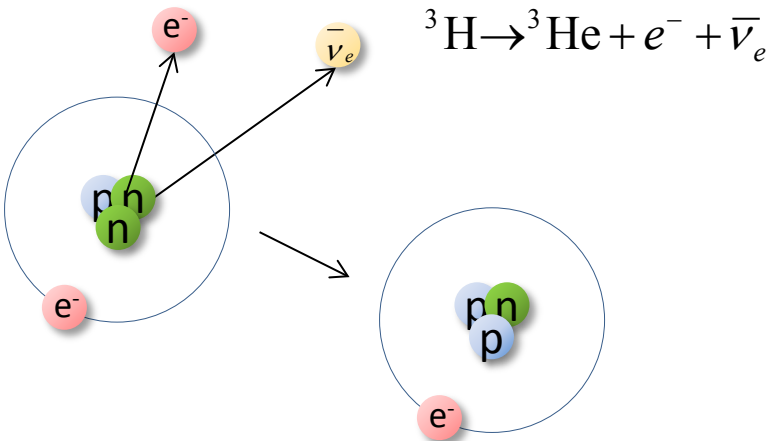
E.G. Myers et al., *Phys. Rev. Lett.* **114** (2015) 013003

- $\tau_{1/2} \cong 4570 \text{ years}$ ($2 \cdot 10^{11}$ atoms for 1 Bq)

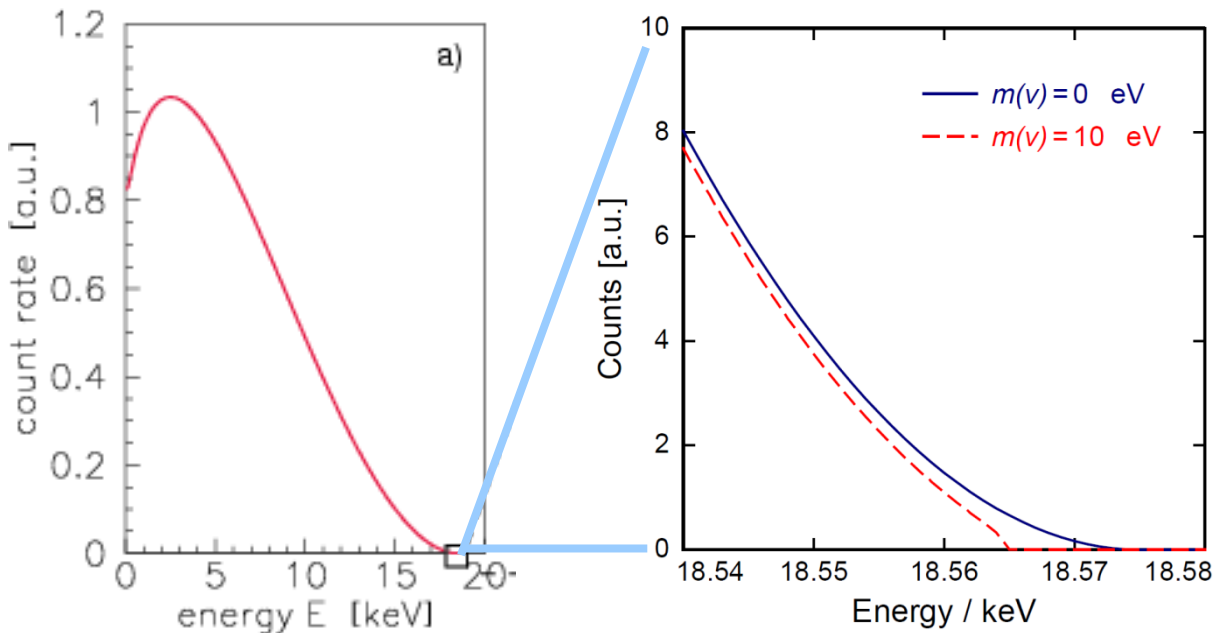
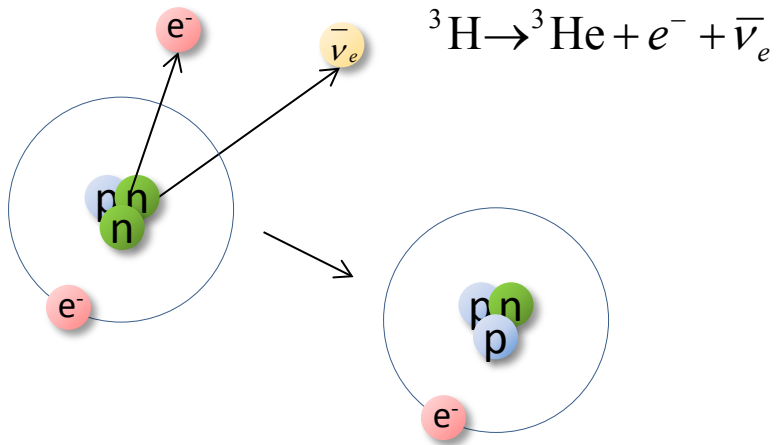
- $Q_{EC} = (2.833 \pm 0.030^{\text{stat}} \pm 0.015^{\text{syst}}) \text{ keV}$

S. Eliseev et al., *Phys. Rev. Lett.* **115** (2015) 062501

Beta decay of ^3H



Beta decay of ^3H

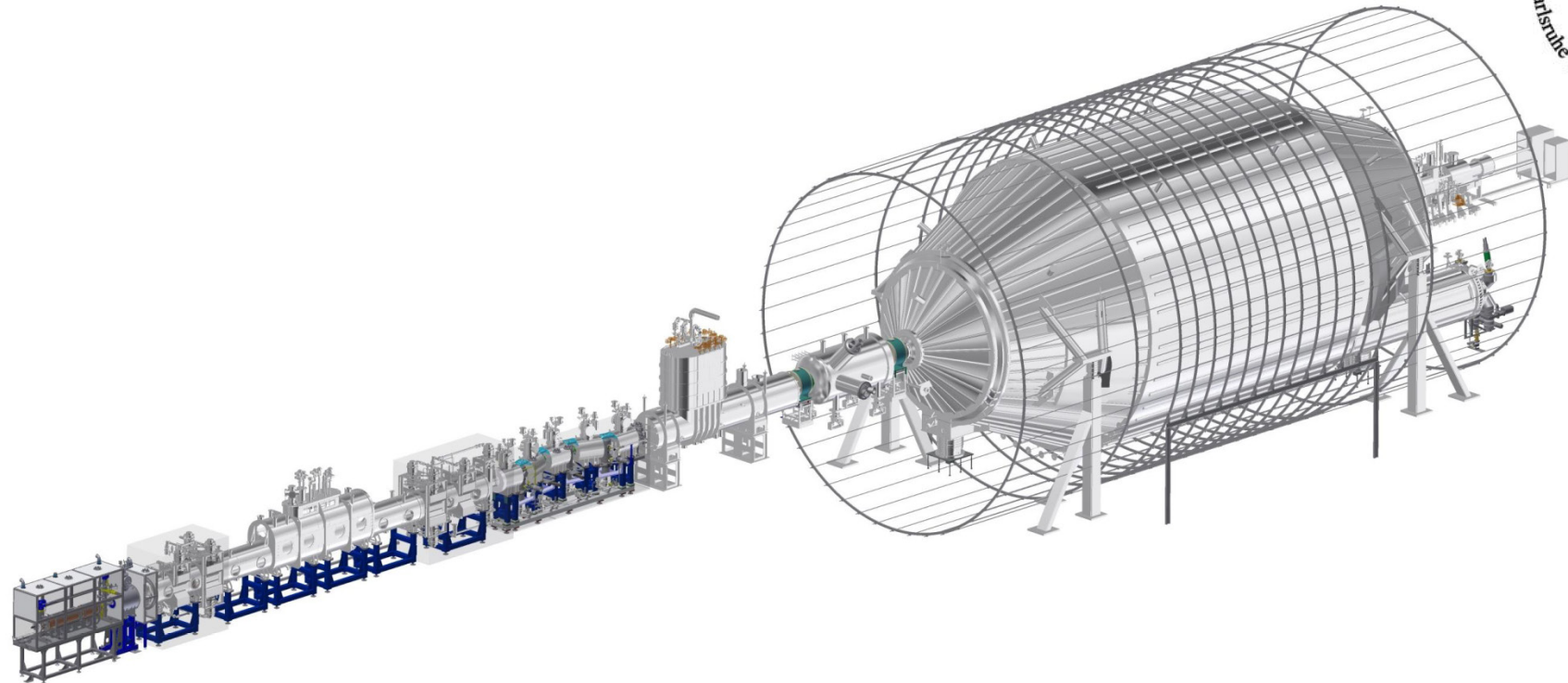


Only a small fraction of events
in the last eV below the endpoint:
 $2 \cdot 10^{-13}$

Very low background is required

The KATRIN experiment

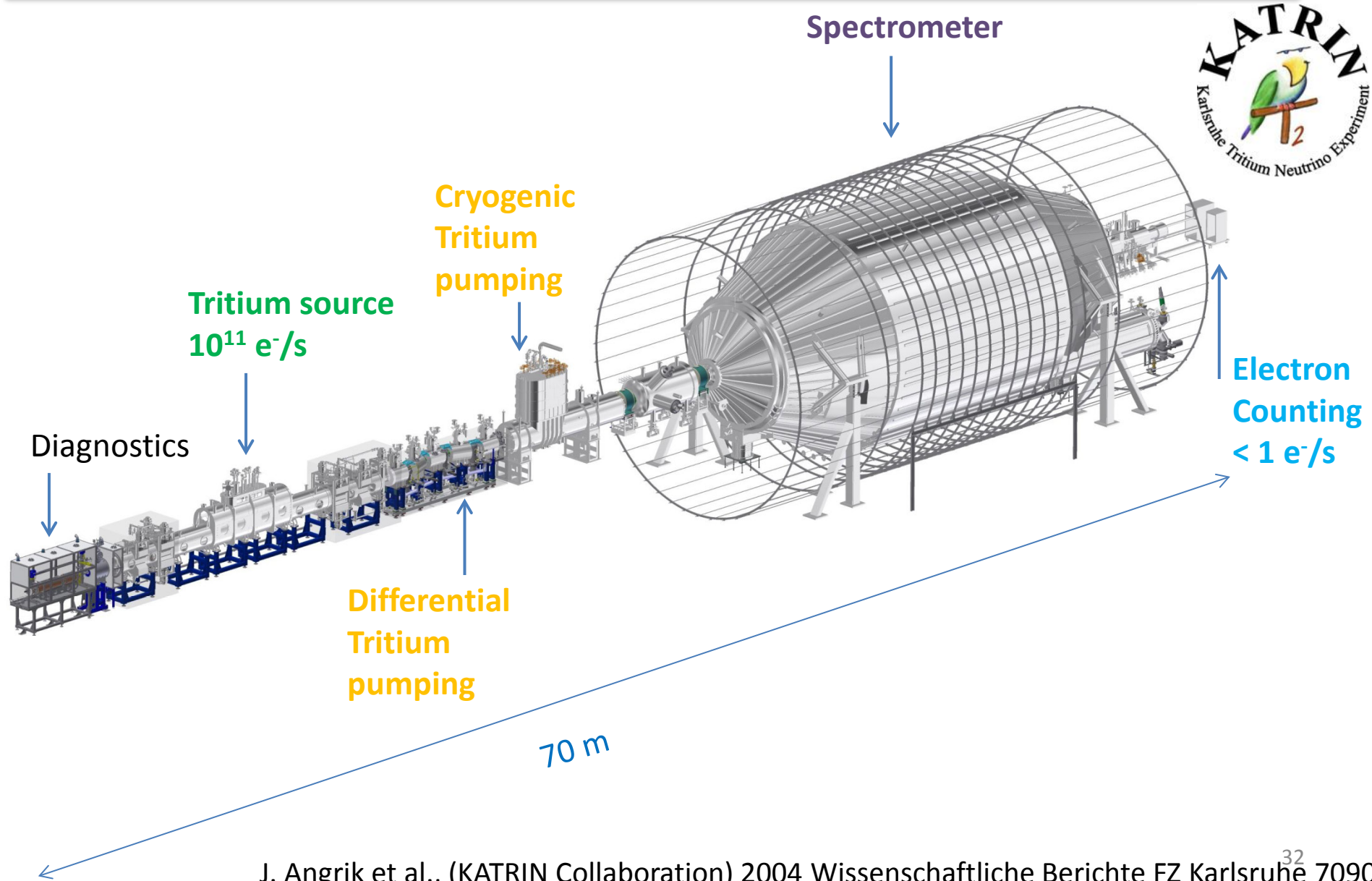
❖ KATRIN - Karlsruhe Tritium Neutrino Experiment



Main ideas:

- high activity source 10^{11} e⁻/s
- high resolution MAC-E* filter to select electrons close to the end point
- count electrons as function of retarding potential
→ integral spectrum

The KATRIN experiment



The KATRIN experiment: present status



Large Helmholtz coil system

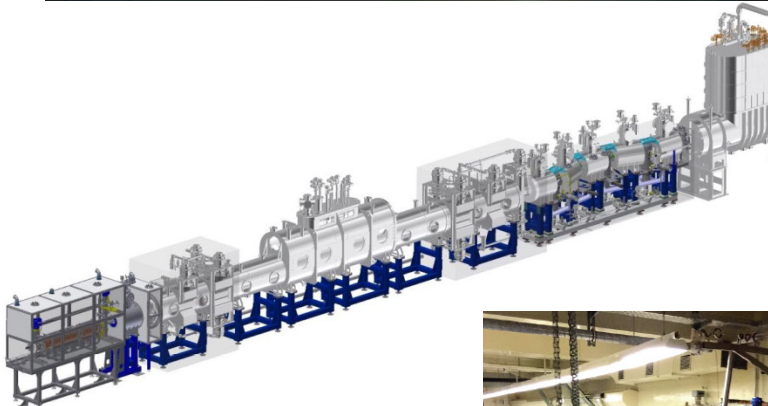
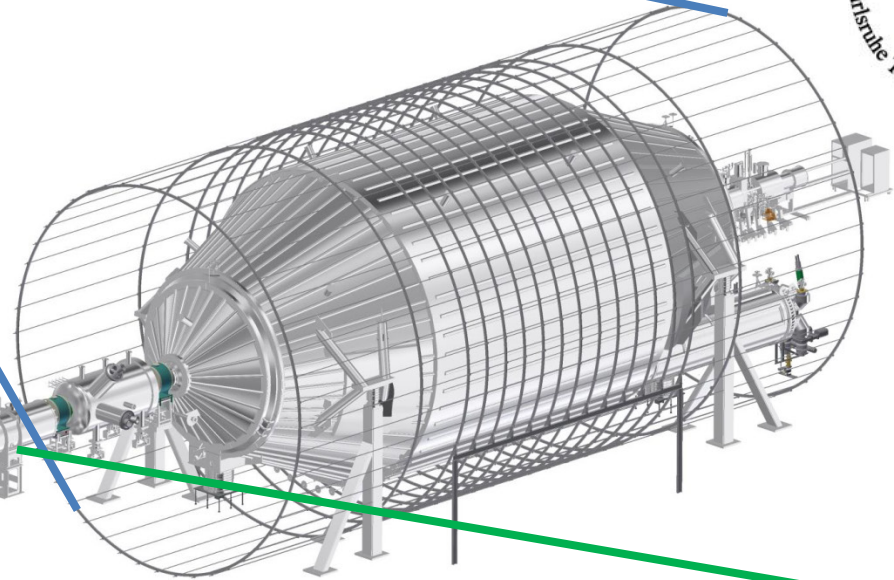


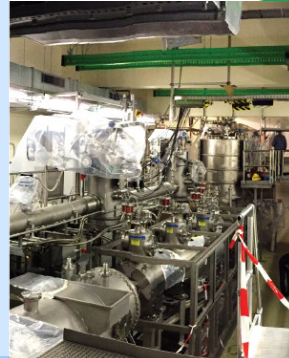
Photo K. Valerius

The KATRIN experiment: present status



First light 14th October 2016

Photo Patrick Langer



^3H based experiments

❖ KATRIN - Karlsruhe Tritium Neutrino Experiment

- Main ideas:
- high activity source: 10^{11} e⁻/s
 - high resolution MAC-E filter to select electrons close to the end point
 - count electrons as function of retarding potential
→ integral spectrum



❖ Project8

- Main ideas:
- Source = detector: $10^{11} - 10^{13}$ $^3\text{H}_2$ molecules /cm³
 - Use cyclotron frequency to extract electron energy
 - Differential spectrum

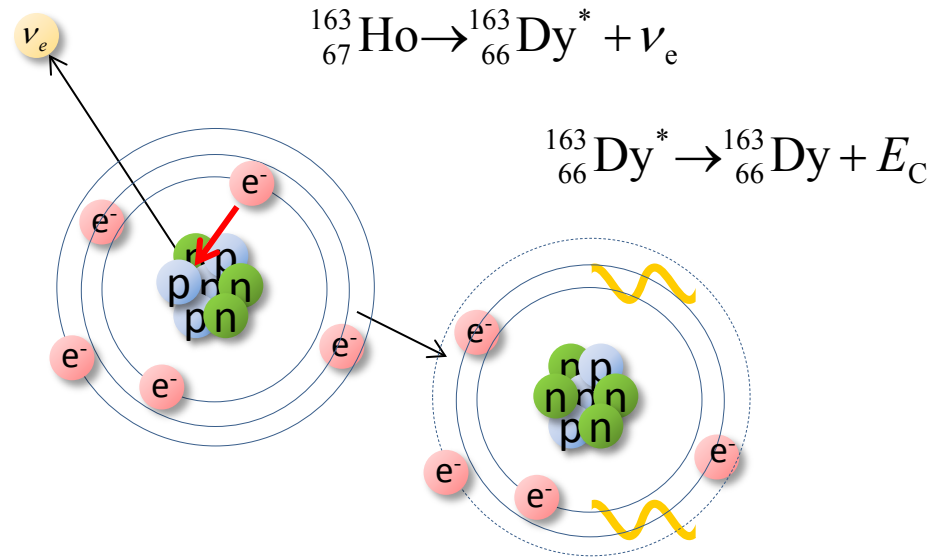
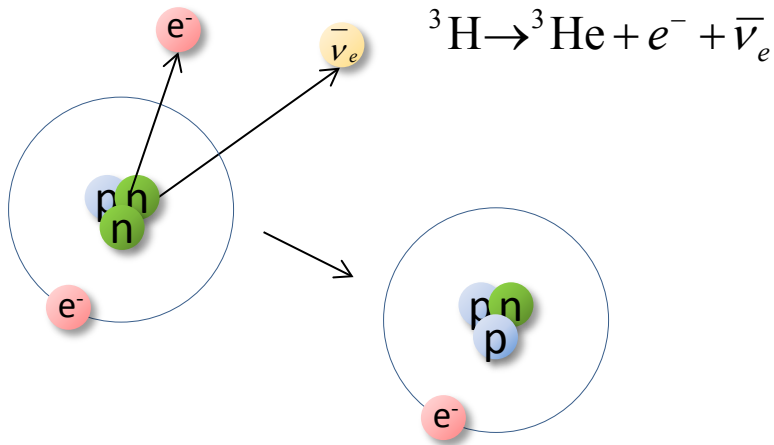


❖ PTOLEMY - Princeton Tritium Observatory for Light, Early-Universe, Massive-Neutrino Yield

- Main ideas:
- large area tritium source: 100 g atomic ^3H
 - MAC-E filter to select electrons close to the end point
 - RF tracking and time-of-flight systems
 - cryogenic calorimetry → differential spectrum



Beta decay and electron capture



- $\tau_{1/2} \cong 12.3 \text{ years}$ ($4 \cdot 10^8$ atoms for 1 Bq)

- $Q_{EC} = 18\,592.01(7) \text{ eV}$

E.G. Myers et al., *Phys. Rev. Lett.* **114** (2015) 013003

- $\tau_{1/2} \cong 4570 \text{ years}$ ($2 \cdot 10^{11}$ atoms for 1 Bq)

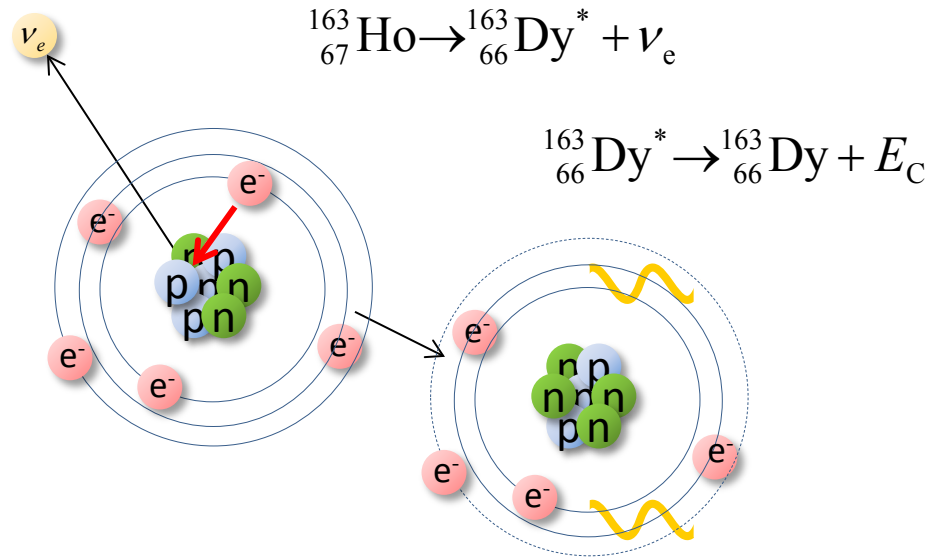
- $Q_{EC} = (2.833 \pm 0.030^{\text{stat}} \pm 0.015^{\text{syst}}) \text{ keV}$

S. Eliseev et al., *Phys. Rev. Lett.* **115** (2015) 062501

Electron capture in ^{163}Ho : spectrum

Atomic de-excitation:

- X-ray emission
- Auger electrons
- Coster-Kronig transitions



• $\tau_{1/2} \cong 4570$ years ($2 \cdot 10^{11}$ atoms for 1 Bq)

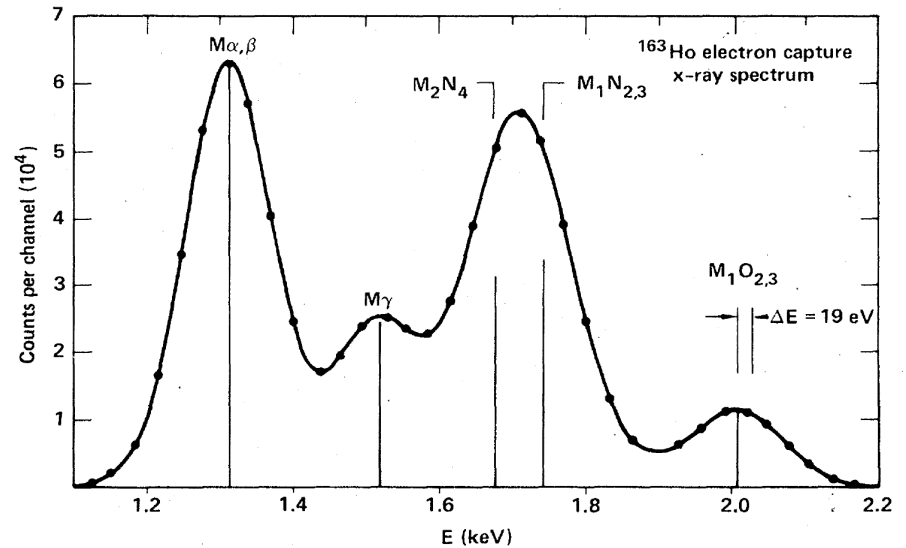
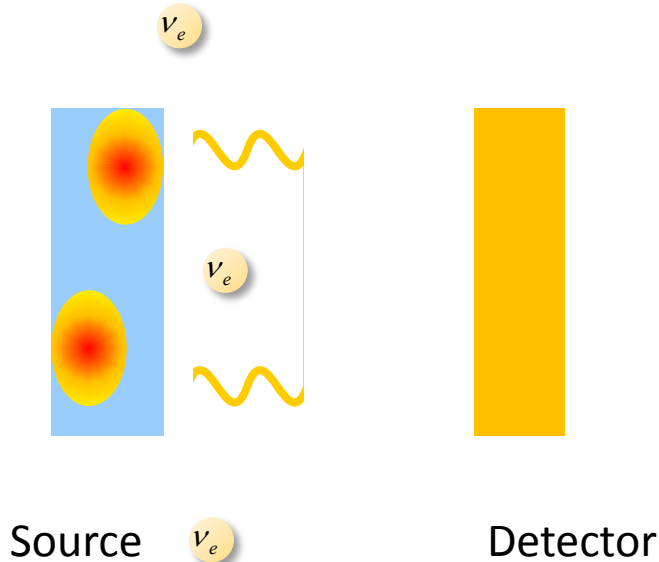
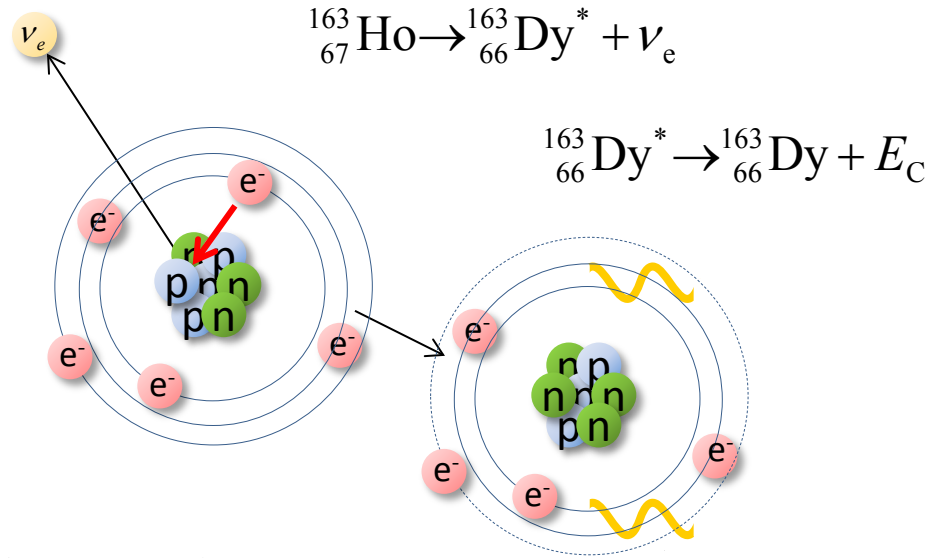
• $Q_{\text{EC}} = (2.833 \pm 0.030^{\text{stat}} \pm 0.015^{\text{syst}})$ keV

S. Eliseev et al., *Phys. Rev. Lett.* **115** (2015) 062501

Electron capture in ^{163}Ho : spectrum

Atomic de-excitation:

- X-ray emission
- Auger electrons
- Coster-Kronig transitions

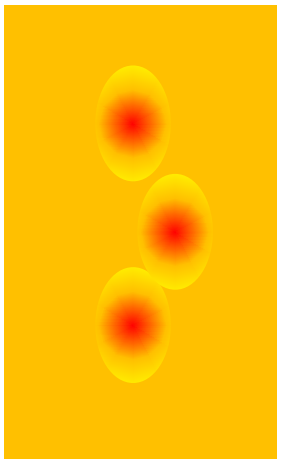
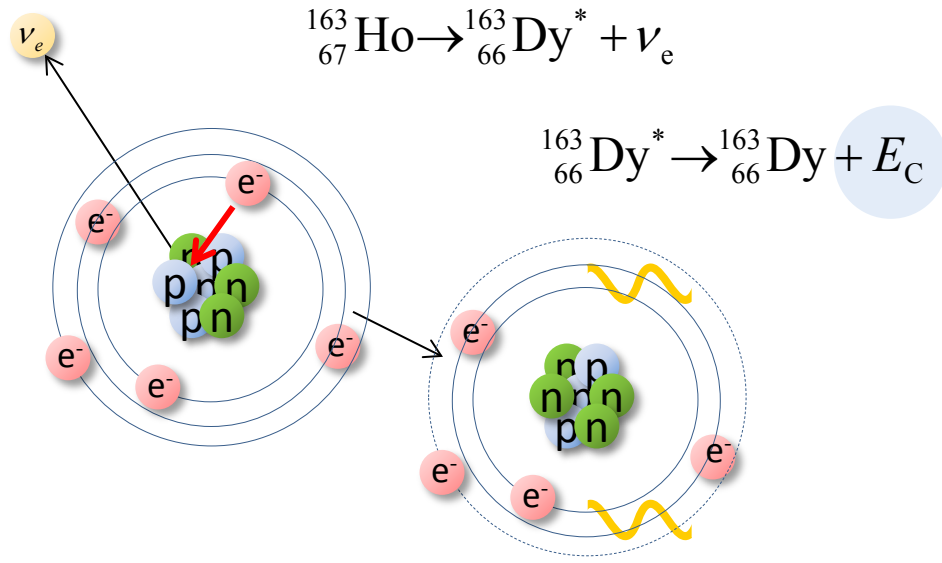


Electron capture in ^{163}Ho : spectrum

Atomic de-excitation:

- X-ray emission
- Auger electrons
- Coster-Kronig transitions

Calorimetric measurement



Source = Detector

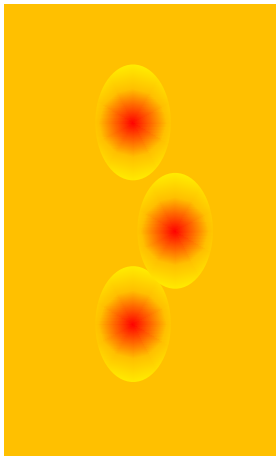


Electron capture in ^{163}Ho : spectrum

Atomic de-excitation:

- X-ray emission
- Auger electrons
- Coster-Kronig transitions

Calorimetric measurement

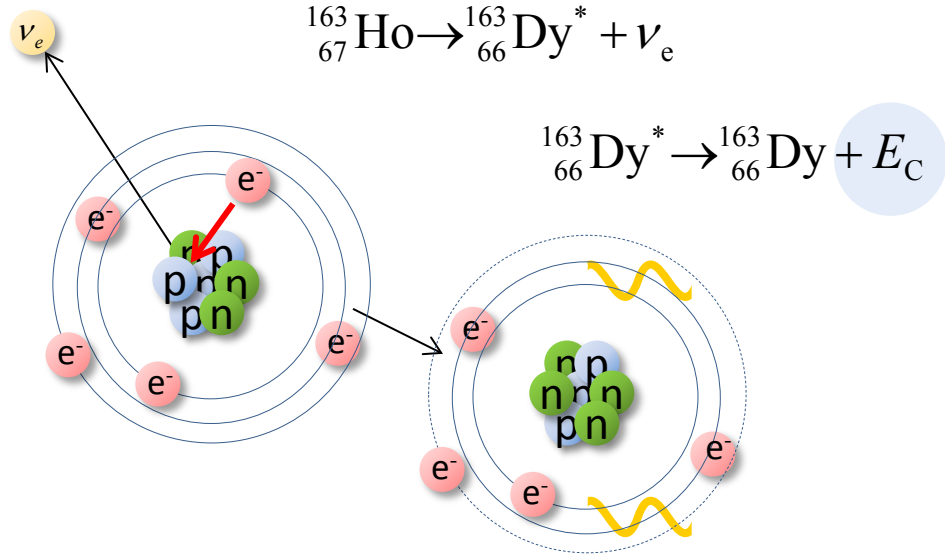


Source = Detector

ν_e

ν_e

ν_e

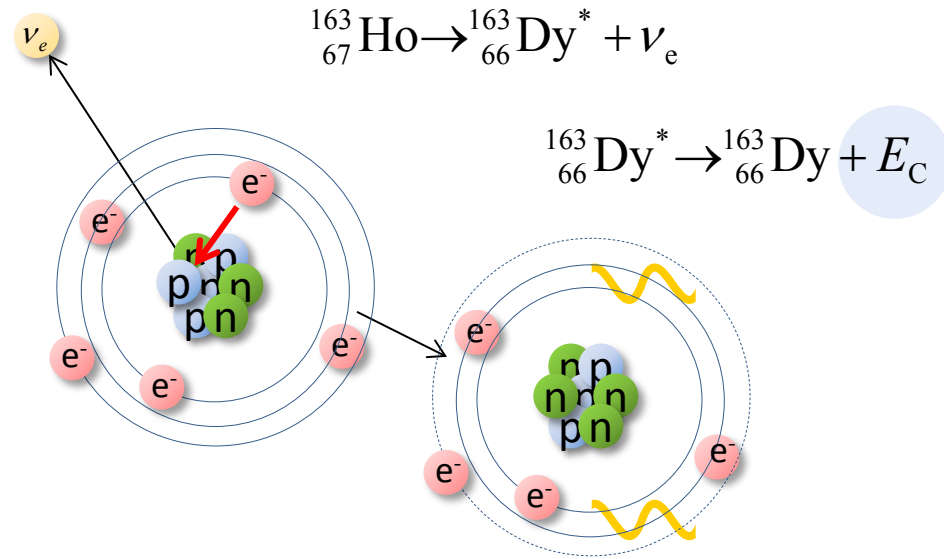


Electron capture in ^{163}Ho : spectrum

Atomic de-excitation:

- X-ray emission
- Auger electrons
- Coster-Kronig transitions

Calorimetric measurement



Volume 118B, number 4, 5, 6

PHYSICS LETTERS

9 December 1982

CALORIMETRIC MEASUREMENTS OF ^{163}Ho DECAY AS TOOLS TO DETERMINE THE ELECTRON NEUTRINO MASS

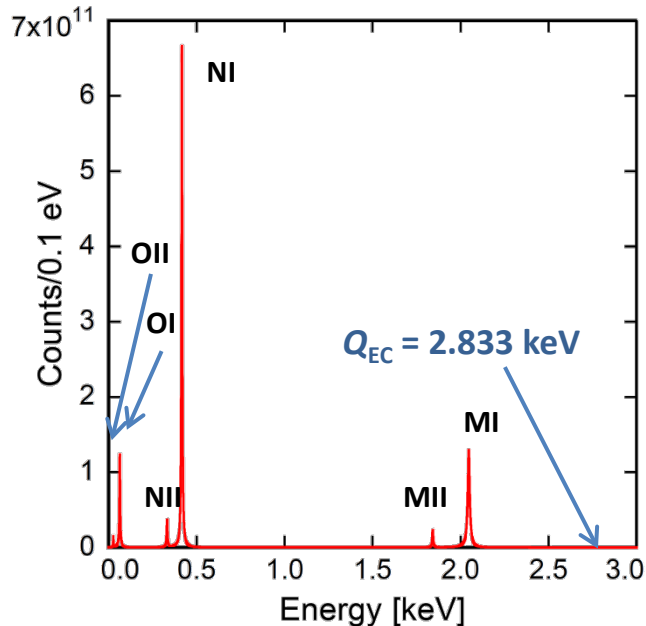
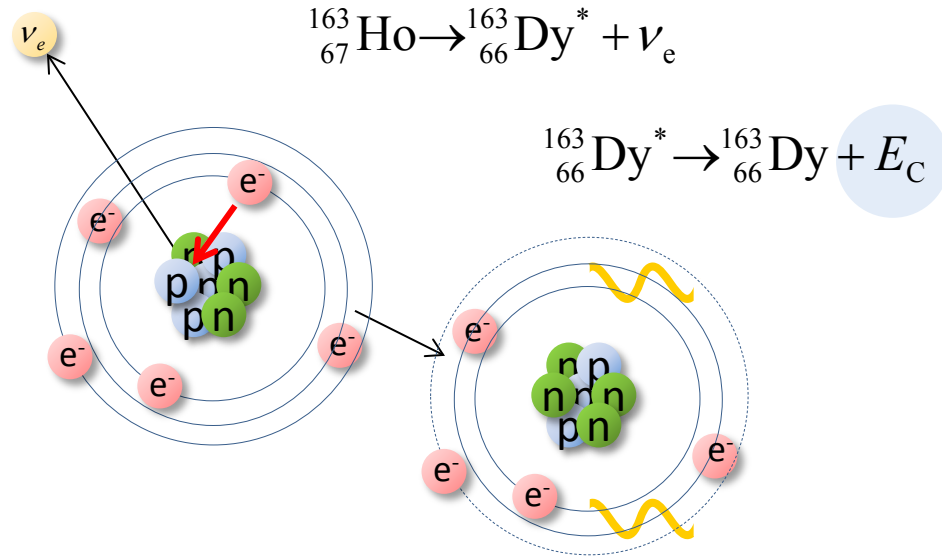
A. DE RÚJULA and M. LUSIGNOLI ¹
CERN, Geneva, Switzerland

Electron capture in ^{163}Ho : spectrum

Atomic de-excitation:

- X-ray emission
- Auger electrons
- Coster-Kronig transitions

Calorimetric measurement

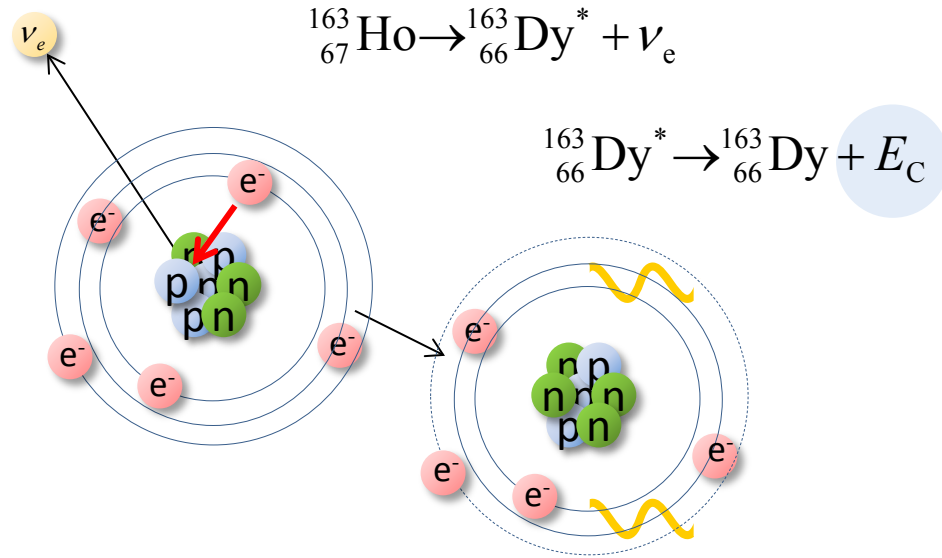


$$\frac{dW}{dE_C} = A(Q_{\text{EC}} - E_C)^2 \sqrt{1 - \frac{m_\nu^2}{(Q_{\text{EC}} - E_C)^2}} \sum_{\text{H}} B_{\text{H}} \phi_{\text{H}}^2(0) \frac{\frac{\Gamma_{\text{H}}}{2\pi}}{(E_C - E_{\text{H}})^2 + \frac{\Gamma_{\text{H}}^2}{4}}$$

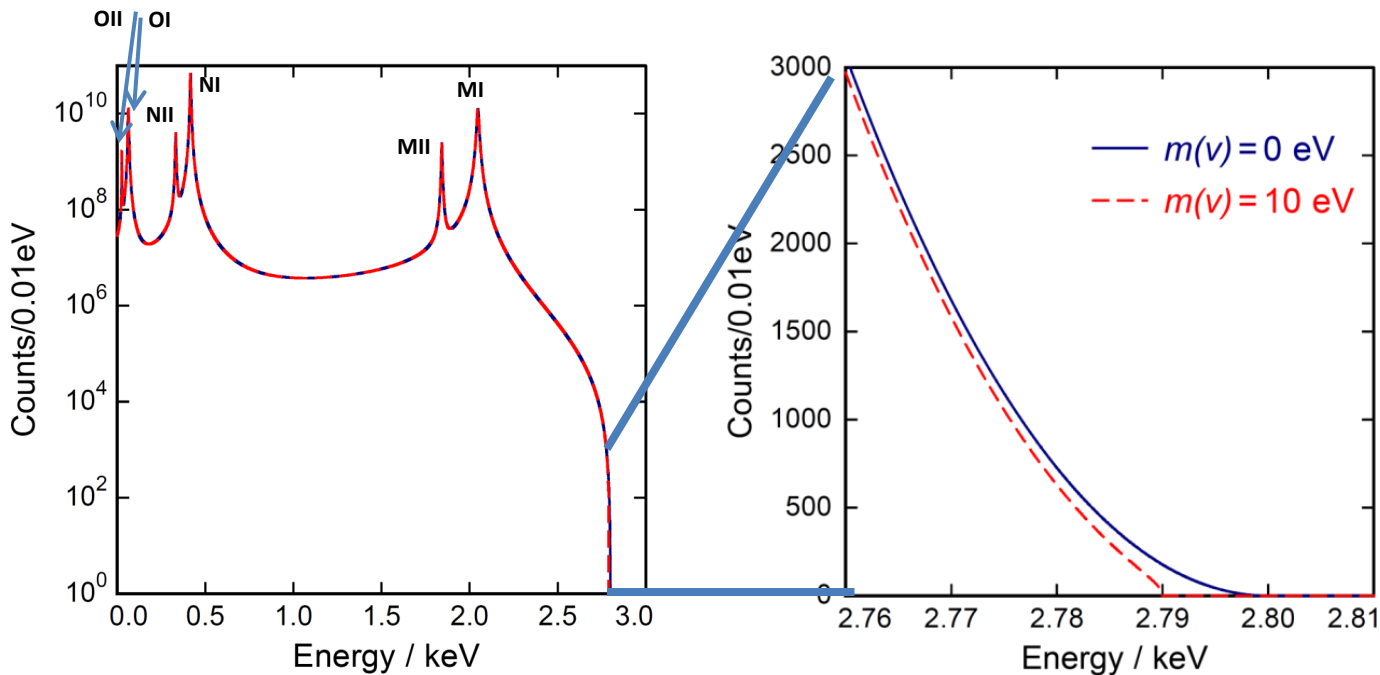
Electron capture in ^{163}Ho : spectrum

Atomic de-excitation:

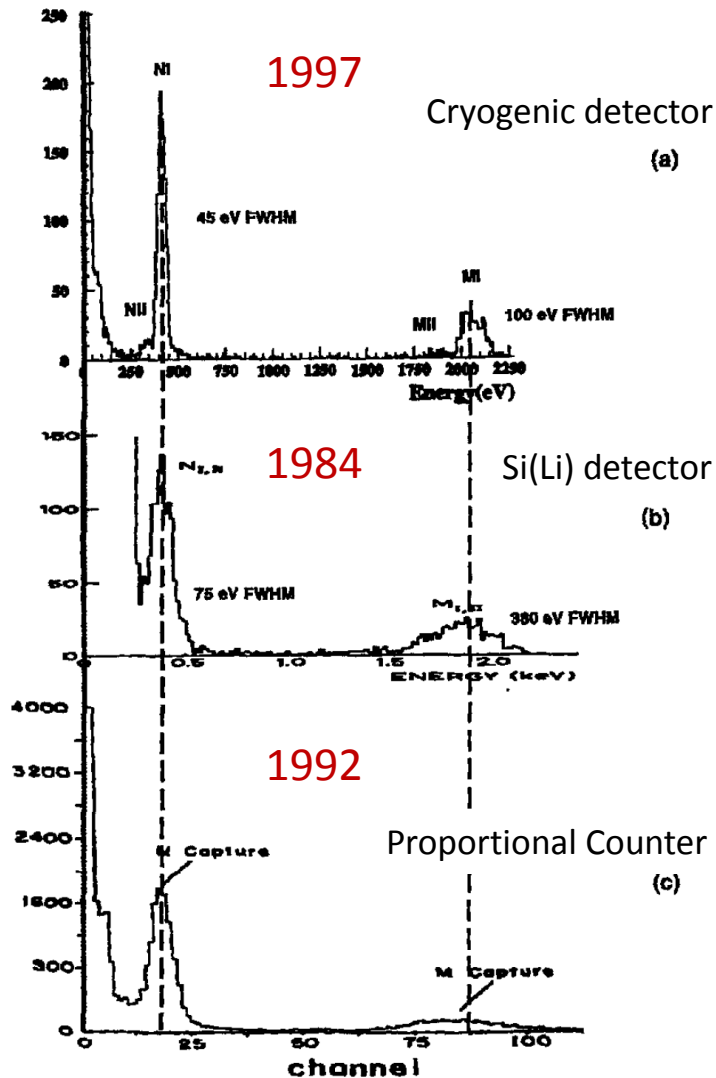
- X-ray emission
- Auger electrons
- Coster-Kronig transitions



Calorimetric measurement



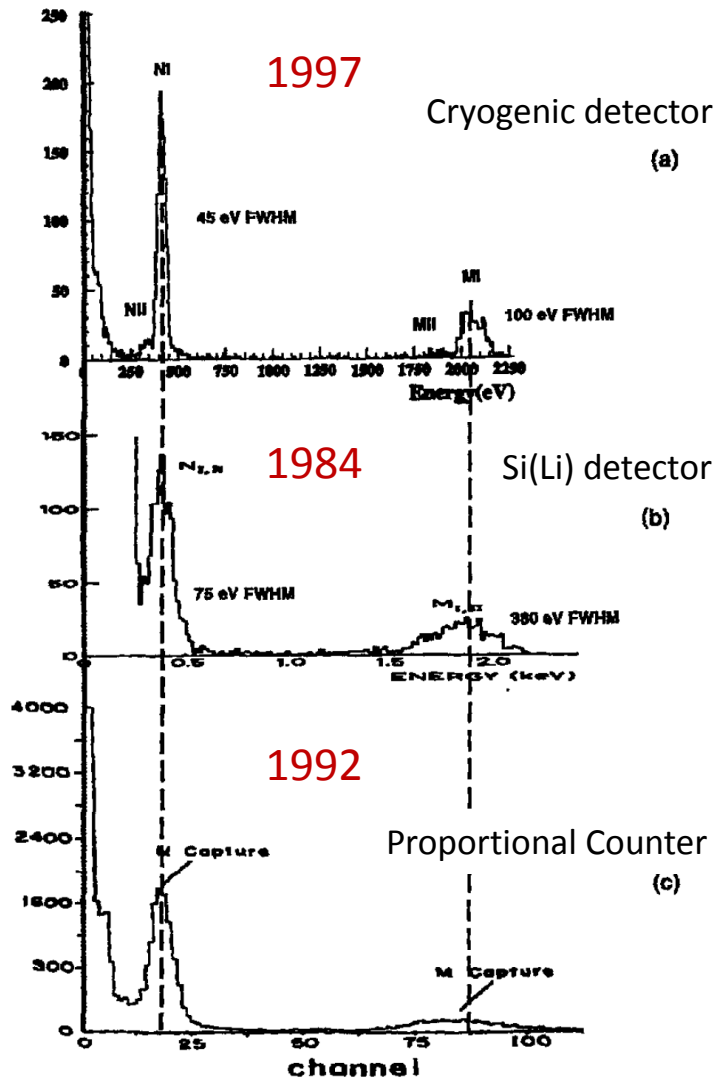
Electron capture in ^{163}Ho : history



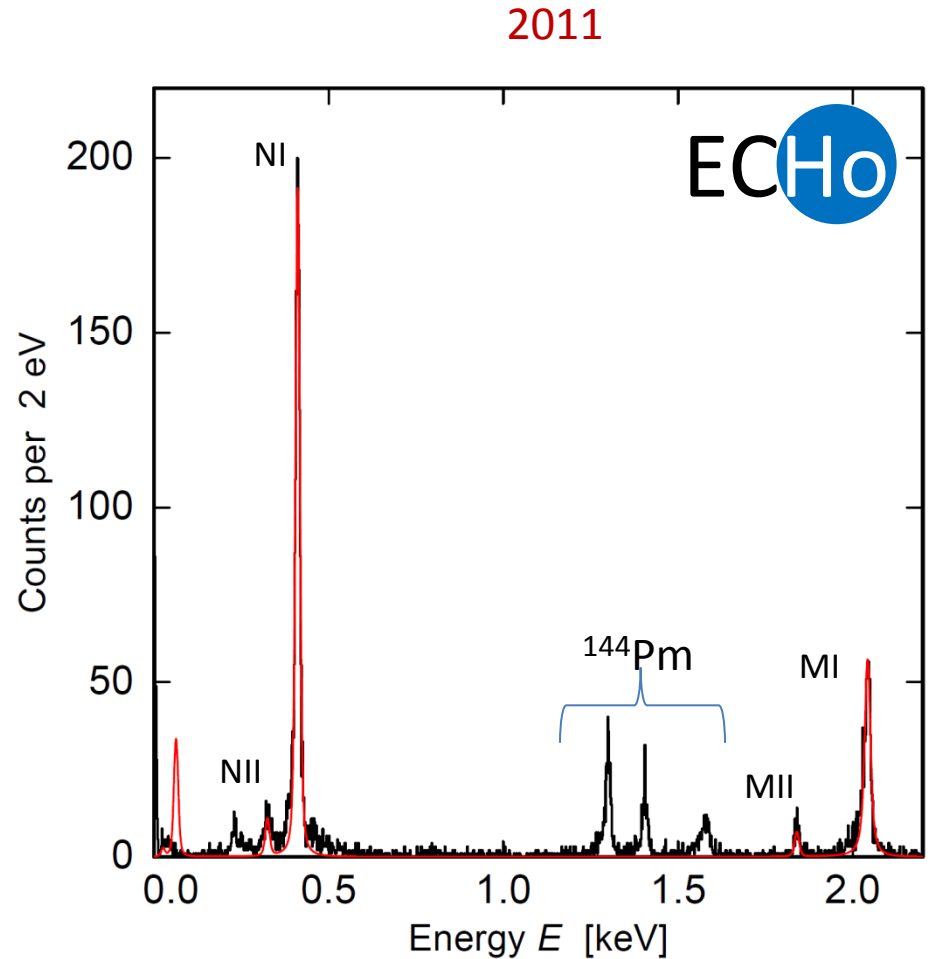
F. Gatti et al., Physics Letters B 398 (1997) 415-419

- (a) F. Gatti et al., Physics Letters B 398 (1997) 415-419
- (b) E. Laesgaard et al., Proceeding of 7th International Conference on Atomic Masses and Fundamental Constants (AMCO-7), (1984).
- (c) F.X. Hartmann and R.A. Naumann, Nucl. Instr. Meth. A 3 13 (1992) 237.

Electron capture in ^{163}Ho : history

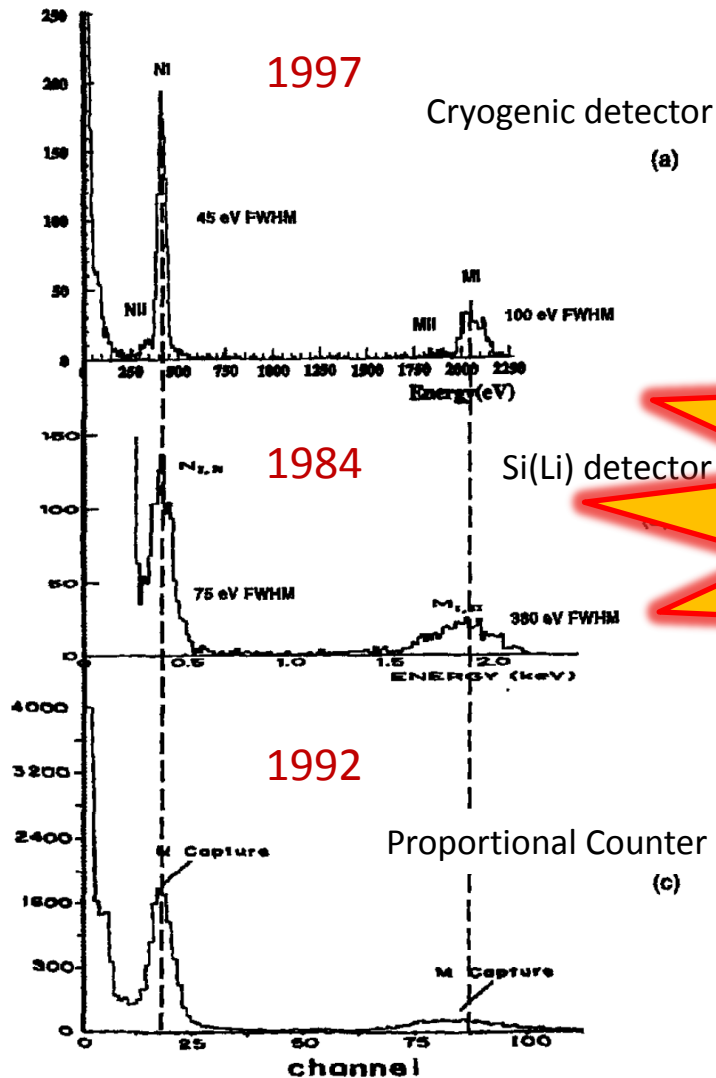


F. Gatti et al., Physics Letters B 398 (1997) 415-419

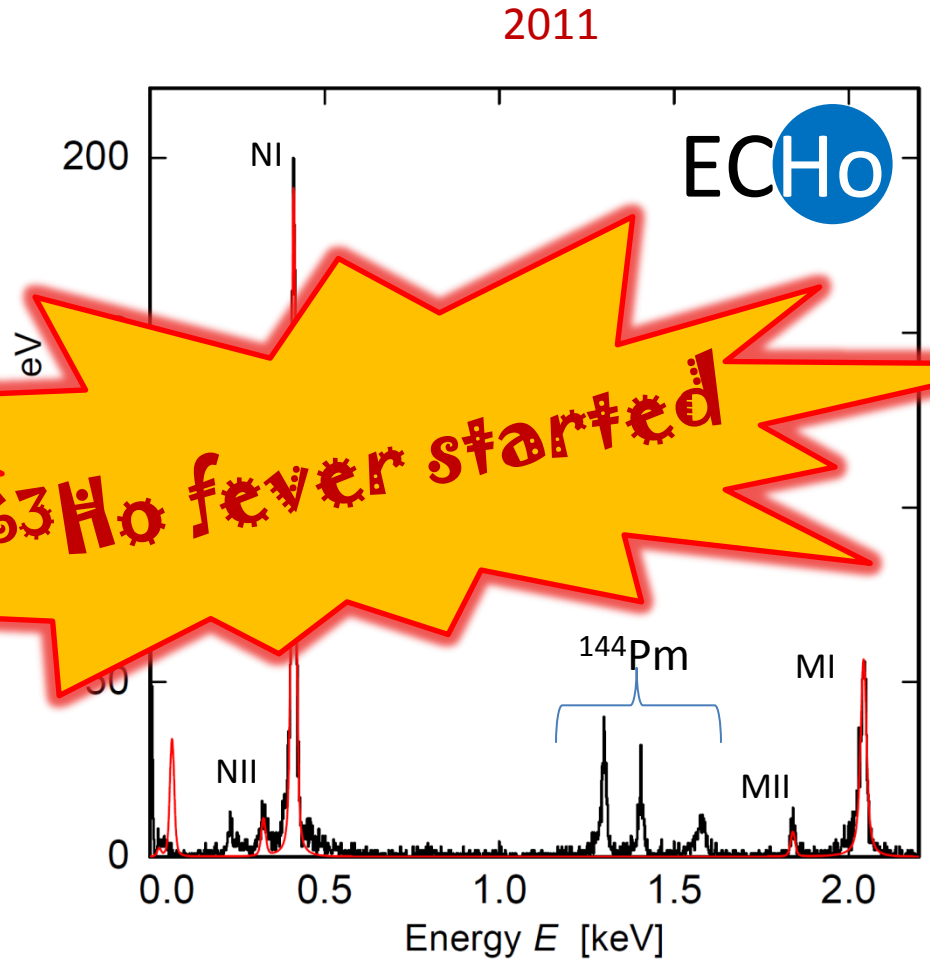


- (a) F. Gatti et al., Physics Letters B 398 (1997) 415-419
- (b) E. Laesgaard et al., Proceeding of 7th International Conference on Atomic Masses and Fundamental Constants (AMCO-7), (1984).
- (c) F.X. Hartmann and R.A. Naumann, Nucl. Instr. Meth. A 3 13 (1992) 237.

Electron capture in ^{163}Ho : history



F. Gatti et al., Physics Letters B 398 (1997) 415-419



- (a) F. Gatti et al., Physics Letters B 398 (1997) 415-419
- (b) E. Laesgaard et al., Proceeding of 7th International Conference on Atomic Masses and Fundamental Constants (AMCO-7), (1984).
- (c) F.X. Hartmann and R.A. Naumann, Nucl. Instr. Meth. A 3 13 (1992) 237.

Electron capture in ^{163}Ho : present

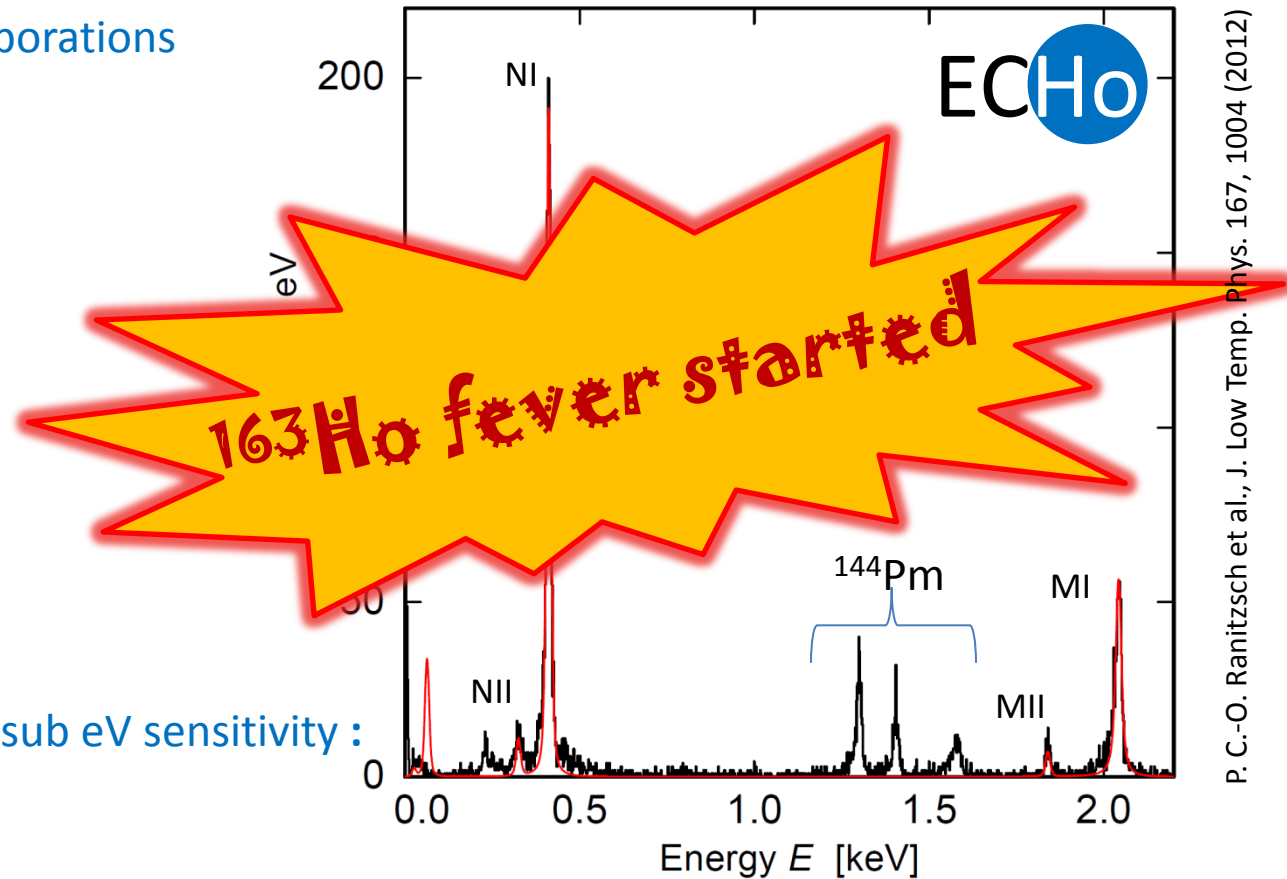
- Calorimetric measurement of the ^{163}Ho spectrum
- Three international collaborations

2011

ECHo (1)

HLMES (2)

NuMECS (3)



P. C.-O. Ranitzsch et al., J. Low Temp. Phys. 167, 1004 (2012)

Common challenges to reach sub eV sensitivity :

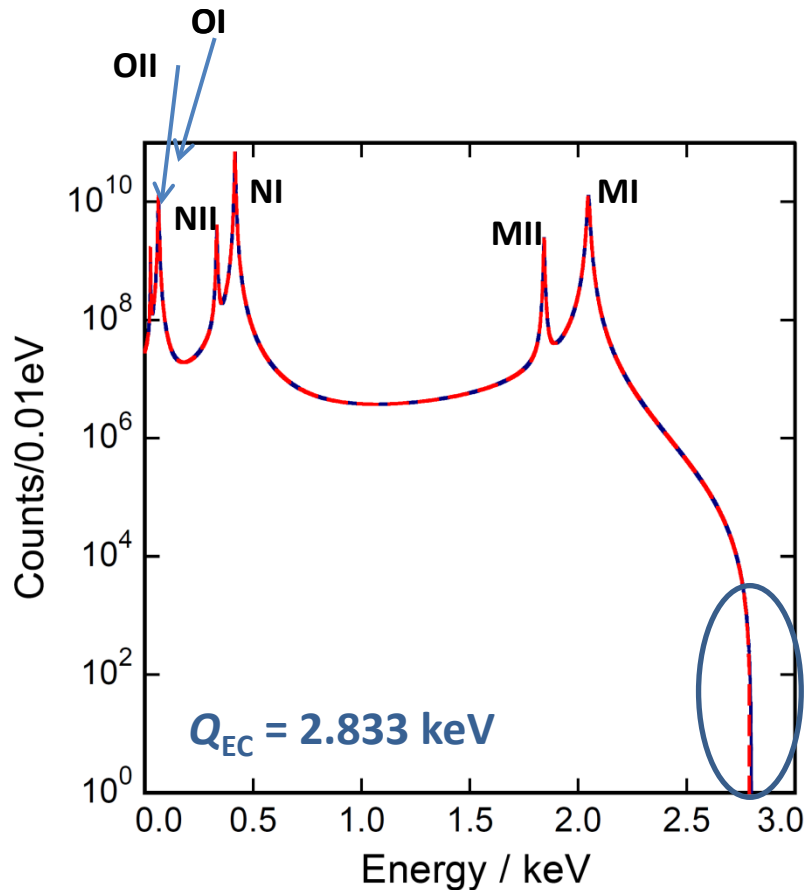
- Detector performance
- High purity ^{163}Ho source
- Background reduction
- Description of the ^{163}Ho EC spectrum

- (1) The ECHo Collaboration EPJ-ST 226 8 (2017) 1623
- (2) B. Alpert et al, Eur. Phys. J. C (2015) 75:112
- (3) M. Croce et al., arXiv:1510.03874

Requirements for sub-eV sensitivity in ECHO

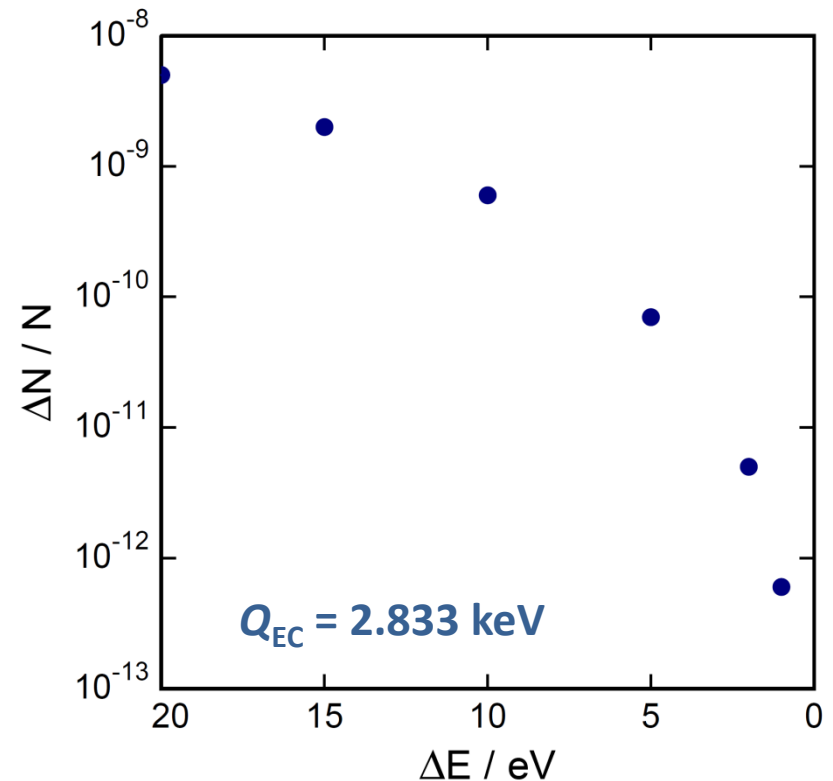
Statistics in the end point region

- $N_{ev} > 10^{14} \rightarrow A \approx 1 \text{ MBq}$



Fraction of events at endpoint regions

- In the interval 2.832 -2.833 keV only 6×10^{-13}



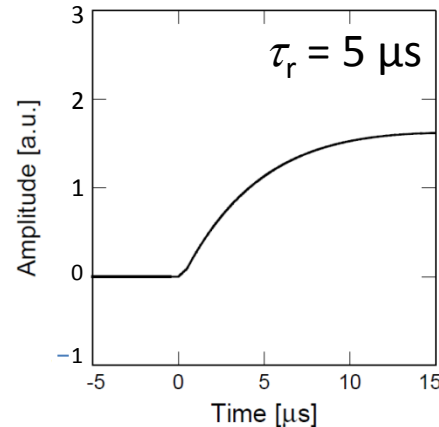
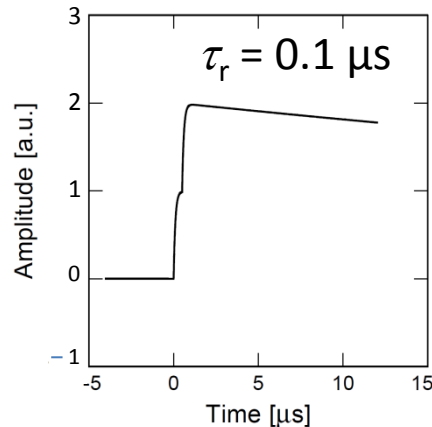
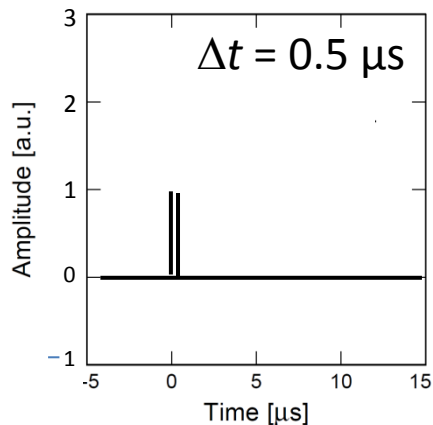
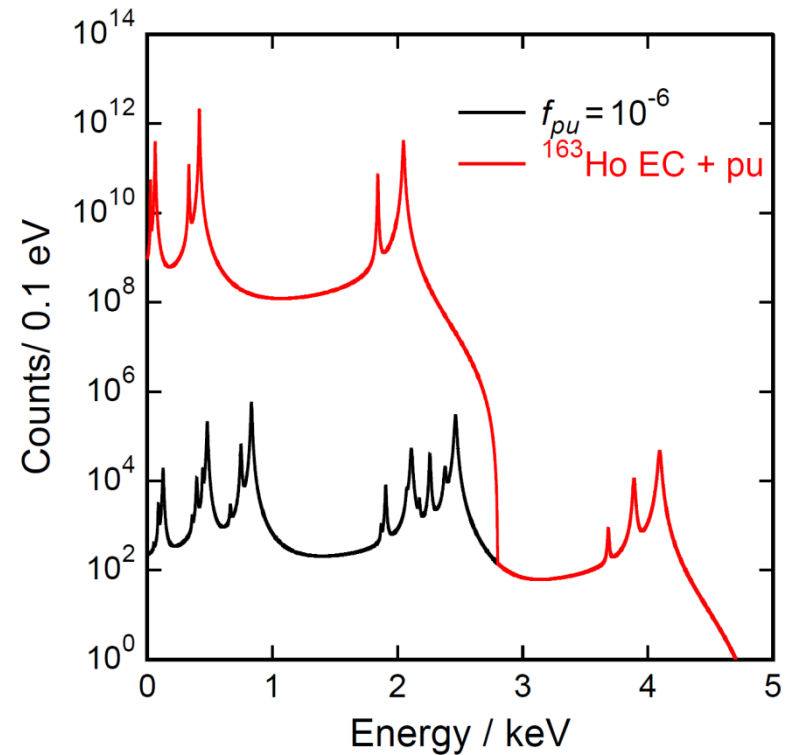
Requirements for sub-eV sensitivity in ECHO

Statistics in the end point region

- $N_{\text{ev}} > 10^{14} \rightarrow A \approx 1 \text{ MBq}$

Unresolved pile-up ($f_{\text{pu}} \sim a \cdot \tau_r$)

- $f_{\text{pu}} < 10^{-5}$
- $\tau_r < 1 \mu\text{s} \rightarrow a \sim 10 \text{ Bq}$
- 10^5 pixels



Requirements for sub-eV sensitivity in ECHO

Statistics in the end point region

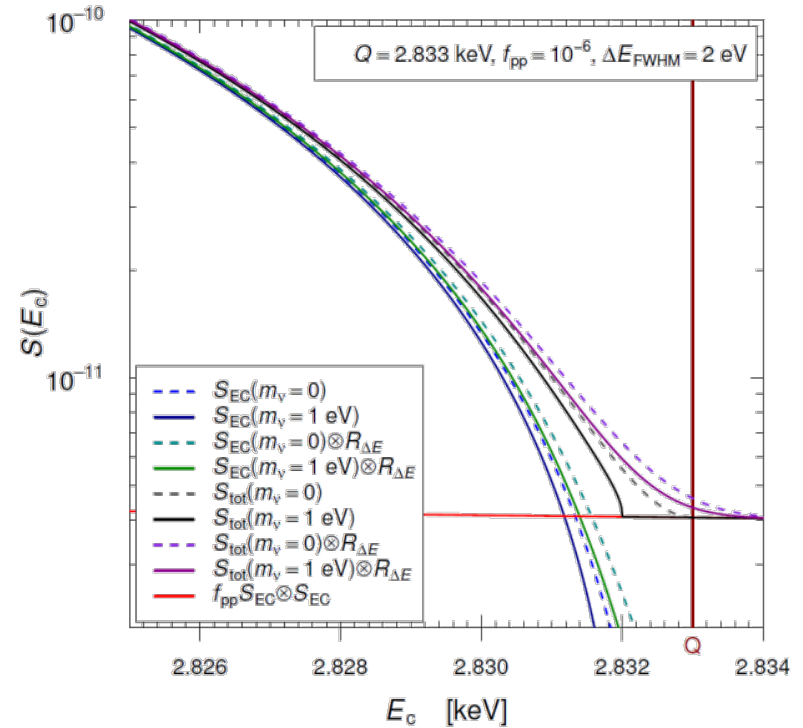
- $N_{\text{ev}} > 10^{14} \rightarrow A \approx 1 \text{ MBq}$

Unresolved pile-up ($f_{\text{pu}} \sim a \cdot \tau_r$)

- $f_{\text{pu}} < 10^{-5}$
- $\tau_r < 1 \mu\text{s} \rightarrow a \sim 10 \text{ Bq}$
- $10^5 \text{ pixels} \rightarrow \text{multiplexing}$

Precision characterization of the endpoint region

- $\Delta E_{\text{FWHM}} < 3 \text{ eV}$



Requirements for sub-eV sensitivity in ECHO

Statistics in the end point region

- $N_{\text{ev}} > 10^{14} \rightarrow A \approx 1 \text{ MBq}$

Unresolved pile-up ($f_{\text{pu}} \sim a \cdot \tau_r$)

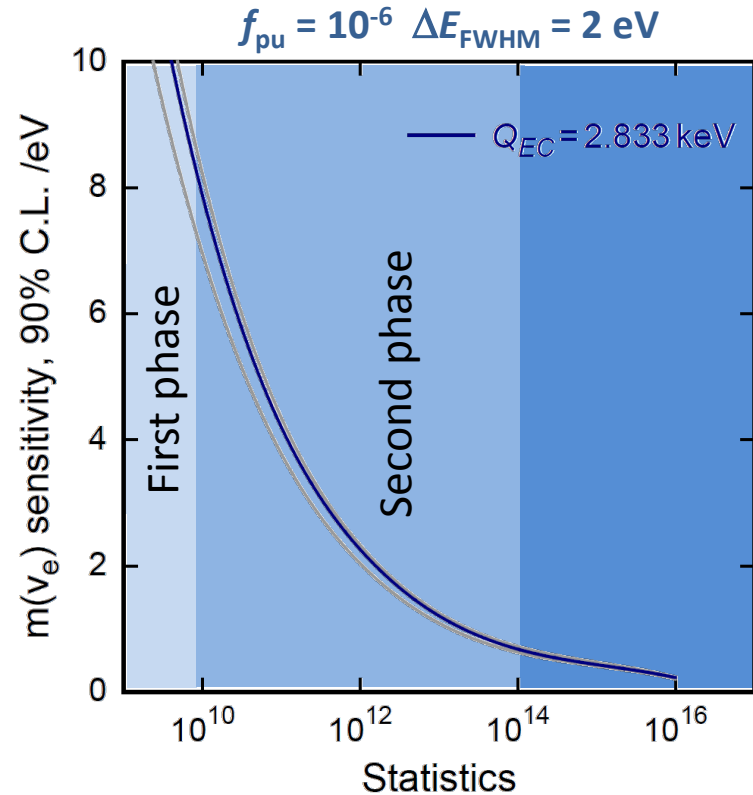
- $f_{\text{pu}} < 10^{-5}$
- $\tau_r < 1 \mu\text{s} \rightarrow a \sim 10 \text{ Bq}$
- $10^5 \text{ pixels} \rightarrow \text{multiplexing}$

Precision characterization of the endpoint region

- $\Delta E_{\text{FWHM}} < 3 \text{ eV}$

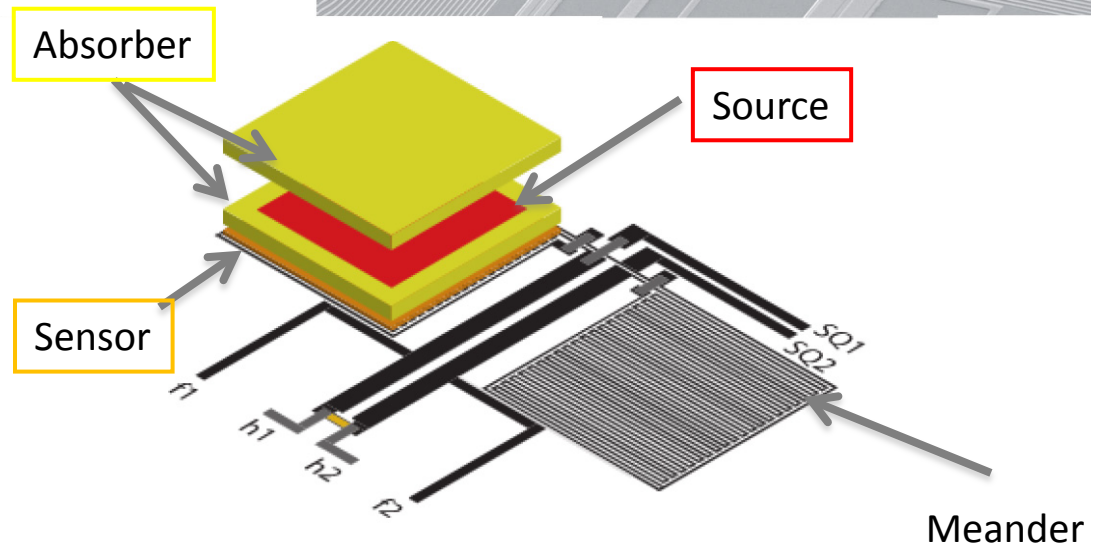
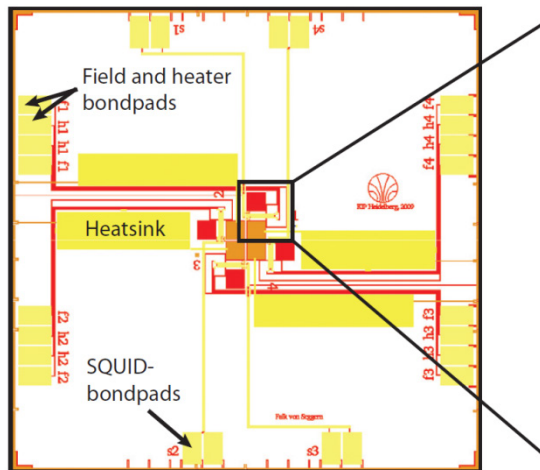
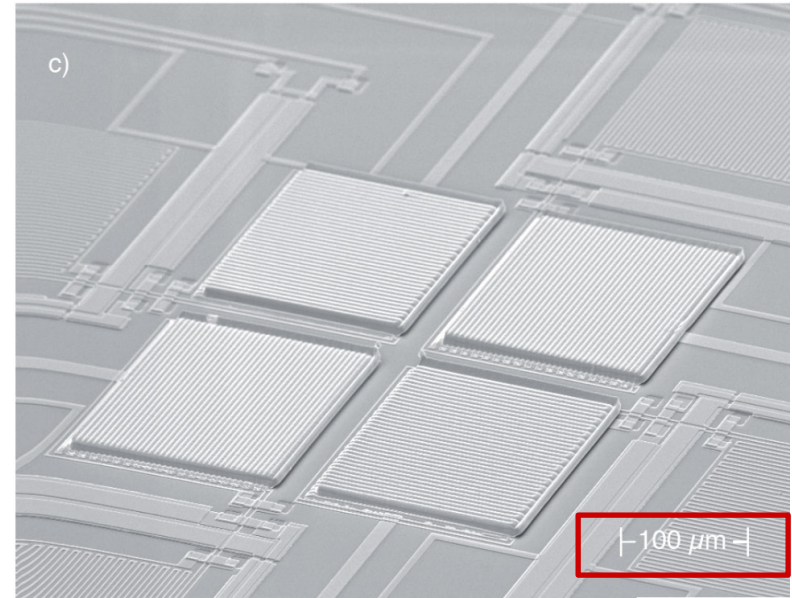
Background level

- $< 10^{-6} \text{ events/eV/det/day}$



First detector prototype for ^{163}Ho

- Absorber for calorimetric measurement
→ ion implantation @ ISOLDE-CERN in 2009
on-line process
- About 0.01 Bq per pixel
- Operated over more than 4 years

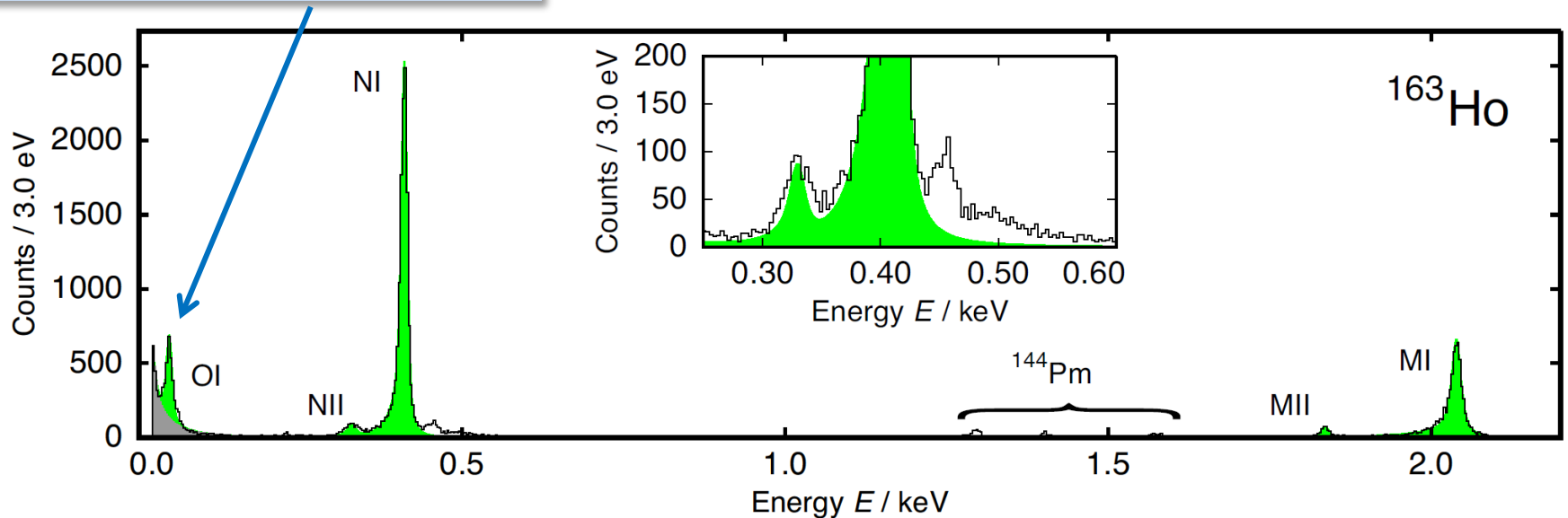


Calorimetric spectrum

- Rise Time ~ 130 ns
- $\Delta E_{\text{FWHM}} = 7.6$ eV @ 6 keV (2013)
- Non-Linearity $< 1\%$ @ 6keV

	E_{H} bind.	E_{H} exp.	Γ_{H} lit.	Γ_{H} exp
MI	2.047	2.040	13.2	13.7
MII	1.845	1.836	6.0	7.2
NI	0.420	0.411	5.4	5.3
NII	0.340	0.333	5.3	8.0
OI	0.050	0.048	5.0	4.3

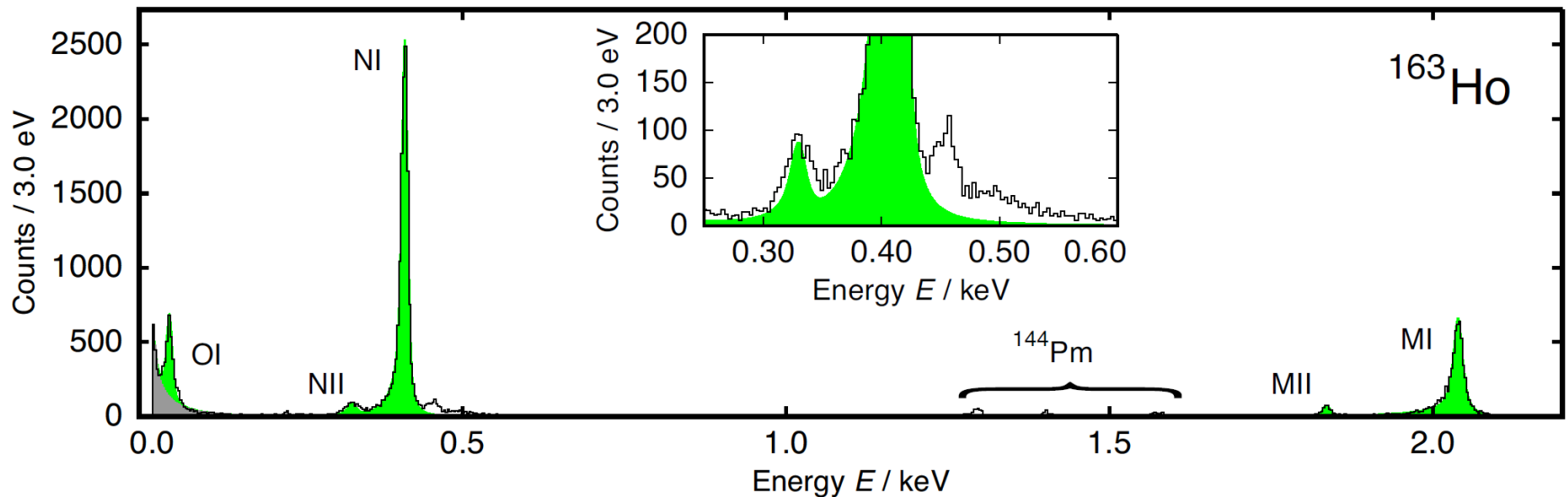
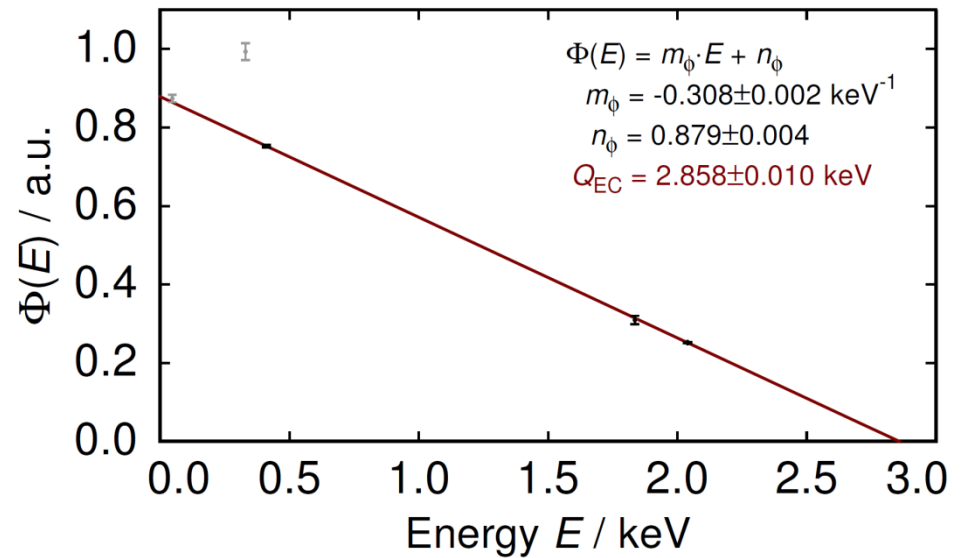
First calorimetric measurement of the OI-line



Q_{EC} determination

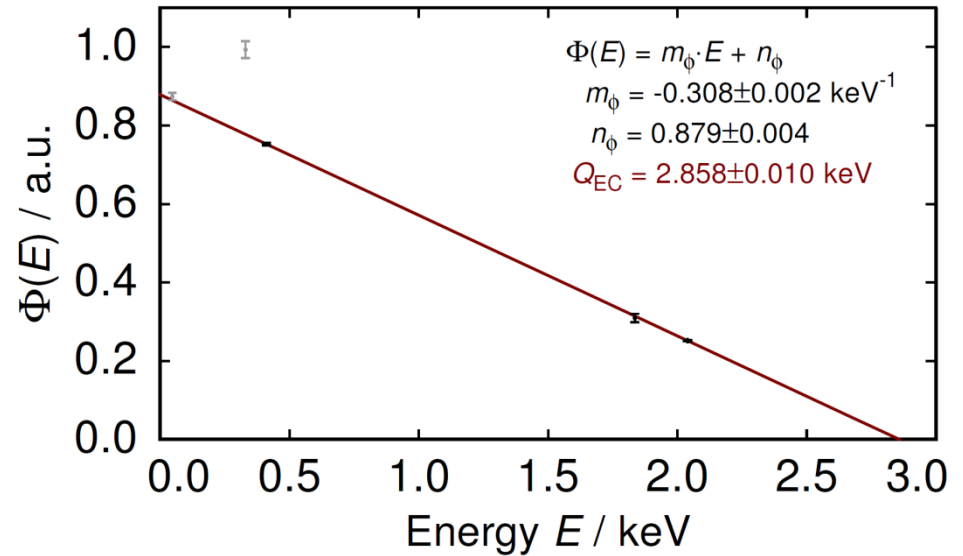
$$\Phi_H(E) = \sqrt{\frac{n_H}{\phi_H^2(0)B_H}} = \sqrt{C}(Q_{EC} - E_H)$$

Line amplitudes are affected
by the phase space factor



Q_{EC} determination

$$\Phi_H(E) = \sqrt{\frac{n_H}{\phi_H^2(0)B_H}} = \sqrt{C}(Q_{EC} - E_H)$$

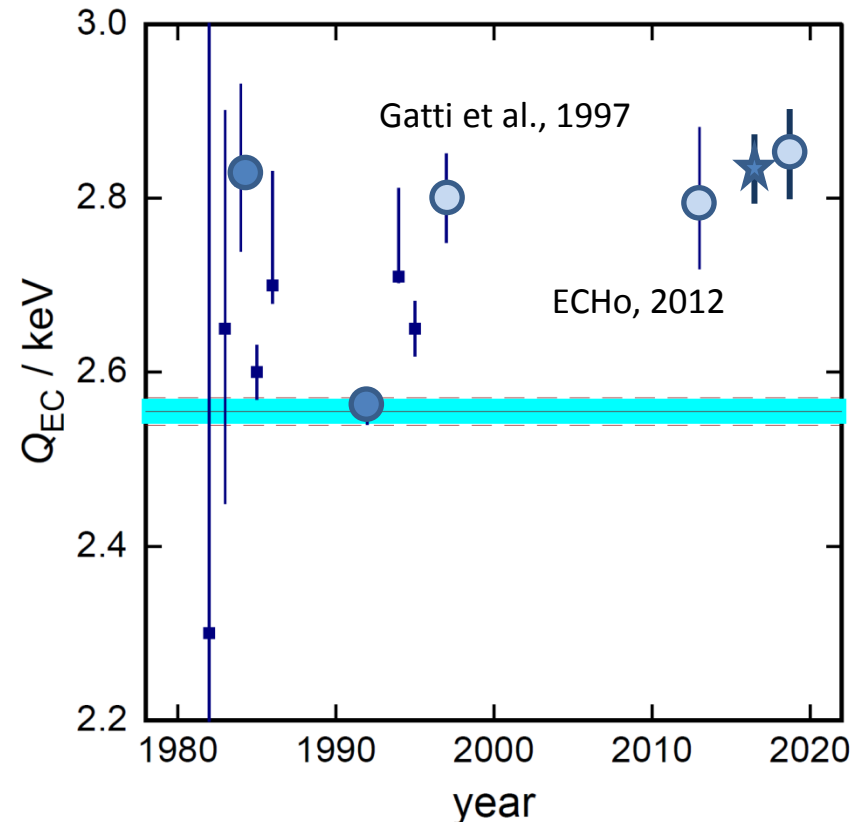


Our result:

$$Q_{EC} = (2.858 \pm 0.010^{\text{stat}} \pm 0.05^{\text{syst}}) \text{ keV}$$

Penning Trap Mass Spectrometry result:

$$Q_{EC} = (2.833 \pm 0.030^{\text{stat}} \pm 0.015^{\text{syst}}) \text{ keV}$$







Scaling up

^{163}Ho high purity source

Required activity in the detectors: Final experiment $\rightarrow >10^6 \text{ Bq} \rightarrow >10^{17}$ atoms

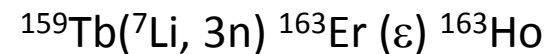
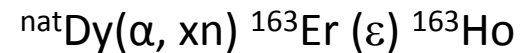
- Neutron irradiation
(n, γ)-reaction on ^{162}Er

High cross-section 

Radioactive contaminants 

Er161 3.21 h 3/2- EC	Er162 0+ 0.14	Er163 75.0 m 5/2 EC	Er164 0+ 1.61	Er165 10.36 h 5/2- EC	Er166 0+ 33.6
Ho160 25.6 m 5+ EC *	Ho161 2.48 h 7/2- EC *	Ho162 15.0 m 1+ EC *	Ho163 1.70 y 2- EC	Ho164 29 m 1+ EC, β^-	Ho165 2.3 y 3/2- EC *
Dy159 144.4 d 3/2- EC	Dy160 0+ 2.34	Dy161 5/2+ 18.9	Dy162 0+ 25.5	Dy163 5/2- 24.9	Dy164 0+ 28.2
Tb158 180 y 3- EC, β^- *	Tb159 3/2+ 100	Tb160 72.3 d 3- β^-	Tb161 6.88 d 3/2+ β^-	Tb162 7.60 m 1- β^-	Tb163 19.5 m 3/2+ β^-

- Charged particle activation



Small cross-section 

Few radioactive contaminants 

^{163}Ho high purity source

Required activity in the detectors: Final experiment $\rightarrow >10^6 \text{ Bq} \rightarrow >10^{17}$ atoms

- Neutron irradiation
(n,γ)-reaction on ^{162}Er

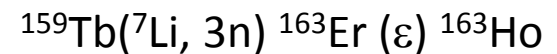
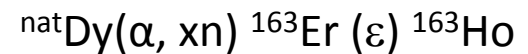
High cross-section 

Radioactive contaminants 



Er161 3.21 h 3/2- EC	Er162 0+ 0.14	Er163 75.0 m 5/2 EC	Er164 0+ 1.61	Er165 10.36 h 5/2- EC	Er166 0+ 33.6
Ho160 25.6 m 5+ EC *	Ho161 2.48 h 7/2- EC *	Ho162 15.0 m 1+ EC *	Ho163 1.70 y 2- EC	Ho164 29 m 1+ EC,β-	Ho165 2.3 y 3/2- β-
Dy159 144.4 d 3/2- EC	Dy160 0+ 2.34	Dy161 5/2+ 18.9	Dy162 0+ 25.5	Dy163 5/2- 24.9	Dy164 0+ 28.2
Tb158 180 y 3- EC,β- *	Tb159 3/2+ 100	Tb160 72.3 d 3- β-	Tb161 6.88 d 3/2+ β-	Tb162 7.60 m 1- β-	Tb163 19.5 m 3/2+ β-

- Charged particle activation

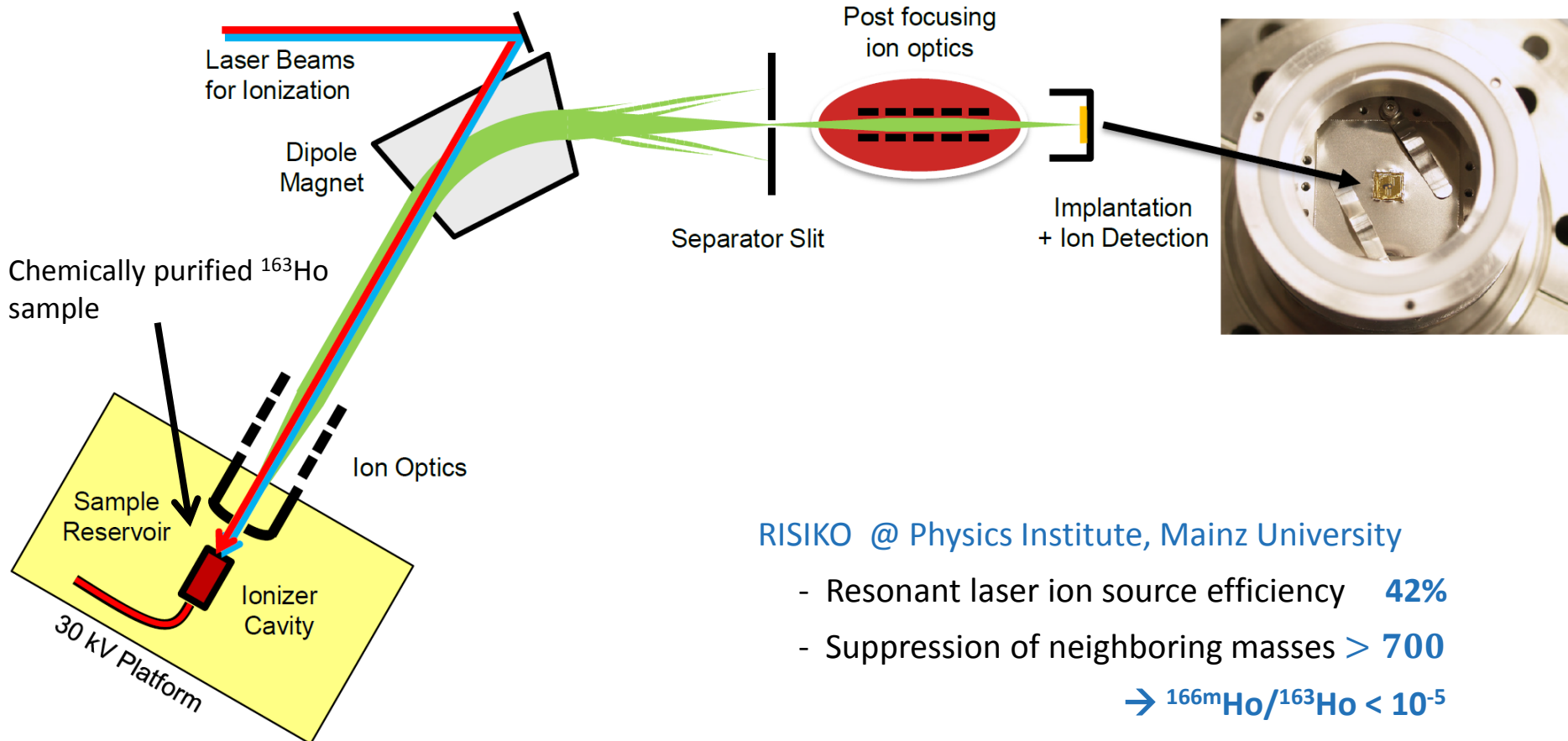


Small cross-section 

Few radioactive contaminants 

NuMECS

Mass separation and ^{163}Ho ion-implantation

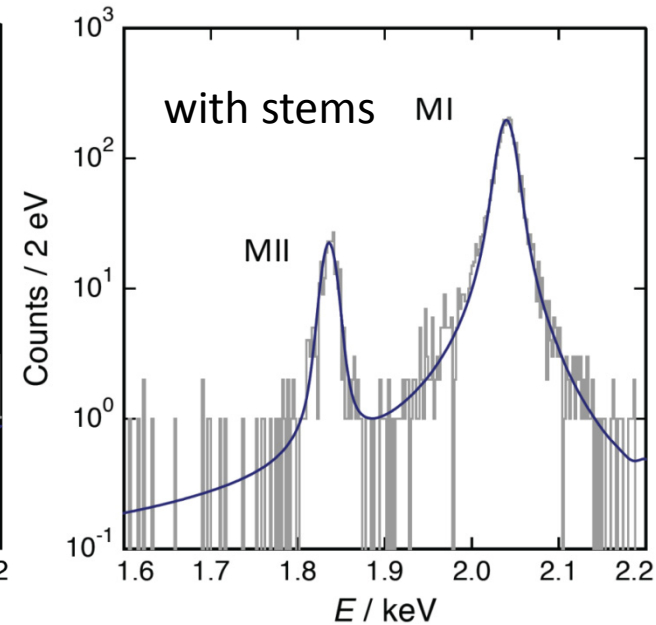
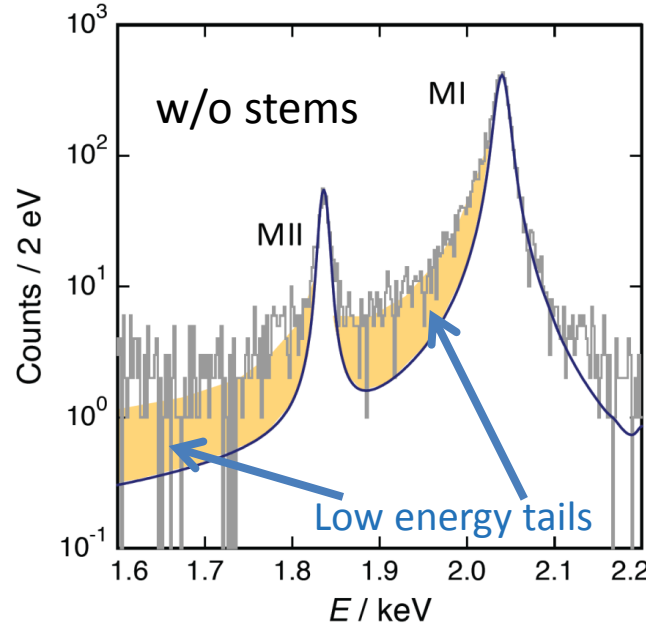
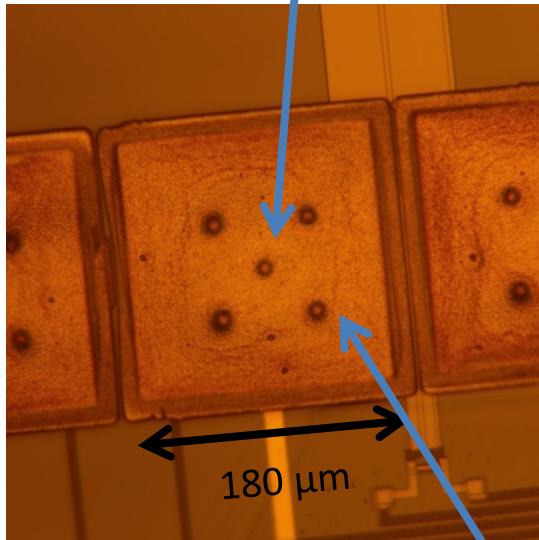


RISIKO @ Physics Institute, Mainz University

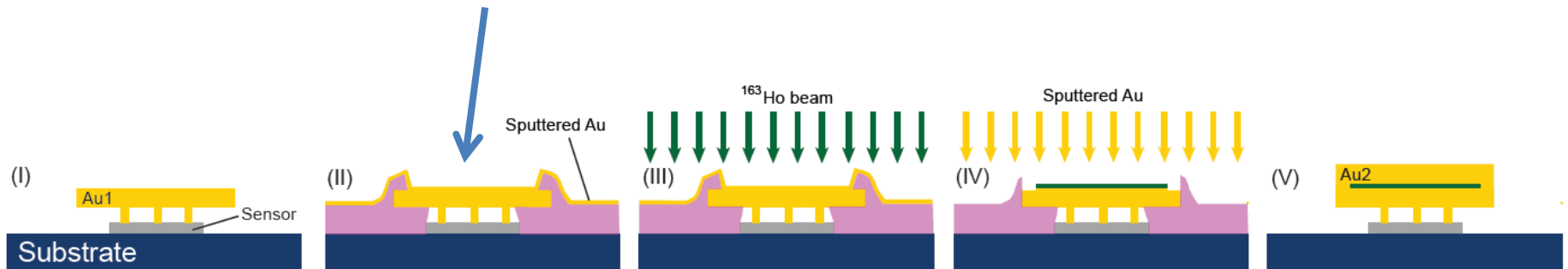
- Resonant laser ion source efficiency **42%**
- Suppression of neighboring masses **> 700**
→ $^{166}\text{mHo}/^{163}\text{Ho} < 10^{-5}$
- Optimization of beam focalization

Fabrication 4π absorber

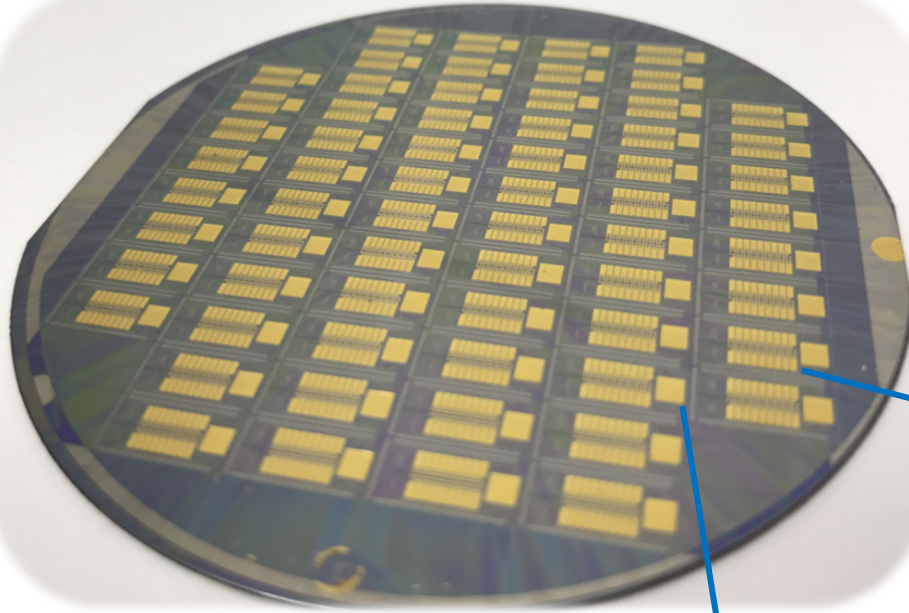
Stems between absorber and sensor prevent athermal phonon loss to the substrate



Definition of the implantation area by microstructuring a photoresist layer



ECHO-1k array



3" wafer with 64 ECHO-1k chip

Suitable for
parallel and multiplexed readout

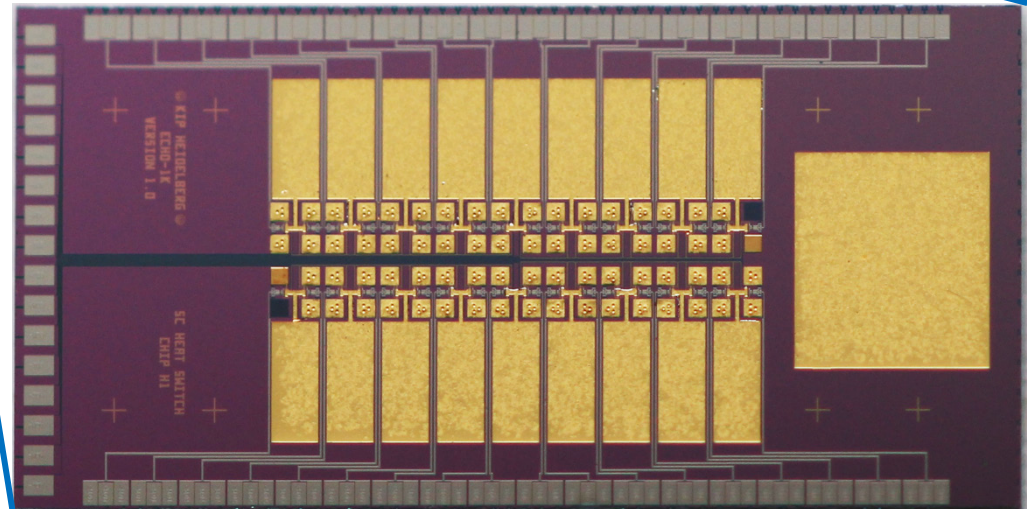
64 pixels which can be loaded with ^{163}Ho
+ 4 detectors for diagnostics

Design performance:

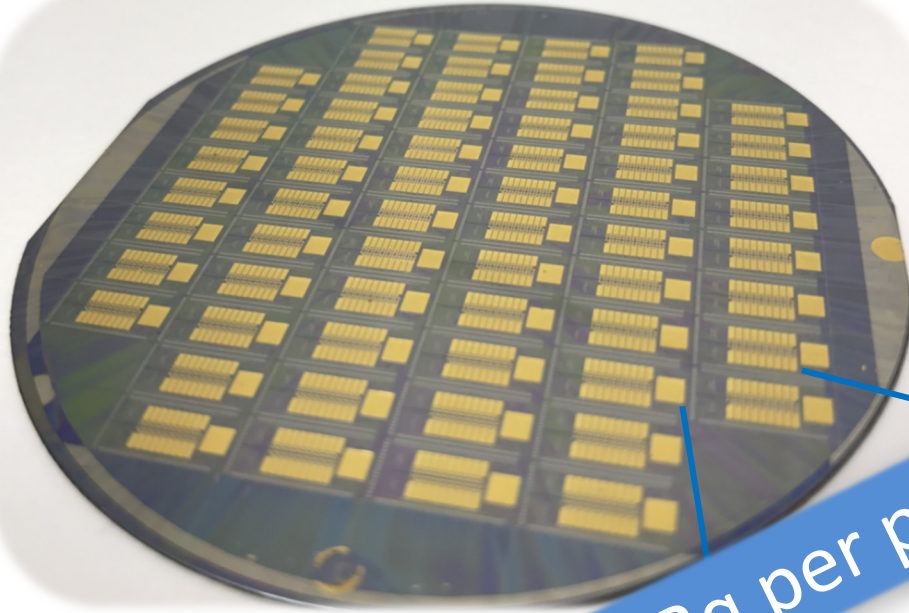
$$\Delta E_{\text{FWHM}} \sim 5 \text{ eV}$$

$$\tau_r \sim 90 \text{ ns (single channel readout)}$$

$$\tau_r \sim 300 \text{ ns (multiplexed read-out)}$$



ECHO-1k array



3" wafer with 64 ECHO-1k chip

Suitable for parallel and multiplexed read-out

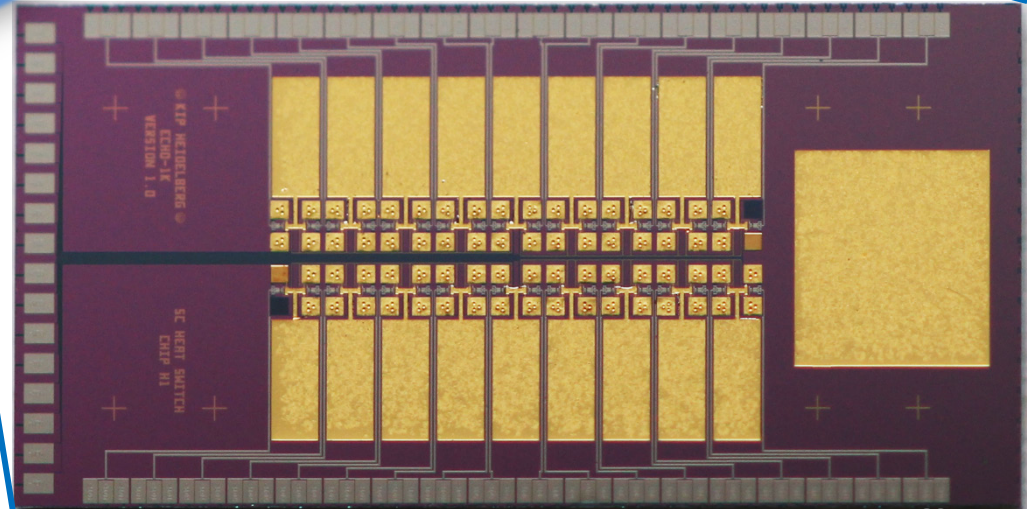
64 pixels which can be read out with ^{163}Ho
+ 4 detector channels

Design performance:

$$\Delta E_{\text{FWHM}} \sim 5 \text{ eV}$$

$$\tau_r \sim 90 \text{ ns (single channel readout)}$$

$$\tau_r \sim 300 \text{ ns (multiplexed read-out)}$$



ECHo-1k (2015 - 2018)

^{163}Ho activity: $A_t = 1 \text{ kBq}$

Detectors: **Metallic Magnetic Calorimeters**

→ Energy resolution $\Delta E_{\text{FWHM}} \leq 5 \text{ eV}$

→ Time resolution $\tau \leq 1 \mu\text{s}$

Unresolved pile-up fraction $f_{\text{pu}} \leq 10^{-5}$

→ activity per pixel: $A = 10 \text{ Bq}$

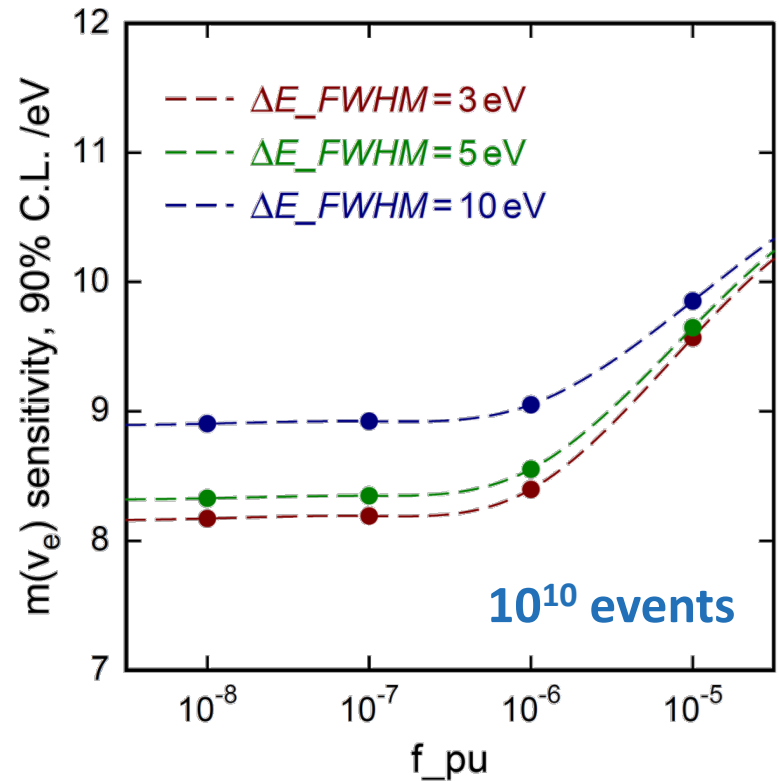
→ number of detectors $N = 100$

Read-out : **Microwave SQUID Multiplexing**

→ 2 arrays with ~ 50 single pixels

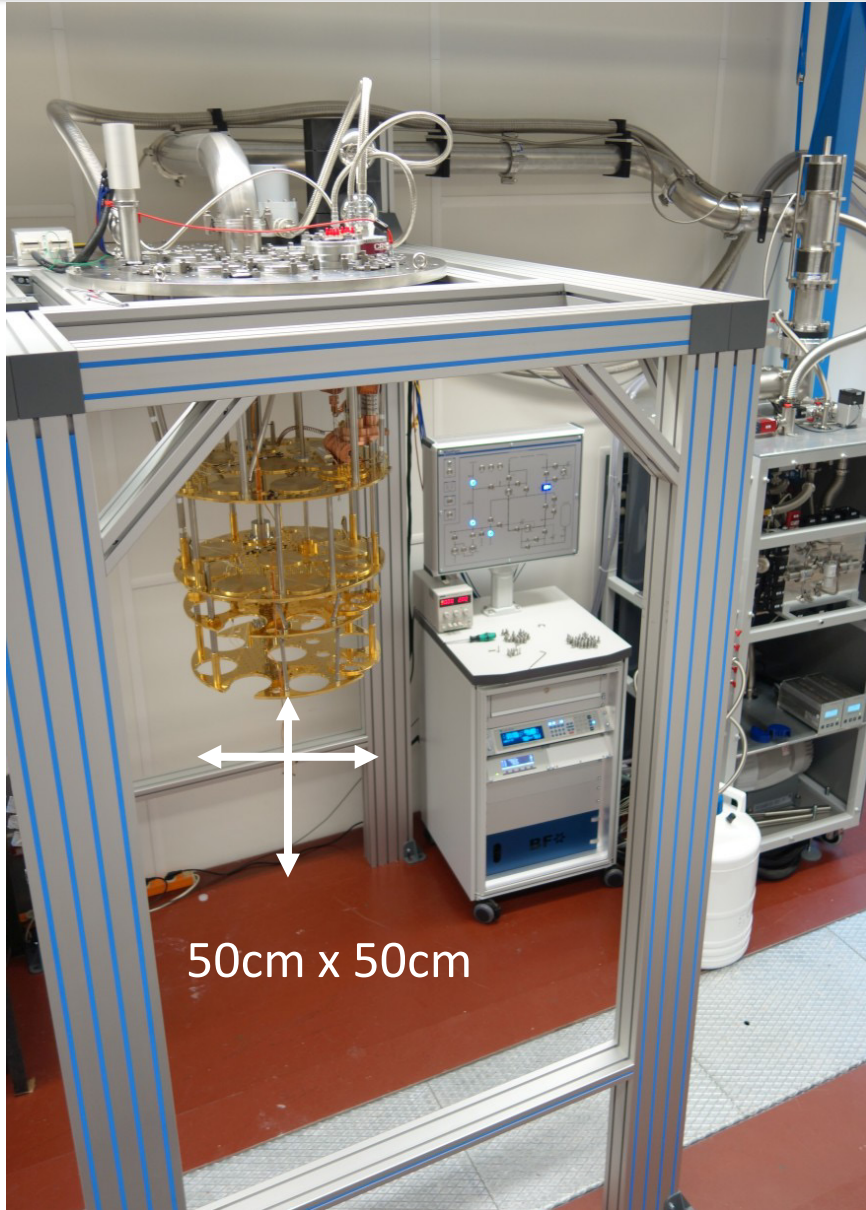
Background $b < 10^{-5} \text{ /eV/det/day}$

Measuring time $t = 1 \text{ year}$

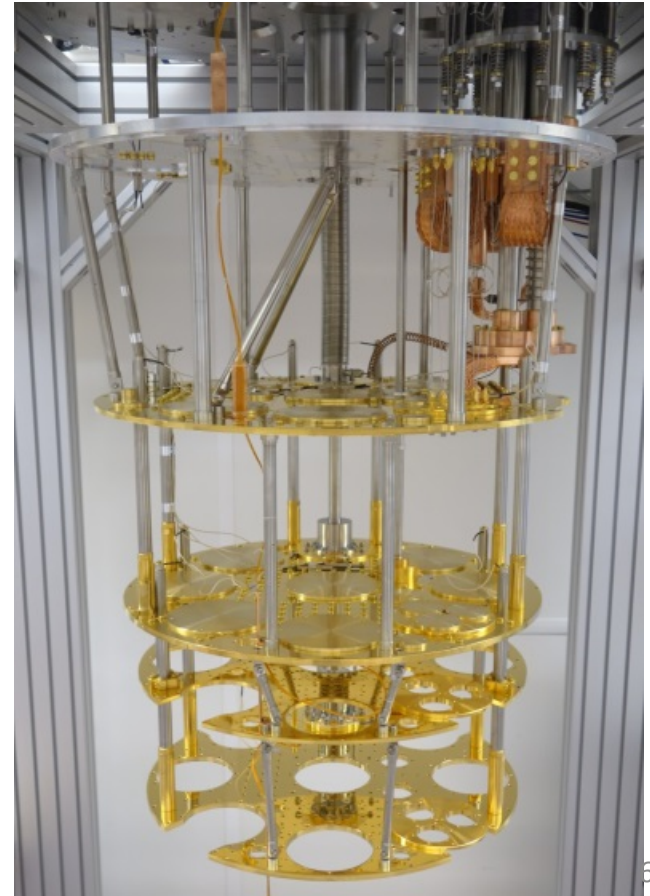


$m(\nu_e) < 10 \text{ eV } 90\% \text{ C.L.}$

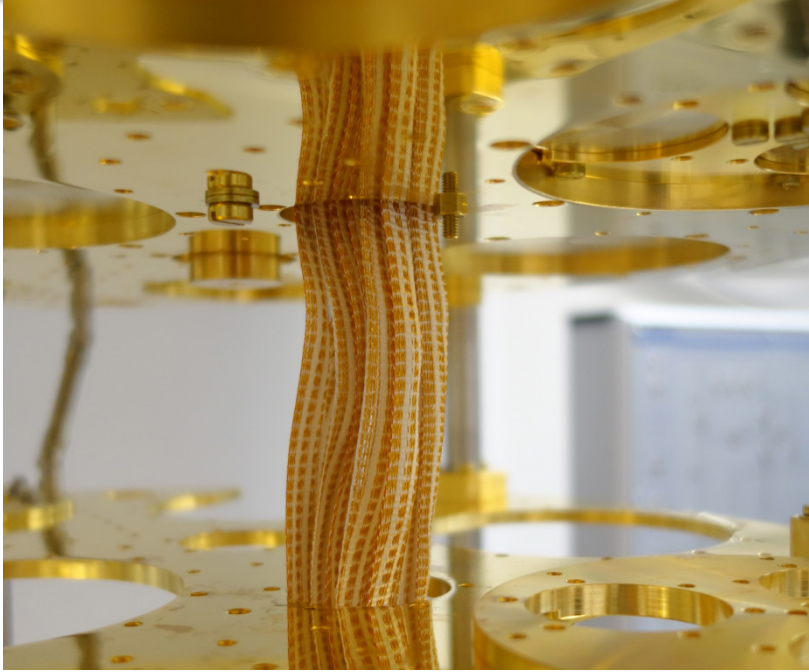
ECHO cryogenic platform



- Large space at MXC enough for several ECHO phases
- cooling power: $15\mu\text{W}$ @ 20 mK
- Possibility to load 200kg for passive shielding



ECHo cryogenic platform

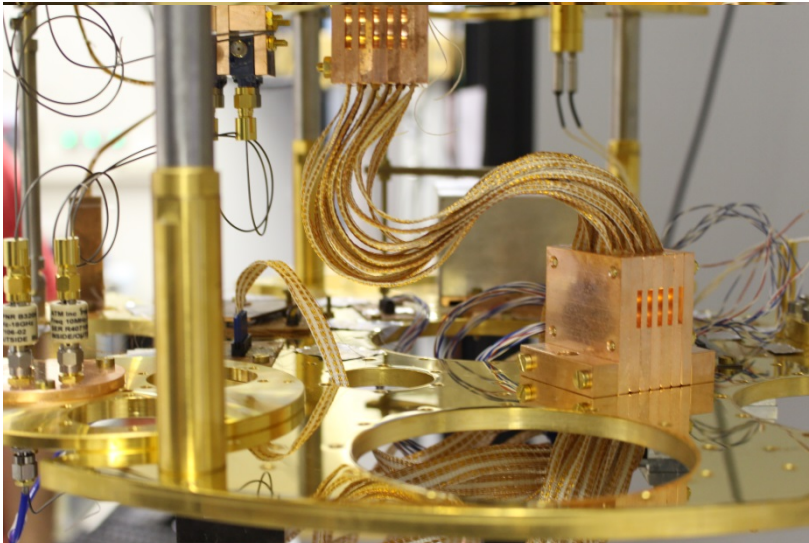


- Large space at MXC enough for several ECHo phases
- cooling power: $15\mu\text{W}$ @ 20 mK
- Possibility to load 200kg for passive shielding
- Presently equipped with:

2 RF lines for microwave multiplexing readout of 2 MMC arrays

12 ribbons each with 30 Cu98Ni2 0.2 mm, 1.56 Ohm/m, cables from RT to mK

→ allows for parallel readout of 36 two-stage SQUID set-up



ECHo-1M (next future)

^{163}Ho activity: $A_t = 1 \text{ MBq}$

Detectors: **Metallic Magnetic Calorimeters**

→ Energy resolution $\Delta E_{\text{FWHM}} \leq 3 \text{ eV}$

→ Time resolution $\tau \leq 0.1 \mu\text{s}$

Unresolved pile-up fraction $f_{\text{pu}} \leq 10^{-6}$

→ activity per pixel: $A = 10 \text{ Bq}$

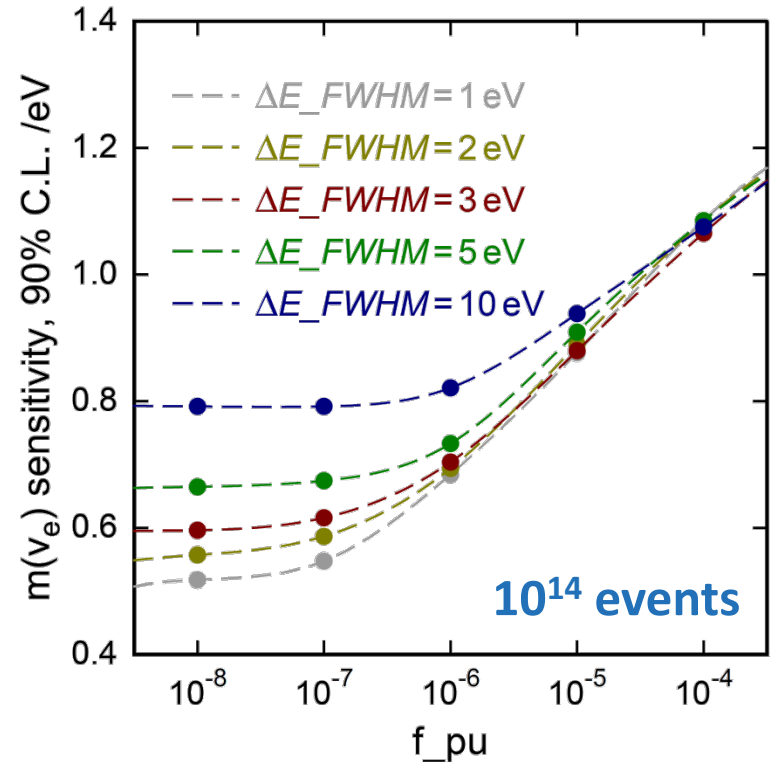
→ number of detectors $N = 10^5$

Read-out : **Microwave SQUID Multiplexing**

→ 100 arrays with ~ 1000 single pixels

Background $b < 10^{-6} \text{ /eV/det/day}$

Measuring time $t = 1 - 3 \text{ year}$

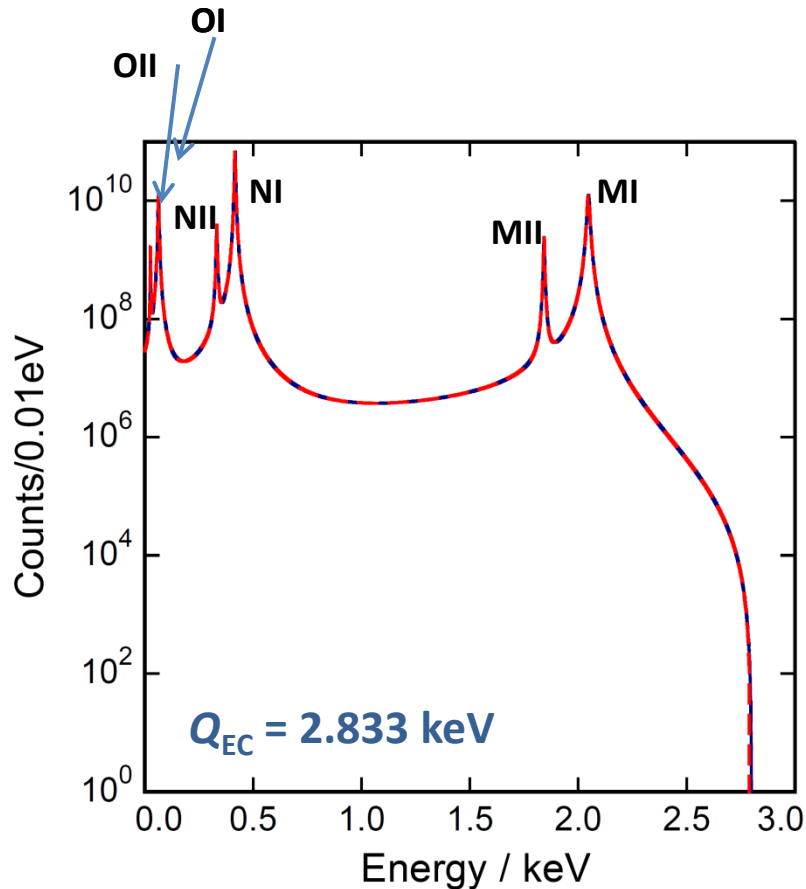


$m(\nu_e) < 1 \text{ eV } 90\% \text{ C.L.}$

How does
the existence of sterile neutrino
affect the EC spectrum?

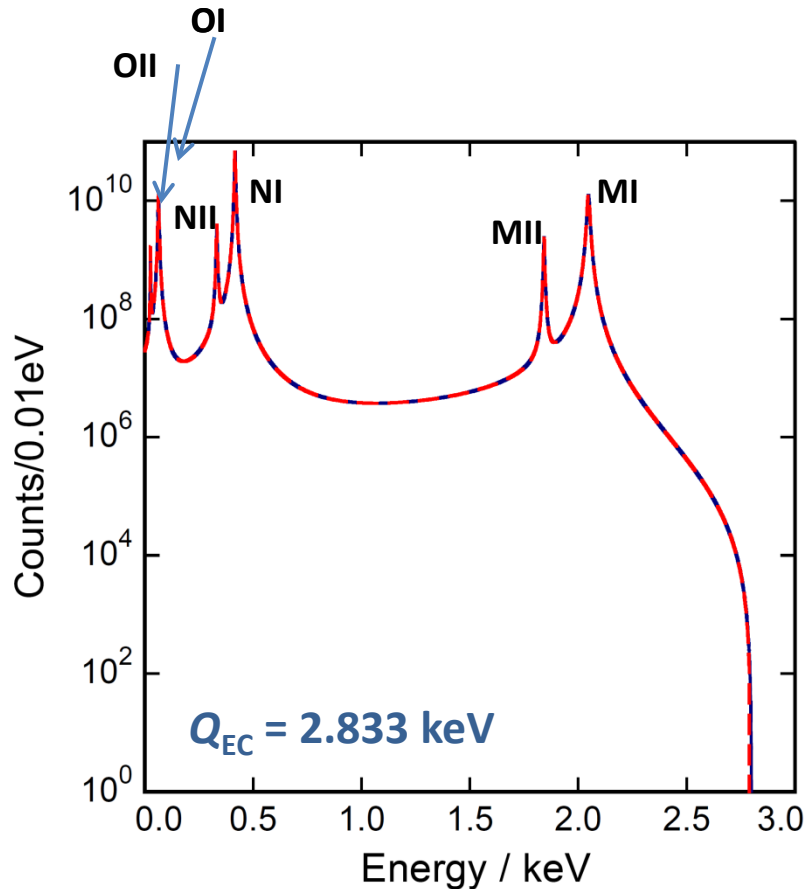
Sterile Neutrino and ^{163}Ho

$$\frac{dW}{dE_C} = A(Q_{\text{EC}} - E_C)^2 \sqrt{1 - \frac{m_\nu^2}{(Q_{\text{EC}} - E_C)^2}} \sum_H B_H \phi_H^2(0) \frac{\frac{\Gamma_H}{2\pi}}{(E_C - E_H)^2 + \frac{\Gamma_H^2}{4}}$$



Sterile Neutrino and ^{163}Ho

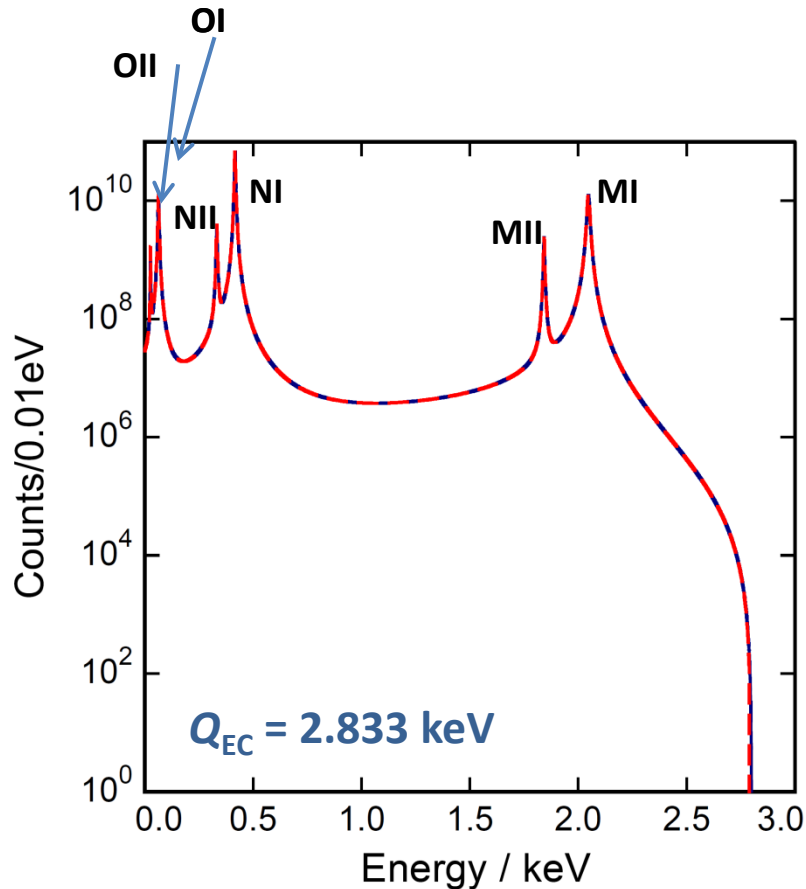
$$\frac{dW}{dE_C} = A(Q_{\text{EC}} - E_C)^2 \sum_i |U_{ei}|^2 \sqrt{1 - \frac{m_i^2}{(Q_{\text{EC}} - E_C)^2}} \sum_H B_H \varphi_H^2(0) \frac{\frac{\Gamma_H}{2\pi}}{(E_C - E_H)^2 + \frac{\Gamma_H^2}{4}}$$



- Electron neutrino as superposition of mass eigenstates

Sterile Neutrino and ^{163}Ho

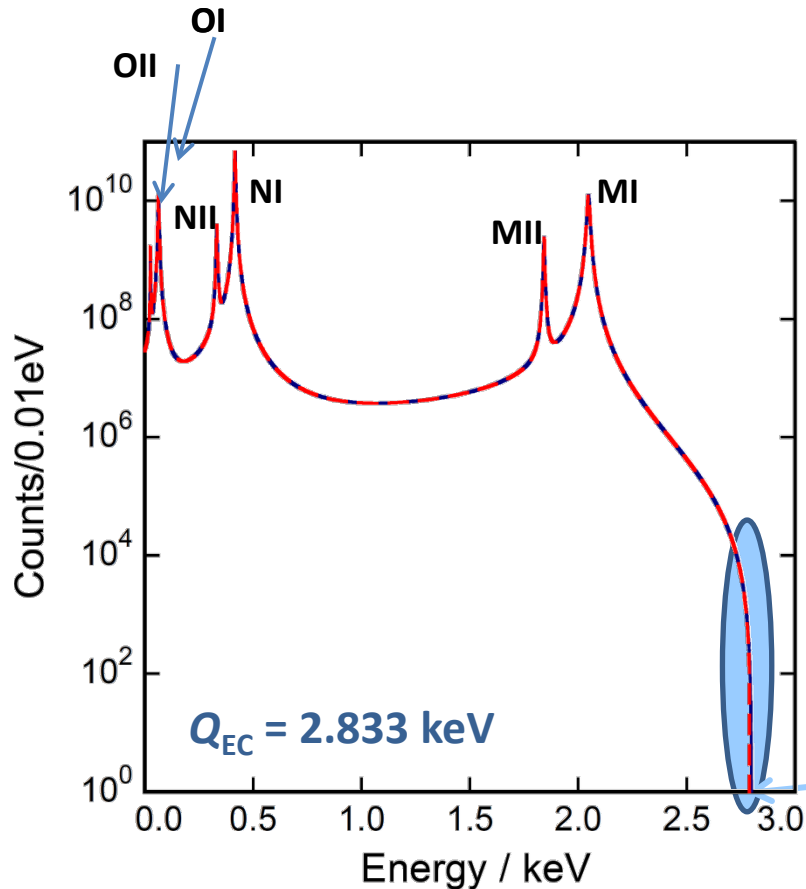
$$\frac{dW}{dE_C} = A(Q_{\text{EC}} - E_C)^2 \left[(1 - |U_{e4}|^2) + |U_{e4}|^2 \sqrt{1 - \frac{m_4^2}{(Q_{\text{EC}} - E_C)^2}} H(Q_{\text{EC}} - E_C - m_4) \right] \sum_H B_H \varphi_H^2(0) \frac{\frac{\Gamma_H}{2\pi}}{(E_C - E_H)^2 + \frac{\Gamma_H^2}{4}}$$



- Electron neutrino as superposition of mass eigenstates
- $m_{i=1,2,3} \ll m_4 \longrightarrow m_{i=1,2,3} \sim 0 \text{ eV}$

Sterile Neutrino and ^{163}Ho

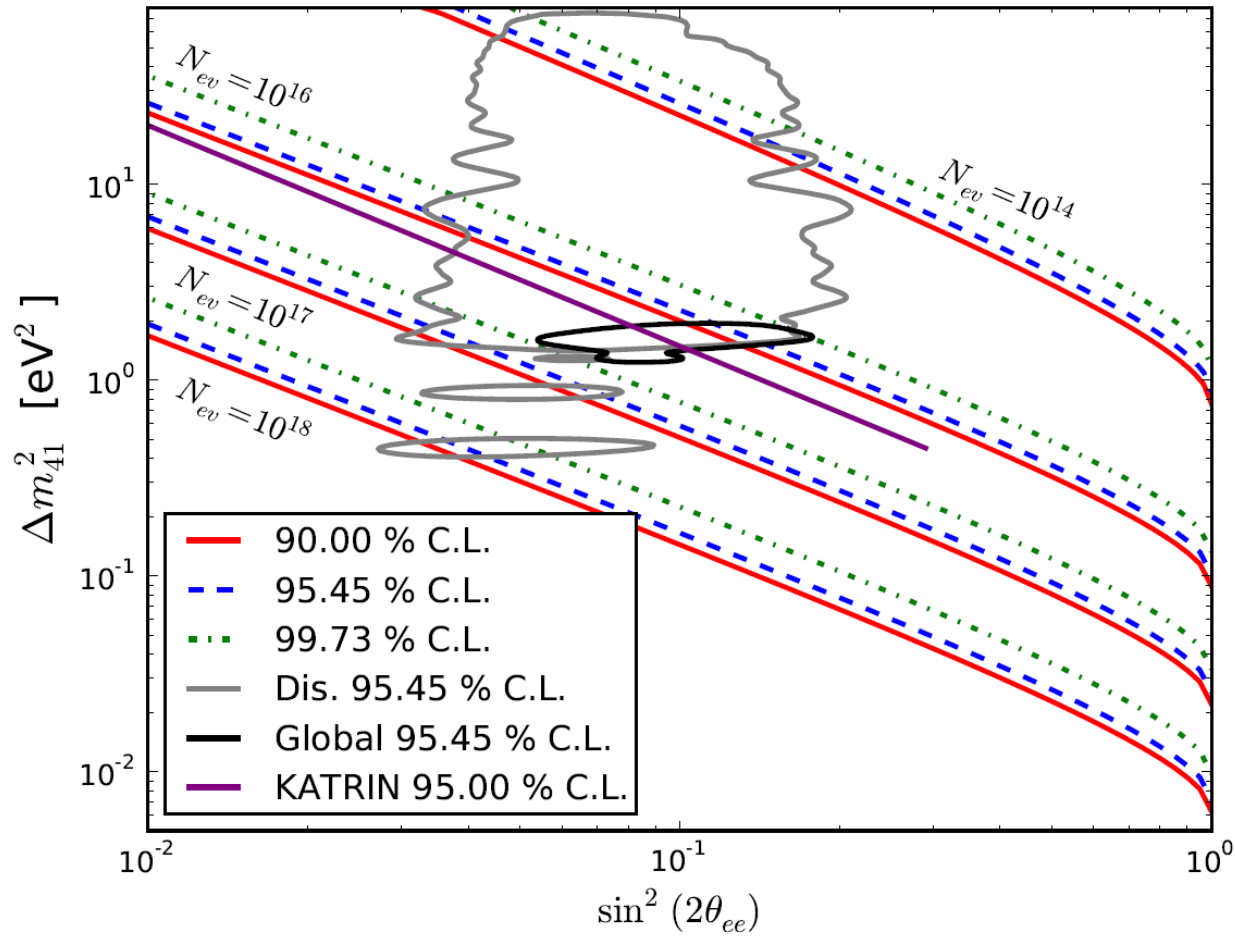
$$\frac{dW}{dE_C} = A(Q_{EC} - E_C)^2 \left[(1 - |U_{e4}|^2) + |U_{e4}|^2 \sqrt{1 - \frac{m_4^2}{(Q_{EC} - E_C)^2}} H(Q_{EC} - E_C - m_4) \right] \sum_H B_H \varphi_H^2(0) \frac{\frac{\Gamma_H}{2\pi}}{(E_C - E_H)^2 + \frac{\Gamma_H^2}{4}}$$



- Electron neutrino as superposition of mass eigenstates
- $m_{i=1,2,3} \ll m_4 \longrightarrow m_{i=1,2,3} \sim 0 \text{ eV}$

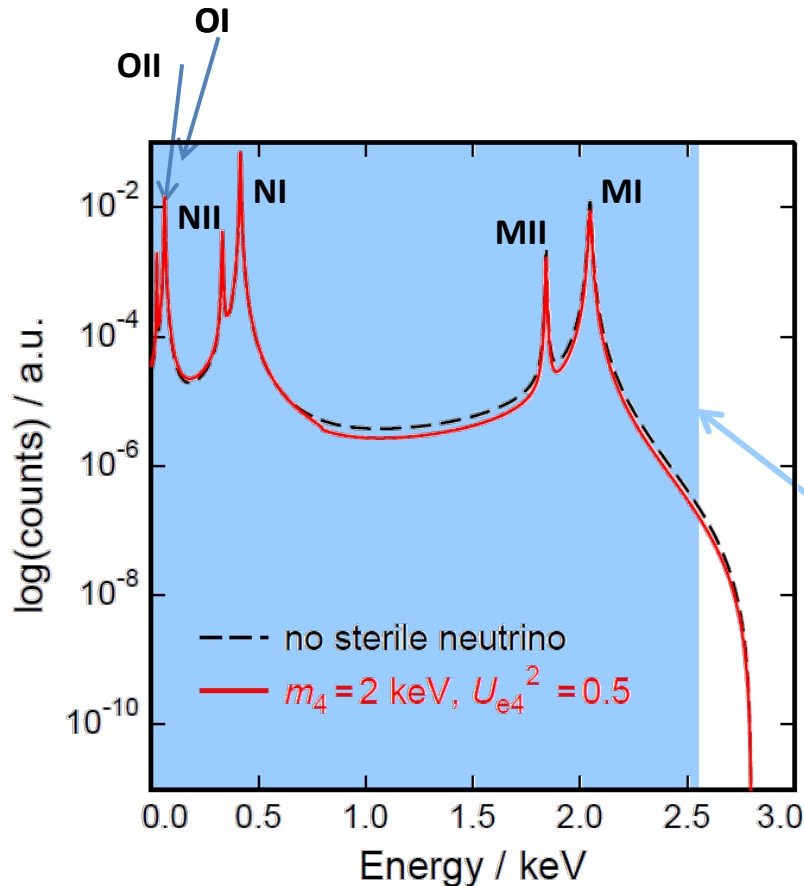
eV-scale sterile neutrinos

eV-scale sterile neutrino



keV-scale sterile neutrino

$$\frac{dW}{dE_C} = A(Q_{EC} - E_C)^2 \left[(1 - |U_{e4}|^2) + |U_{e4}|^2 \sqrt{1 - \frac{m_4^2}{(Q_{EC} - E_C)^2}} H(Q_{EC} - E_C - m_4) \right] \sum_H B_H \varphi_H^2(0) \frac{\frac{\Gamma_H}{2\pi}}{(E_C - E_H)^2 + \frac{\Gamma_H^2}{4}}$$



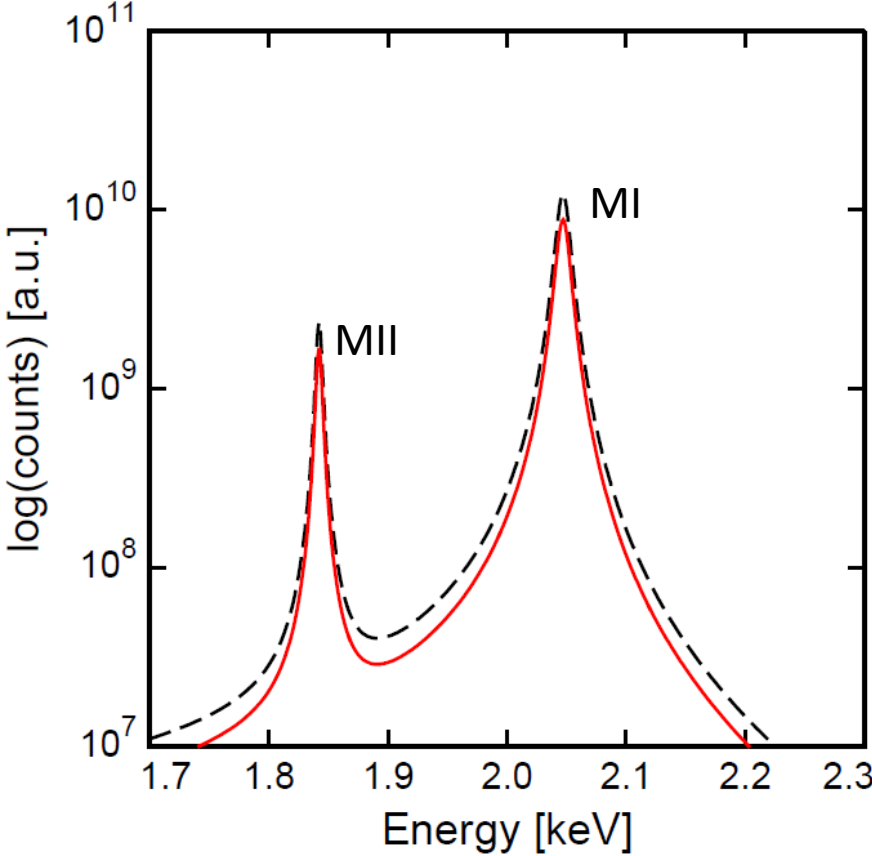
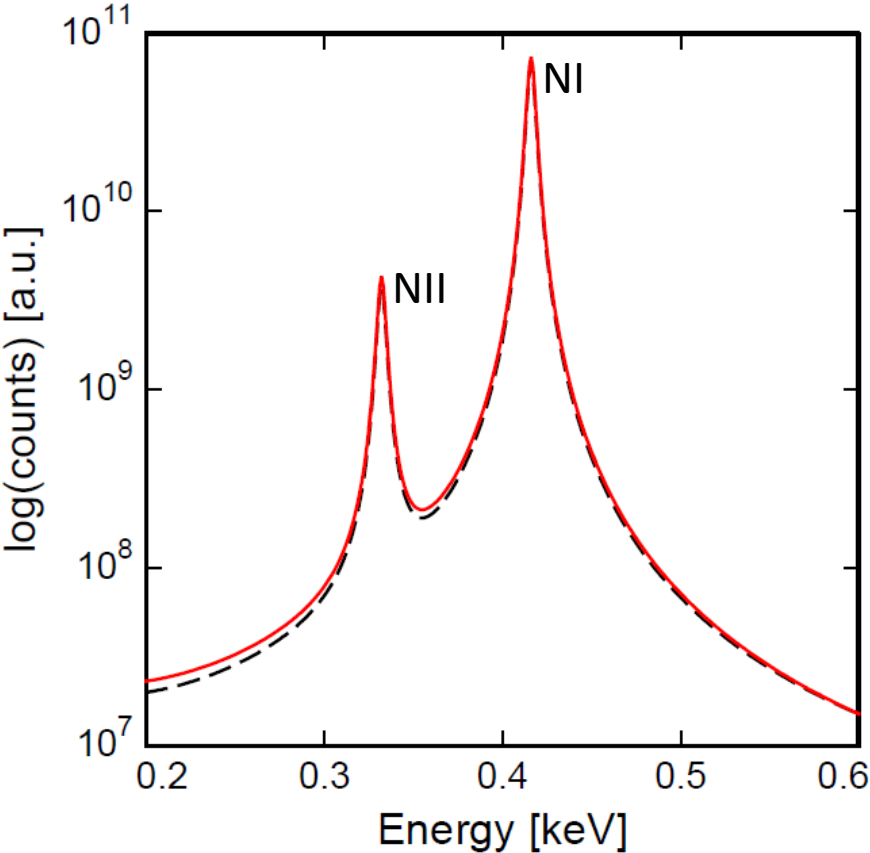
- Electron neutrino as superposition of mass eigenstates
- $m_{i=1,2,3} \ll m_4 \rightarrow m_{i=1,2,3} \sim 0 \text{ eV}$

keV-scale sterile neutrinos

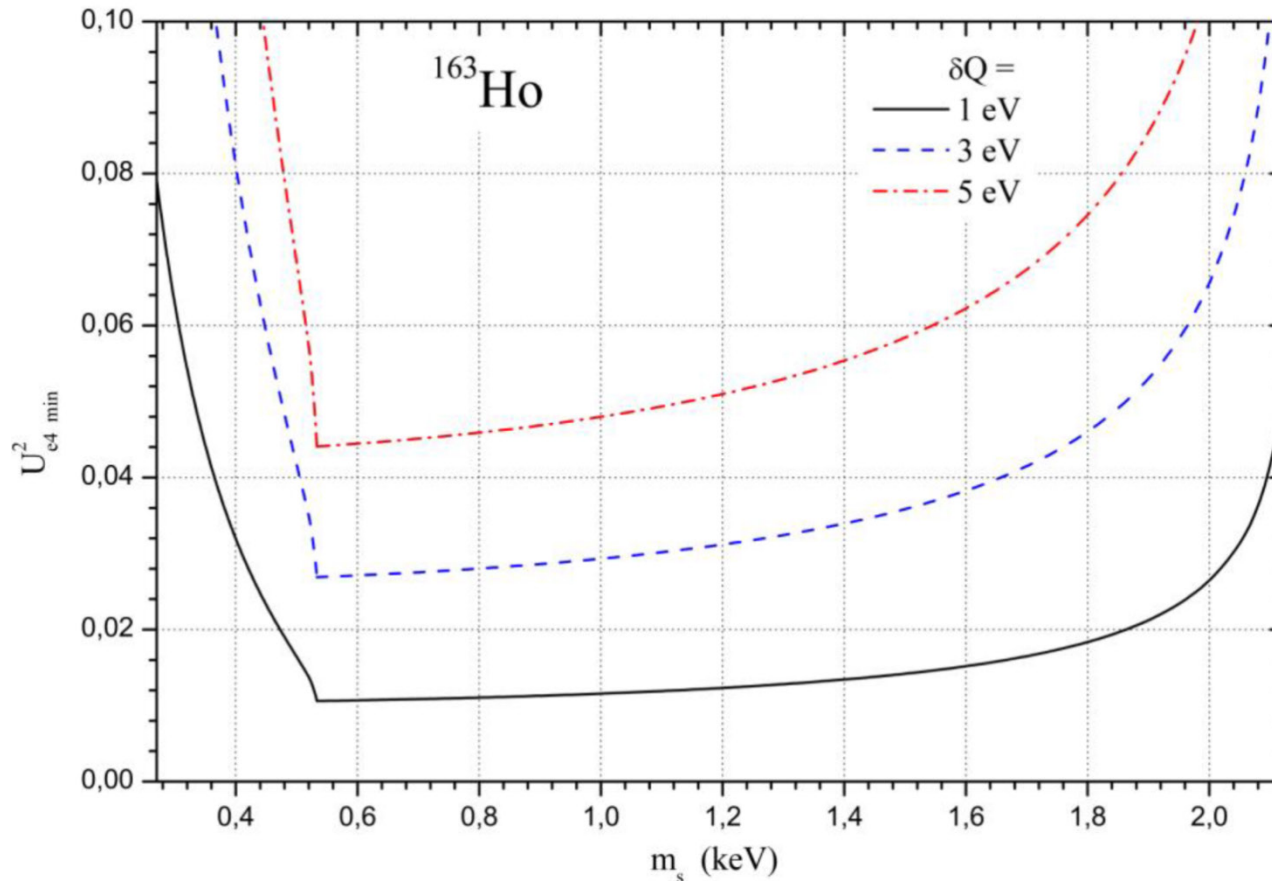
keV-scale sterile neutrino

$m_4=2 \text{ keV}, U_{e4}^2=0.5$

no sterile neutrino

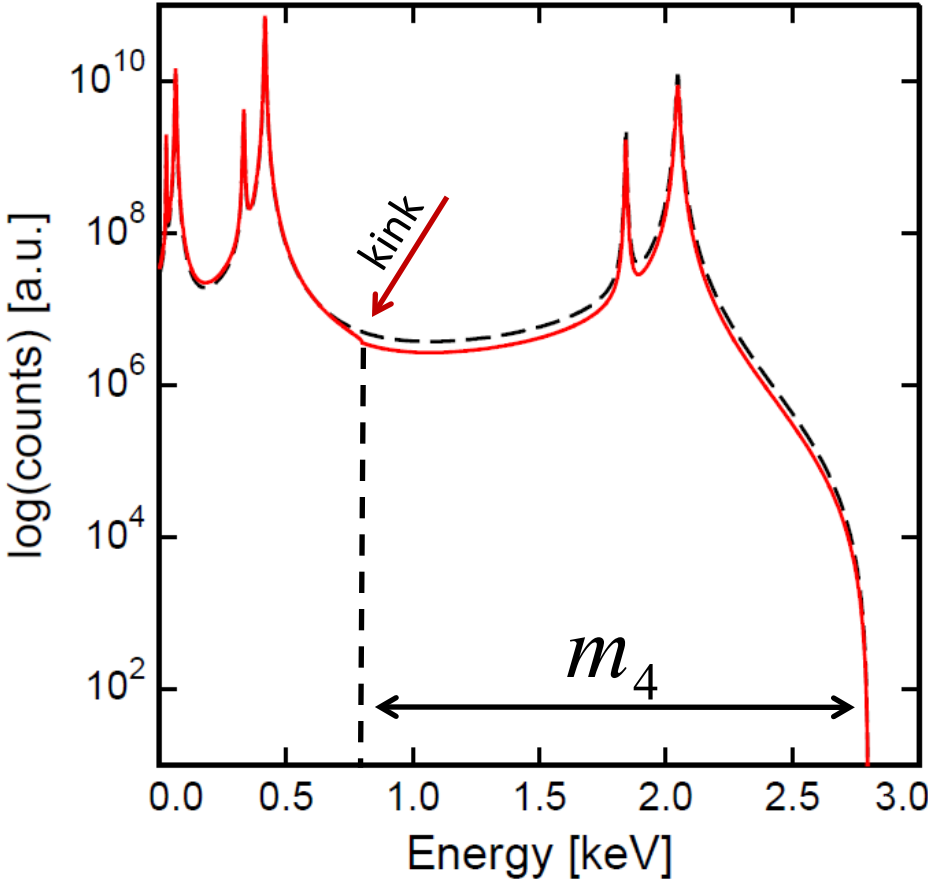


keV-scale sterile neutrino



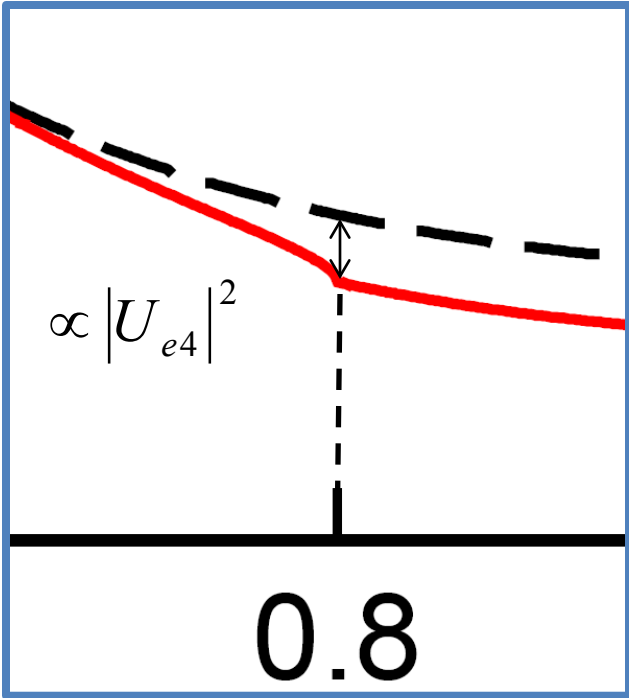
Sensitivity to the mixing matrix element at 90% CL as a function of the sterile neutrino mass achievable with about 10^{10} events in the full EC spectrum.

keV-scale sterile neutrino

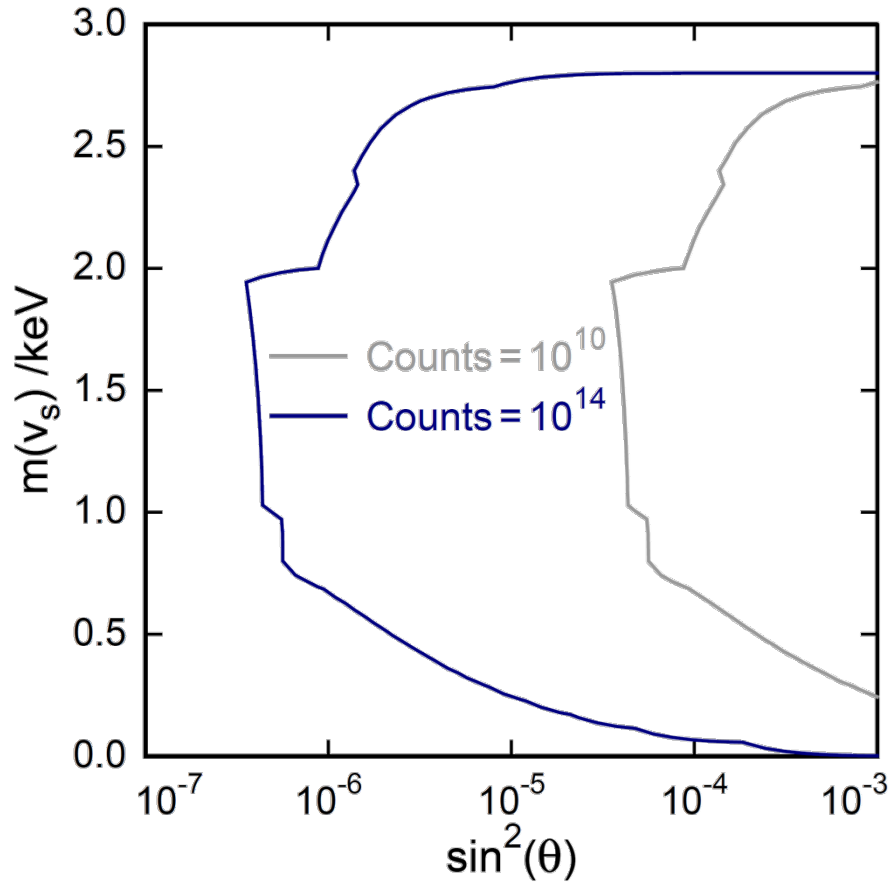


$m_4=2 \text{ keV}, U_{e4}^2=0.5$
no sterile neutrino

- position of kink => m_4
- depth of kink => $|U_{e4}|^2$



keV-scale sterile neutrino



- Statistical Fluctuation
- No Pile Up
- Theoretical Spectrum supposed to be perfectly known

Sterile Neutrino (keV) and Electron Capture

Other candidates in the EC branch:

- $Q_{\text{EC}} < 100 \text{ keV}$
- Reasonable halflife

Nuclide	$T_{1/2}$	EC-transition	Q (keV) [22]	B_i (keV) [23]	B_j (keV) [23]	$ \Psi_i ^2/ \Psi_j ^2$	$Q-B_i$ (keV)
^{123}Te	$>2 \cdot 10^{15} \text{ y}$?	52.7(16)	K: 30.4912(3)	L _I : 4.9392(3)	7.833	22.2
^{157}Tb	71 y	$3/2^+ \rightarrow 3/2^-$	60.04(30)	K: 50.2391(5)	L _I : 8.3756(5)	7.124	9.76
^{163}Ho	4570 y	$7/2^- \rightarrow 5/2^-$	2.555(16)	M _I : 2.0468(5)	N _I : 0.4163(5)	4.151	0.51
^{179}Ta	1.82 y	$7/2^+ \rightarrow 9/2^+$	105.6(4)	K: 65.3508(6)	L _I : 11.2707(4)	6.711	40.2
^{193}Pt	50 y	$1/2^- \rightarrow 3/2^+$	56.63(30)	L _I : 13.4185(3)	M _I : 3.1737(17)	4.077	43.2
^{202}Pb	52 ky	$0^+ \rightarrow 2^-$	46(14)	L _I : 15.3467(4)	M _I : 3.7041(4)	4.036	30.7
^{205}Pb	13 My	$5/2^- \rightarrow 1/2^+$	50.6(5)	L _I : 15.3467(4)	M _I : 3.7041(4)	4.036	35.3
^{235}Np	396 d	$5/2^+ \rightarrow 7/2^-$	124.2(9)	K: 115.6061(16)	L _I : 21.7574(3)	5.587	8.6

Sterile Neutrino (keV) and Electron Capture

Other candidates in the EC branch:

- $Q_{\text{EC}} < 100 \text{ keV}$
- Reasonable halflife

Nuclide	$T_{1/2}$	EC-transition	Q (keV) [22]	B_i (keV) [23]	B_j (keV) [23]	$ \Psi_i ^2/ \Psi_j ^2$	$Q-B_i$ (keV)
^{123}Te	$>2 \cdot 10^{15} \text{ y}$?	52.7(16)	K: 30.4912(3)	L _I : 4.9392(3)	7.833	22.2
^{157}Tb	71 y	$3/2^+ \rightarrow 3/2^-$	60.04(30)	K: 50.2391(5)	L _I : 8.3756(5)	7.124	9.76
^{163}Ho	4570 y	$7/2^- \rightarrow 5/2^-$	2.555(16)	M _I : 2.0468(5)	N _I : 0.4163(5)	4.151	0.51
^{179}Ta	1.82 y	$7/2^+ \rightarrow 9/2^+$	105.6(4)	K: 65.3508(6)	L _I : 11.2707(4)	6.711	40.2
^{193}Pt	50 y	$1/2^- \rightarrow 3/2^+$	56.63(30)	L _I : 13.4185(3)	M _I : 3.1737(17)	4.077	43.2
^{202}Pb	52 ky	$0^+ \rightarrow 2^-$	46(14)	L _I : 15.3467(4)	M _I : 3.7041(4)	4.036	30.7
^{205}Pb	13 My	$5/2^- \rightarrow 1/2^+$	50.6(5)	L _I : 15.3467(4)	M _I : 3.7041(4)	4.036	35.3
^{235}Np	396 d	$5/2^+ \rightarrow 7/2^-$	124.2(9)	K: 115.6061(16)	L _I : 21.7574(3)	5.587	8.6

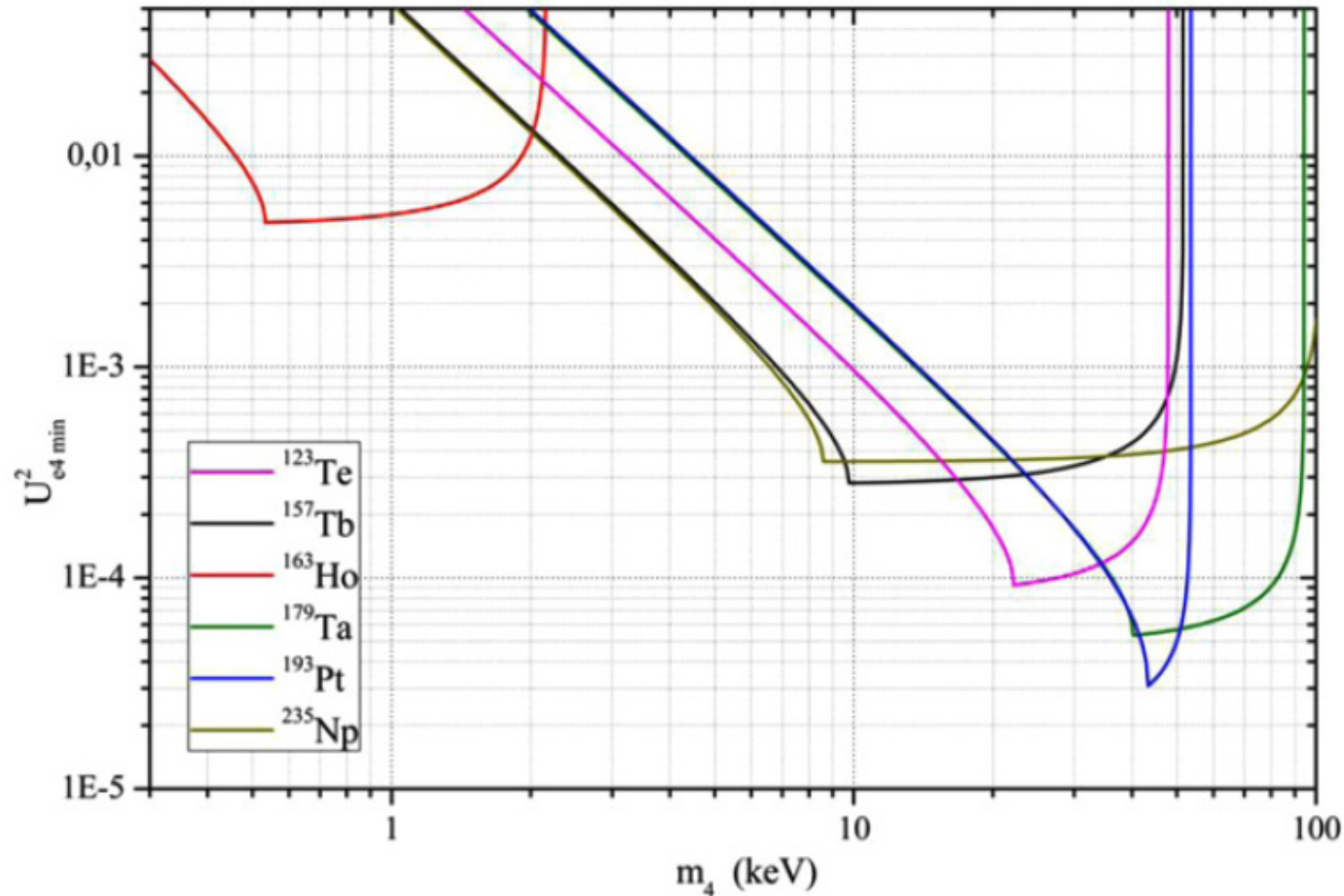
Sterile Neutrino (keV) and Electron Capture

Other candidates in the EC branch:

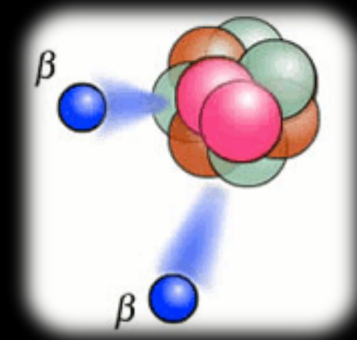
- $Q_{\text{EC}} < 100 \text{ keV}$
- Reasonable halflife

Nuclide	$T_{1/2}$	EC-transition	Q (keV) [22]	B_i (keV) [23]	B_j (keV) [23]	$ \Psi_i ^2/ \Psi_j ^2$	$Q-B_i$ (keV)
^{123}Te	$>2 \cdot 10^{15} \text{ y}$?	52.7(16)	K: 30.4912(3)	L _I : 4.9392(3)	7.833	22.2
^{157}Tb	71 y	$3/2^+ \rightarrow 3/2^-$	60.04(30)	K: 50.2391(5)	L _I : 8.3756(5)	7.124	9.76
^{163}Ho	4570 y	$7/2^- \rightarrow 5/2^-$	2.555(16)	M _I : 2.0468(5)	N _I : 0.4163(5)	4.151	0.51
^{179}Ta	1.82 y	$7/2^+ \rightarrow 9/2^+$	105.6(4)	K: 65.3508(6)	L _I : 11.2707(4)	6.711	40.2
^{193}Pt	50 y	$1/2^- \rightarrow 3/2^+$	56.63(30)	L _I : 13.4185(3)	M _I : 3.1737(17)	4.077	43.2
^{202}Pb	52 ky	$0^+ \rightarrow 2^-$	46(14)	L _I : 15.3467(4)	M _I : 3.7041(4)	4.036	30.7
^{205}Pb	13 My	$5/2^- \rightarrow 1/2^+$	50.6(5)	L _I : 15.3467(4)	M _I : 3.7041(4)	4.036	35.3
^{235}Np	396 d	$5/2^+ \rightarrow 7/2^-$	124.2(9)	K: 115.6061(16)	L _I : 21.7574(3)	5.587	8.6

Sterile Neutrino (keV) and Electron Capture



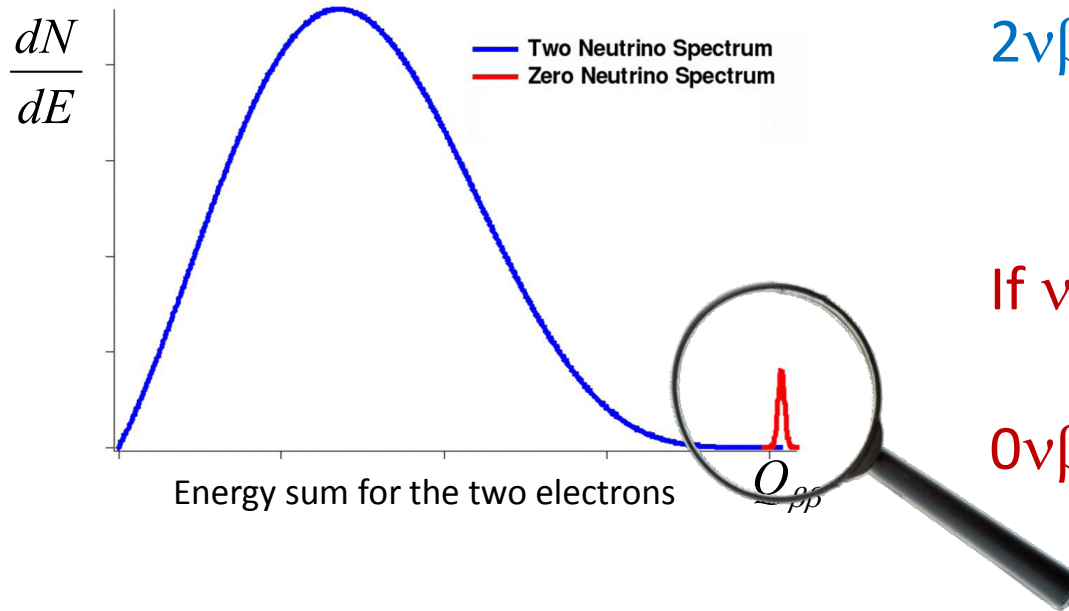
Same statistics + including errors : $(\delta \psi_{ij} = 0)$ $\delta Q_{EC} = 1 \text{ eV}$ $\delta E_{i,j} = 0.1 \text{ eV}$.



Neutrinoless double beta decay

Neutrinoless double beta decay

If conservation rules don't allow simple beta decay:



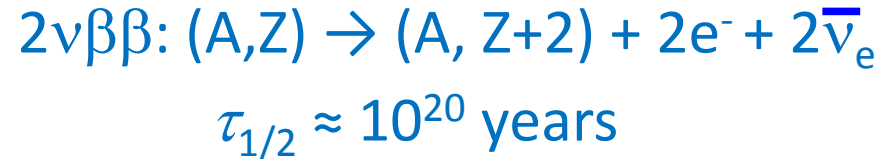
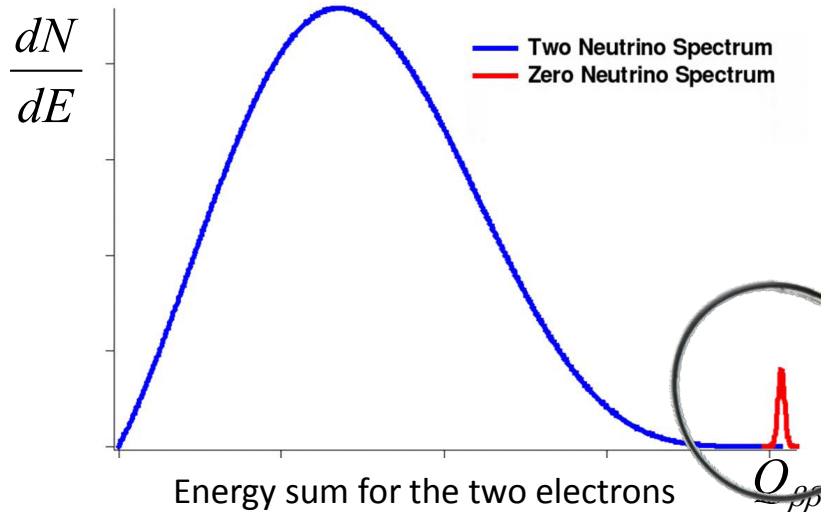
$$2\nu\beta\beta: (A,Z) \rightarrow (A, Z+2) + 2e^- + 2\bar{\nu}_e$$
$$\tau_{1/2} \approx 10^{20} \text{ years}$$

If $\nu = \bar{\nu}$:

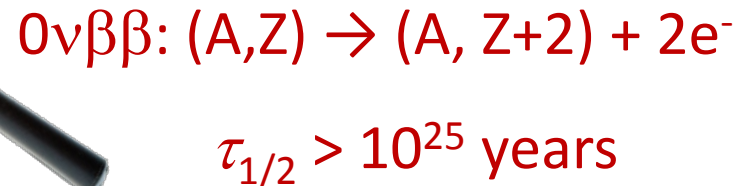
$$0\nu\beta\beta: (A,Z) \rightarrow (A, Z+2) + 2e^-$$
$$\tau_{1/2} > 10^{25} \text{ years}$$

Neutrinoless double beta decay

If conservation rules don't allow simple beta decay:



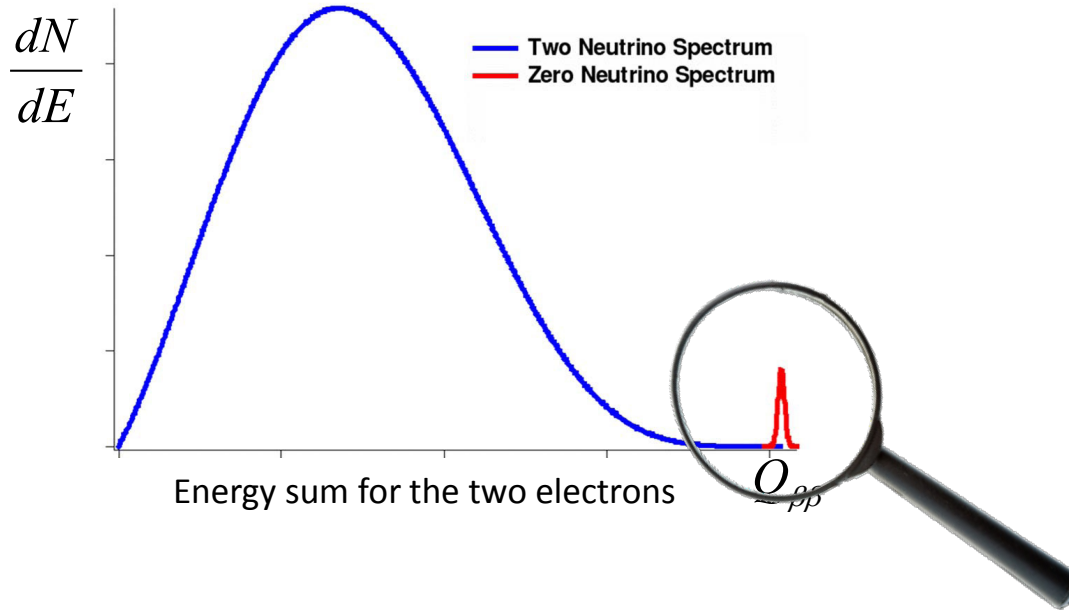
If $\nu = \bar{\nu}$:



Very rare events!

Neutrinoless double beta decay - ν mass

The halflife for $0\nu 2\beta$ decay depends on the neutrino mass



Nuclear matrix element

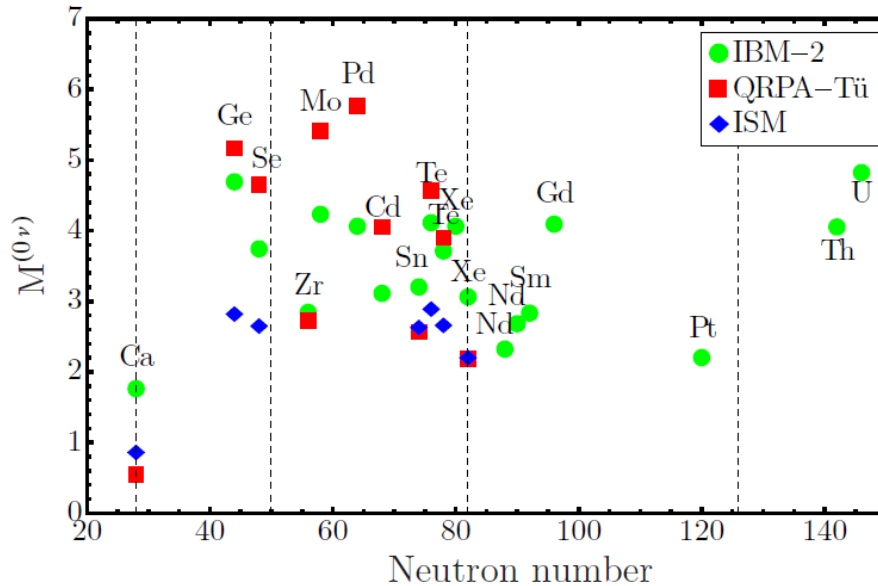
Phase space term

$$\left(\tau_{1/2}^{0\nu}\right)^{-1} = \left|\frac{m_{\beta\beta}}{m_e}\right|^2 \left|M_\nu^{0\nu}\right|^2 G^{0\nu}$$

$$m_{\beta\beta}^2 = \left|\sum U_{ei}^2 m(\nu_i)\right|^2$$

Neutrinoless double beta decay - ν mass

The halflife for $0\nu 2\beta$ decay depends on the neutrino mass



Nuclear matrix element

Phase space term

$$\left(\tau_{1/2}^{0\nu}\right)^{-1} = \left|\frac{m_{\beta\beta}}{m_e}\right|^2 \left|M_{\nu}^{0\nu}\right|^2 G^{0\nu}$$

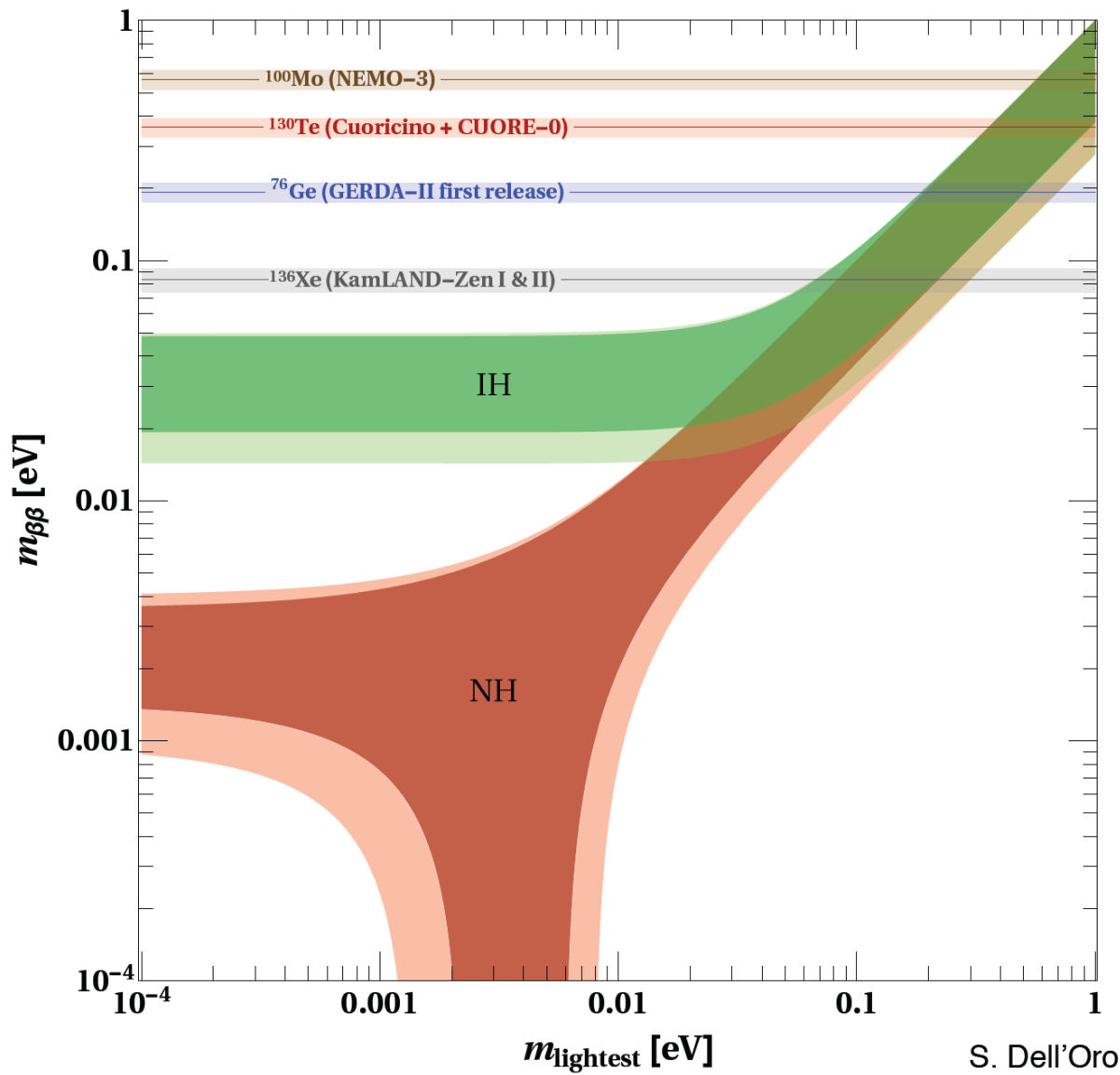
$$m_{\beta\beta}^2 = \left|\sum U_{ei}^2 m(\nu_i)\right|^2$$

Uncertainties to evaluate the effective Majorana mass due to:

➤ Nuclear matrix element

QRPA-Tu: F. Simkovic *et al.*, **Phys. Rev. C** **87**, 045501 (2013)
 pnQRPA: J. Hyvarinen *et al.*, **Phys. Rev. C** **91**, 024613 (2015)
 IBM-2: J. Barea *et al.*, **Phys. Rev. C** **91**, 034304 (2015)
 ISM: J. Menendez *et al.*, **Nucl. Phys. A** **818**, 139 (2009)

Present status



Neutrinoless double beta decay - sensitivity

Typically an excess of events is not found...

A limit on the halflife for $0\nu 2e$ decay can be defined as function of:

Mass of the isotope	M	[kg]	} Exposure $M \times T$	[kg \times year]
Measuring time	T	[year]		
Energy resolution	ΔE	[keV]		
Background index	b	[keV ⁻¹ ton ⁻¹ year ⁻¹]		

Two limits defined by the background index

> 1 background events in ROI

$$\left(\tau_{1/2}^{\text{exp}}\right)^{-1} = (\ln 2) N_a \frac{a}{A} \varepsilon \sqrt{\frac{MT}{b\Delta E}}$$

< 1 background events in ROI

$$\left(\tau_{1/2}^{\text{exp}}\right)^{-1} = (\ln 2) N_a \frac{a}{A} \varepsilon \frac{MT}{n_{CL}}$$

Neutrinoless double beta decay - nuclides

Is there a preferred nuclide?

$$\left(\tau_{1/2}^{\text{exp}}\right)^{-1} = (\ln 2) N_a \frac{a}{A} \varepsilon \sqrt{\frac{MT}{b\Delta E}}$$

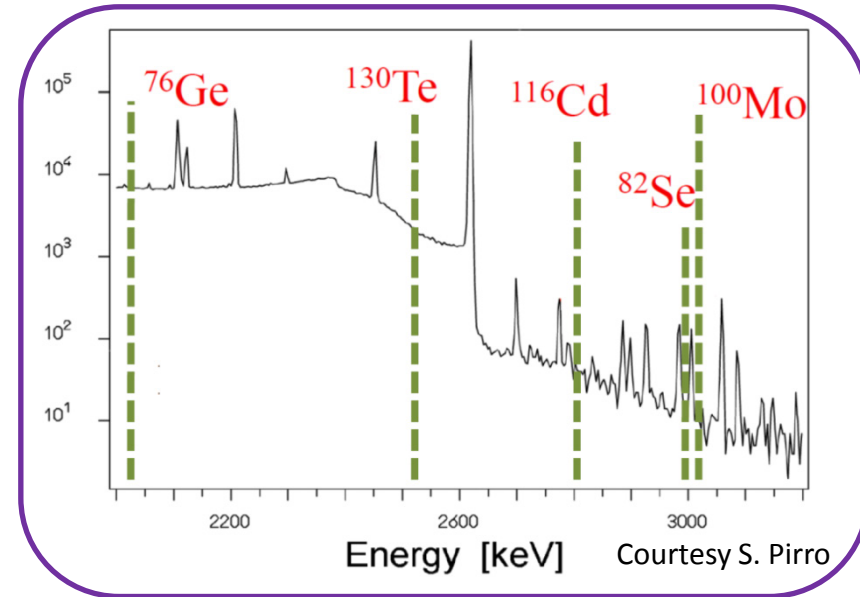
transition	$G^{01}(E_0, Z)$ $\times 10^{14} y$	$Q_{\beta\beta}$ [MeV]	Abund. (%)
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	26.9	3.667	6
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	8.04	4.271	0.2
$^{96}\text{Zr} \rightarrow ^{96}\text{Mo}$	7.37	3.350	3
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	6.24	2.802	7
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	5.92	2.479	9
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	5.74	3.034	10
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	5.55	2.533	34
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	3.53	2.995	9
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	0.79	2.040	8

Neutrinoless double beta decay - nuclides

Is there a preferred nuclide?

$$\left(\tau_{1/2}^{\text{exp}}\right)^{-1} = (\ln 2) N_a \frac{a}{A} \varepsilon \sqrt{\frac{MT}{b\Delta E}}$$

transition	$G^{01}(E_0, Z)$ $\times 10^{14} \text{y}$	$Q_{\beta\beta}$ [MeV]	Abund. (%)
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	26.9	3.667	6
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	8.04	4.271	0.2
$^{96}\text{Zr} \rightarrow ^{96}\text{Mo}$	7.37	3.350	3
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	6.24	2.802	7
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	5.92	2.479	9
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	5.74	3.034	10
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	5.55	2.533	34
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	3.53	2.995	9
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	0.79	2.040	8

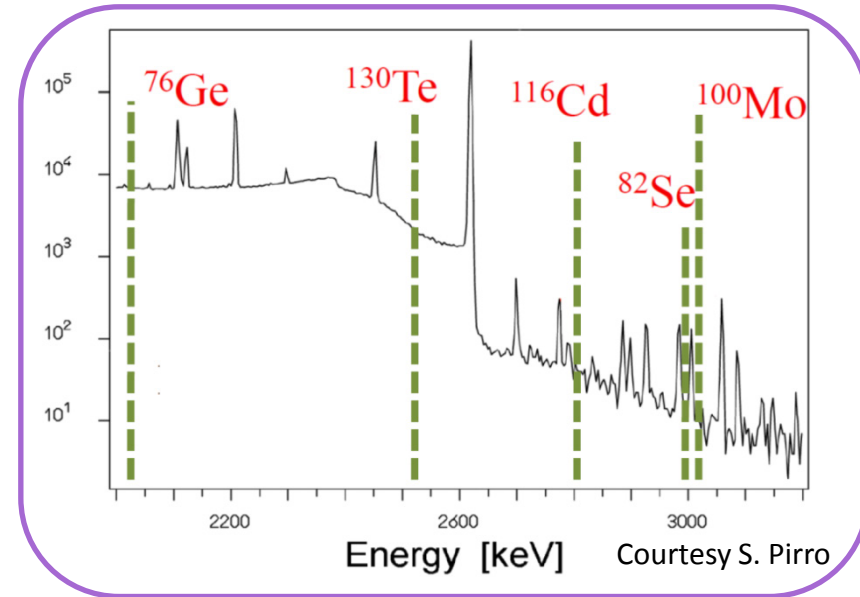


Neutrinoless double beta decay - nuclides

Is there a preferred nuclide?

$$\left(\tau_{1/2}^{\text{exp}}\right)^{-1} = (\ln 2) N_a \frac{a}{A} \frac{\sqrt{MT}}{\varepsilon \gamma b \Delta E}$$

transition	$G^{01}(E_0, Z)$ $\times 10^{14} \text{y}$	$Q_{\beta\beta}$ [MeV]	Abund. (%)
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	26.9	3.667	6
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	8.04	4.271	0.2
$^{96}\text{Zr} \rightarrow ^{96}\text{Mo}$	7.37	3.350	3
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	6.24	2.802	7
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	5.92	2.479	9
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	5.74	3.034	10
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	5.55	2.533	34
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	3.53	2.995	9
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	0.79	2.040	8



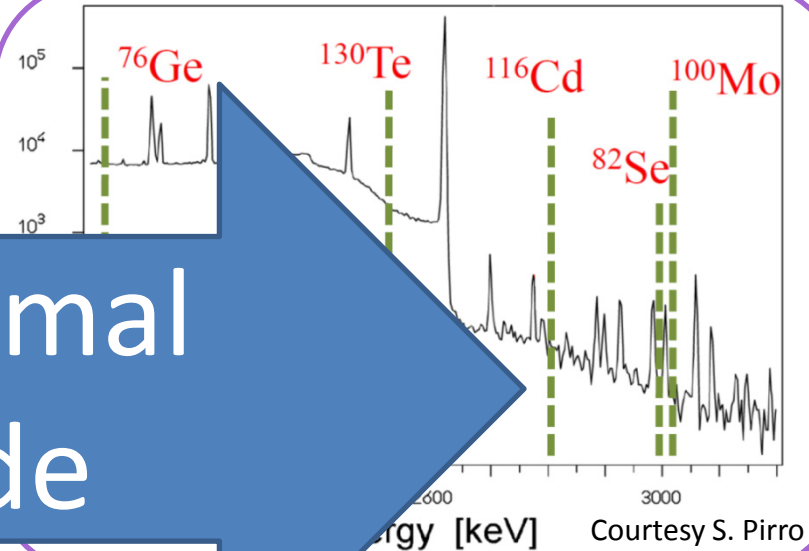
Detector Properties:

- Efficiency
- Energy resolution

Neutrinoless double beta decay - nuclides

Is there a preferred nuclide?

$$\left(\tau_{1/2}^{\text{exp}}\right)^{-1} = (\ln 2) N_a \frac{a}{4} \frac{\epsilon_{\gamma}}{b \Delta E} \sqrt{\frac{MT}{b \Delta E}}$$



No optimal nuclide

transition

$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	7.1	3.350	3
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	7.1	3.350	3
$^{96}\text{Zr} \rightarrow ^{96}\text{Mo}$	7.1	3.350	3
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	6.24	2.802	7
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	5.92	2.479	9
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	5.74	3.034	10
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	5.55	2.533	34
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	3.53	2.995	9
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	0.79	2.040	8

Detector Properties:

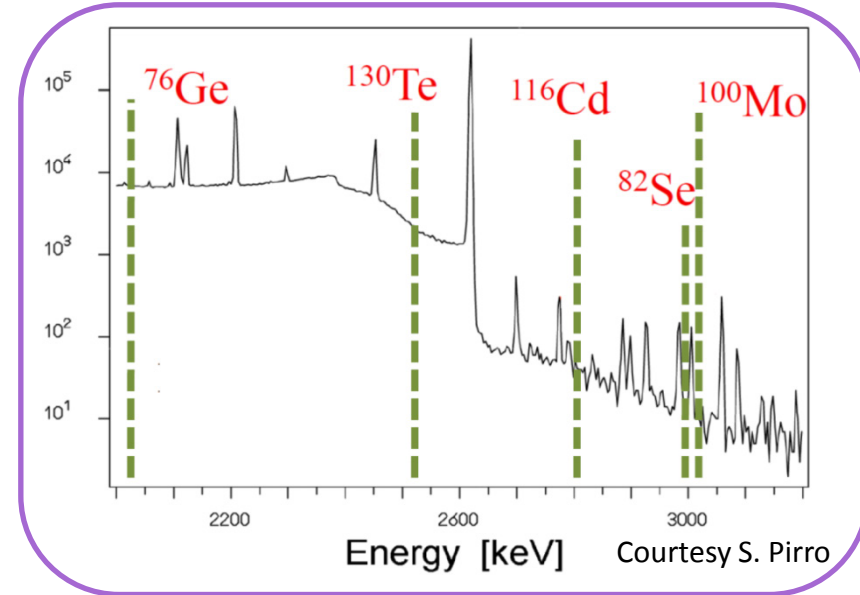
- Efficiency
- Energy resolution

Neutrinoless double beta decay - nuclides

Is there a preferred nuclide?

$$\left(\tau_{1/2}^{\text{exp}}\right)^{-1} = (\ln 2) N_a \frac{a}{A} \frac{\epsilon \sqrt{MT}}{b \Delta E}$$

transition	$G^{01}(E_0, Z)$ $\times 10^{14} y$	$Q_{\beta\beta}$ [MeV]	Abund. (%)
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	26.9	3.667	6
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	8.04	4.271	0.2
$^{96}\text{Zr} \rightarrow ^{96}\text{Mo}$	7.37	3.350	3
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	6.24	2.802	7
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	5.92	2.479	9
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	5.74	3.034	10
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	5.55	2.533	34
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	3.53	2.995	9
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	0.79	2.040	8



Detector Properties:

- Efficiency
- Energy resolution

Background

Many different experiments

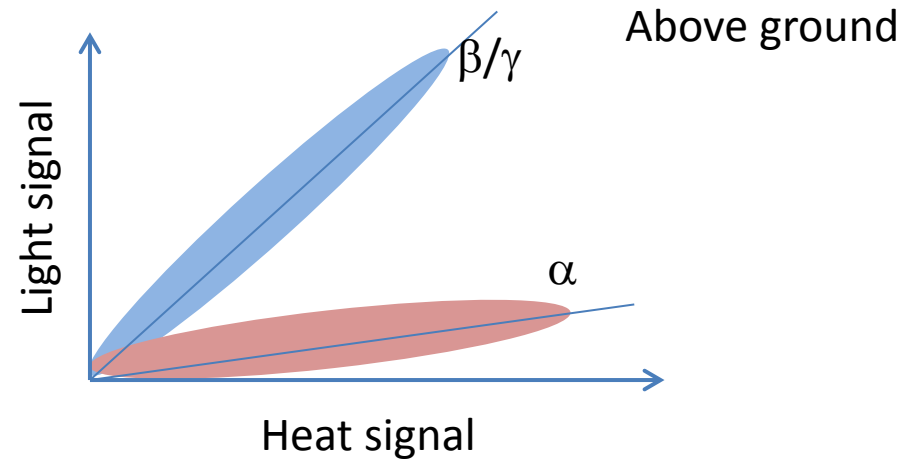
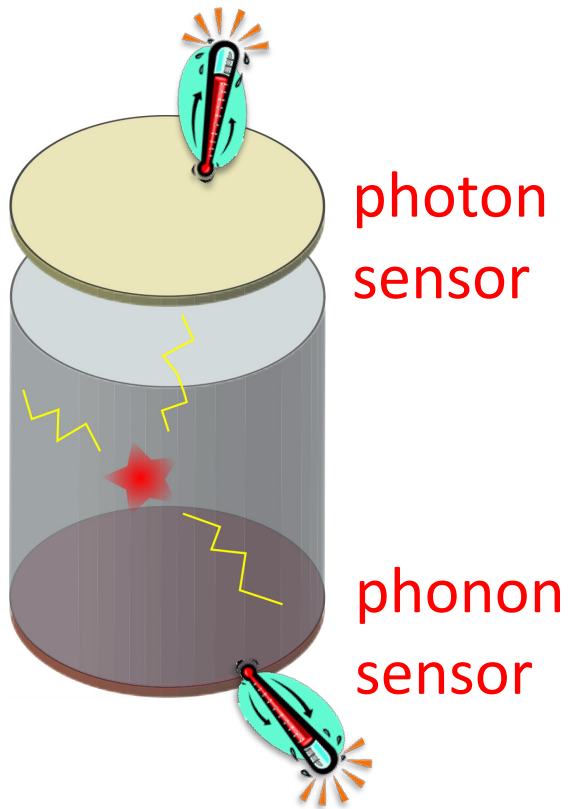
Experiment	Isotope	Technique	Mass $\beta\beta(0\nu)$ isotope
CUORICINO	^{130}Te	TeO ₂ Bolometer	10 kg
NEMO3	$^{100}\text{Mo}/^{82}\text{Se}$	Foils with tracking	6.9/0.9 kg
GERDA I	^{76}Ge	Ge diodes in LAr	15 kg
EXO200	^{136}Xe	Xe liquid TPC	160 kg
KamLAND-ZEN	^{136}Xe	2.7% in liquid scint.	380 kg
CUORE-0	^{130}Te	TeO ₂ Bolometer	11 kg
GERDA II	^{76}Ge	Point contact Ge in LAr	30+35 kg
Majorana D	^{76}Ge	Point contact Ge	30 kg
CUORE	^{130}Te	TeO ₂ Bolometer	206 kg
SNO+	^{130}Te	0.3% natTe suspended in Scint	55 kg
NEXT-100	^{136}Xe	High pressure Xe TPC	80 kg
SuperNEMO D	^{82}Se	Foils with tracking	7 kg
CANDLES	^{48}Ca	305 kg of CaF ₂ crystals - liq. scint	0.3 kg
LUCIFER	^{82}Se	ZnSe scint. bolometer	18 kg
1TGe (GERDA+MJ)	^{76}Ge	Best technology from GERDA and MAJORANA	~ tonne
CUPID	-	Hybrid Bolometers	~ tonne
nEXO	^{136}Xe	Xe liquid TPC	~ tonne
SuperNEMO	^{82}Se	Foils with tracking	100 kg
AMoRE	^{100}Mo	CaMoO ₄ scint. bolometer	50 kg
MOON	^{100}Mo	Mo sheets	200 kg
COBRA	^{116}Cd	CdZnTe detectors	10 kg/183 kg
CARVEL	^{48}Ca	$^{48}\text{CaWO}_4$ crystal scint.	~ tonne
DCBA	^{150}Nd	Nd foils & tracking chambers	20 kg

← Best results

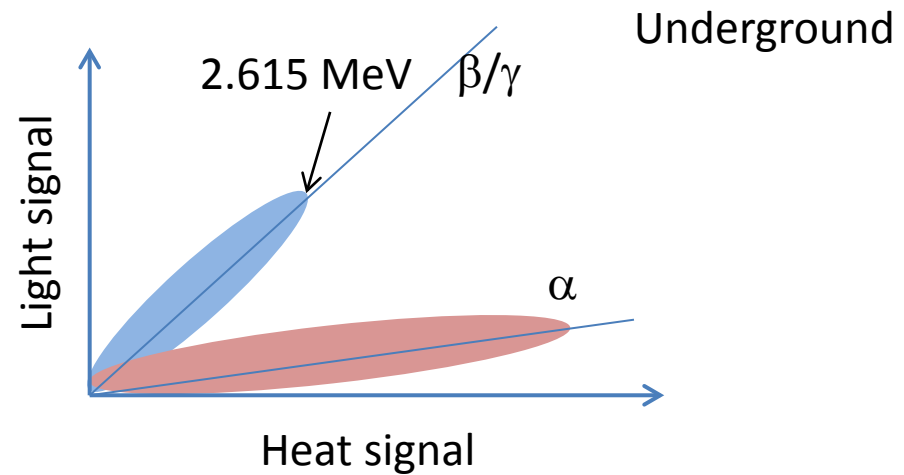
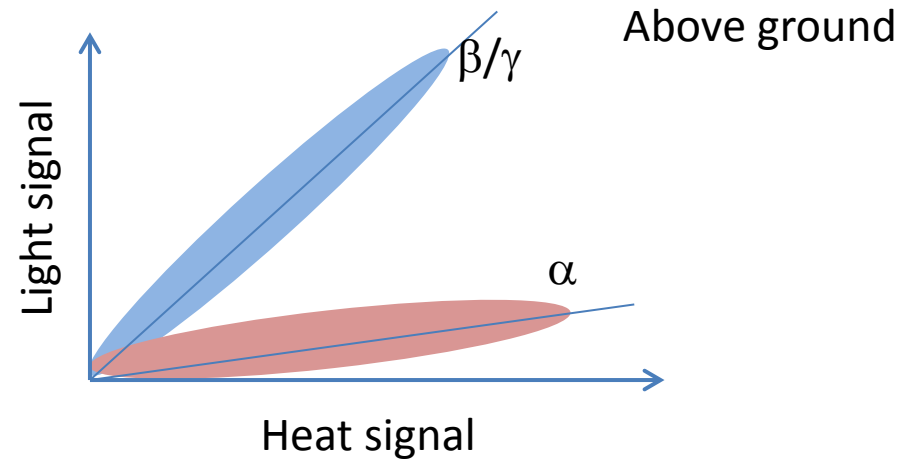
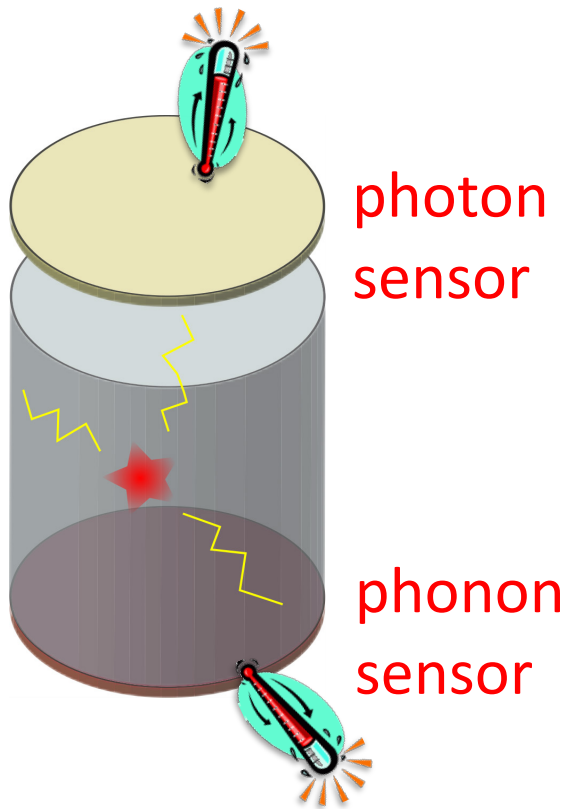


Promising technique

Promising technologies - Scintillating crystals

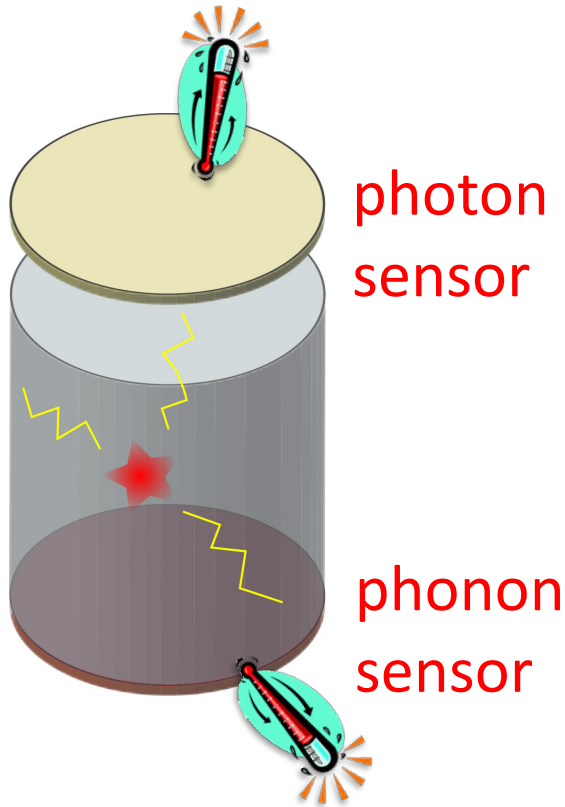


Promising technologies - Scintillating crystals



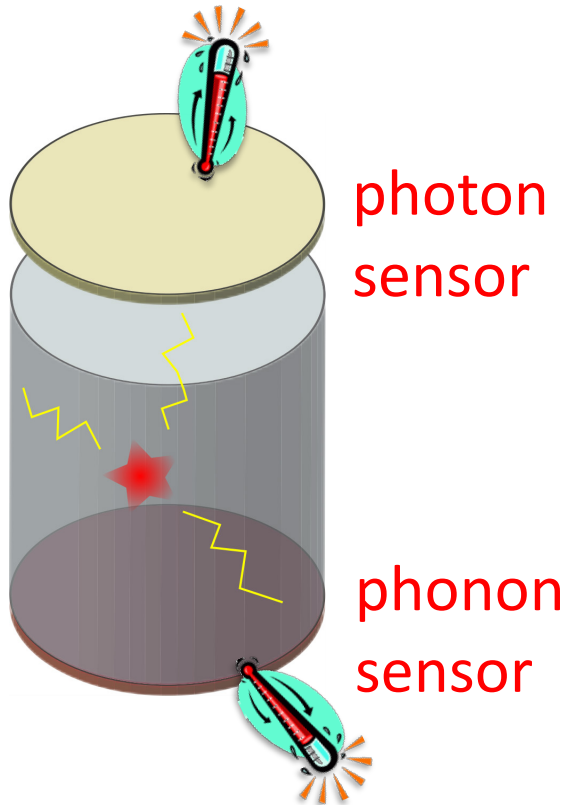
Background due to α particle can be removed

Promising technologies - Scintillating crystals



transition	$G^{01}(E_0, Z)$ $\times 10^{14}y$	$Q_{\beta\beta}$ [MeV]	Abund. (%)
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	26.9	3.667	← 6
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	8.04	4.271	← 0.2
$^{96}\text{Zr} \rightarrow ^{96}\text{Mo}$	7.37	3.350	← 3
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	6.24	2.802	← 7
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	5.92	2.479	9
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	5.74	3.034	← 10
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	5.55	2.533	34
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	3.53	2.995	← 9
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	0.79	2.040	8

Promising technologies - Scintillating crystals

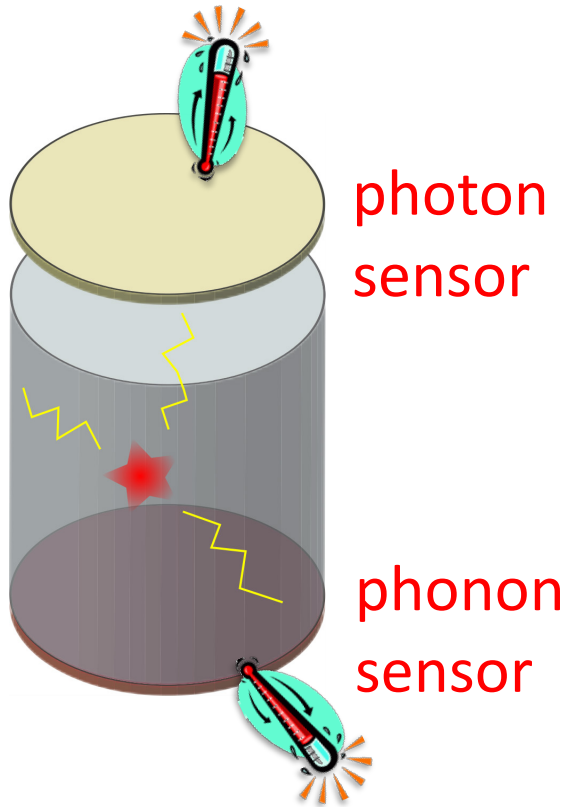


transition	$G^{01}(E_0, Z)$ $\times 10^{14}y$	$Q_{\beta\beta}$ [MeV]	Abund. (%)
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	26.9	3.667	← 6
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	8.04	4.271	← 0.2
$^{96}\text{Zr} \rightarrow ^{96}\text{Mo}$	7.37	3.350	← 3
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	6.24	2.802	← 7
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	5.92	2.479	9
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	5.74	3.034	← 10
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	5.55	2.533	34
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	3.53	2.995	← 9
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	0.79	2.040	8

Temperature signal: $\Delta T \cong \frac{\Delta E_{\text{phonon}}}{C}$

Light signal is also detected as $\Delta T \cong \frac{\Delta E_{\text{photon}}}{C}$ of a suitable photon detector

Promising technologies - Scintillating crystals



transition	$G^{01}(E_0, Z)$ $\times 10^{14}y$	$Q_{\beta\beta}$ [MeV]	Abund. (%)
$^{150}\text{Nd} \rightarrow ^{150}\text{Sm}$	26.9	3.667	← 6
$^{48}\text{Ca} \rightarrow ^{48}\text{Ti}$	8.04	4.271	← 0.2
$^{96}\text{Zr} \rightarrow ^{96}\text{Mo}$	7.37	3.350	← 3
$^{116}\text{Cd} \rightarrow ^{116}\text{Sn}$	6.24	2.802	← 7
$^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$	5.92	2.479	9
$^{100}\text{Mo} \rightarrow ^{100}\text{Ru}$	5.74	3.034	← 10
$^{130}\text{Te} \rightarrow ^{130}\text{Xe}$	5.55	2.533	34
$^{82}\text{Se} \rightarrow ^{82}\text{Kr}$	3.53	2.995	← 9
$^{76}\text{Ge} \rightarrow ^{76}\text{Se}$	0.79	2.040	8

Low T thermal detectors are the best candidate for these measurements

^{100}Mo -based experiments

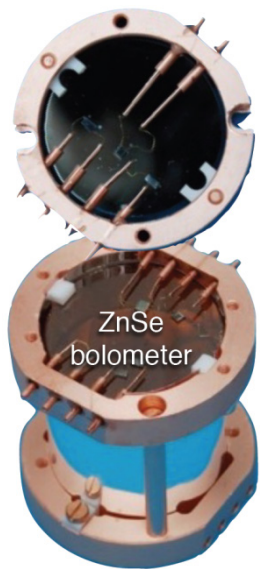
$$^{100}\text{Mo} \rightarrow ^{100}\text{Ru} + 2e^- + (2\nu_e) \quad T_{1/2} = [7.15 \pm 0.37 \text{ (stat)} \pm 0.66 \text{ (syst)}] \times 10^{18} \text{ y}$$

$$Q_{\beta\beta} = 3034 \text{ keV}$$

L. Cardani et al., J. Phys. G: Nucl. Part. Phys. 41 (2014) 075204

LUCIFER

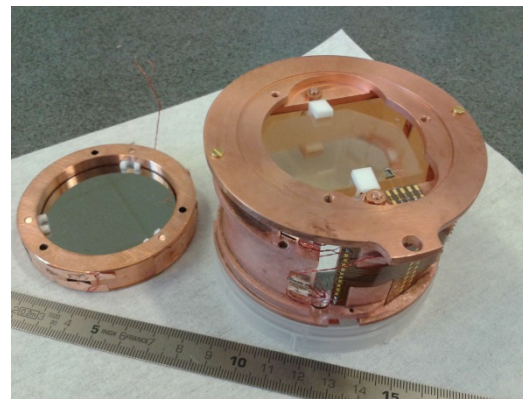
Different scintillating crystals coupled to NTD_Ge



LUCIFER, <http://arxiv.org/abs/1303.4080>
JINST 8 (2013) P05021

LUMINEU

ZnMoO_4 , LiMoO_4



NTD-Ge baseline for photon and phonon channel

MMC R&D for photon channel

LUMINEU arXiv:1704.01758
Submitted to EPJC

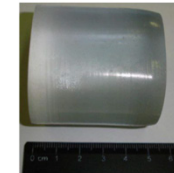
AMoRE

CaMoO_4

• SB28
weight 196 g



• SB29
weight 390 g



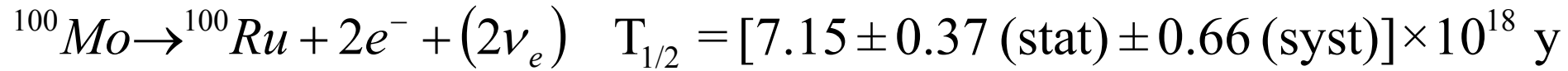
• S35
weight ~300 g



MMC for photon and phonon channel

Technical Design Report for the
AMoRE $0\nu\beta\beta$ Decay Search Experiment
[arXiv:1512.05957](http://arxiv.org/abs/1512.05957) [physics.ins-det]

^{100}Mo -based experiments



$$Q_{\beta\beta} = 3034 \text{ keV}$$

L. Cardani et al., J. Phys. G: Nucl. Part. Phys. 41 (2014) 075204

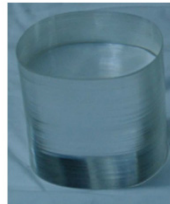
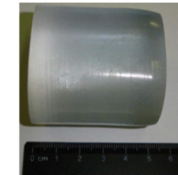
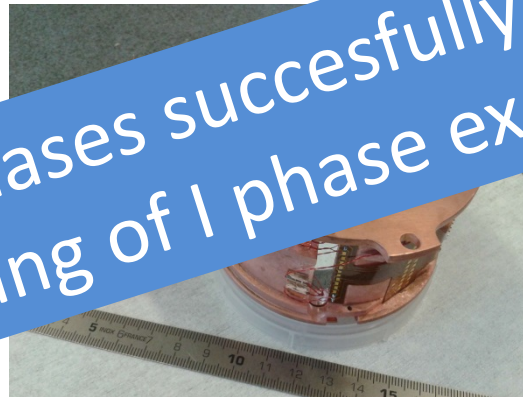
LUCIFER

Different scintillating crystals coupled to NTD_Ge



LUMINEU

ZnMoO_4 , LiMoO_4



• S35
weight ~300 g



R&D phases successfully concluded
Starting of I phase experiments

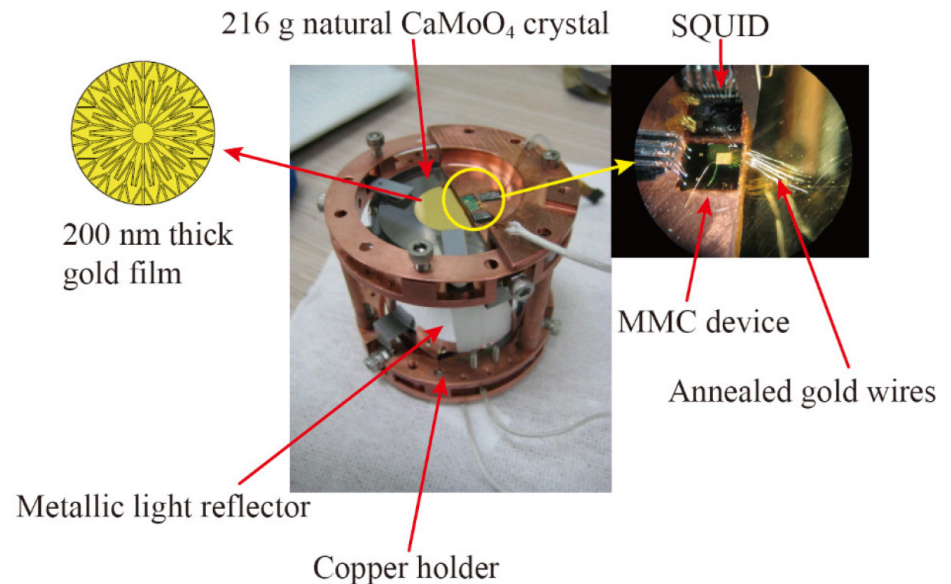
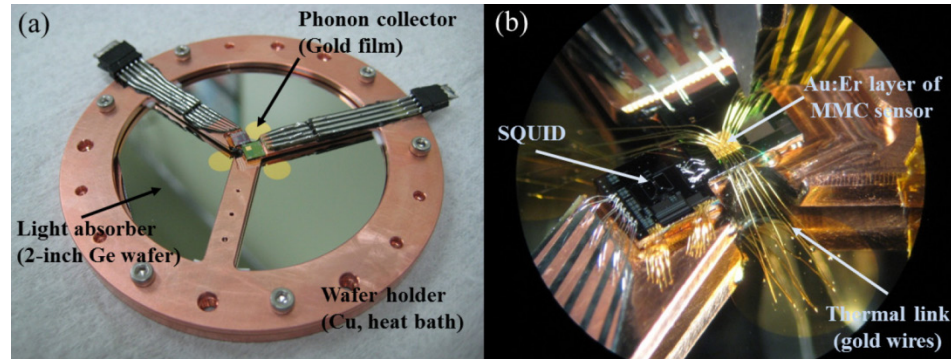
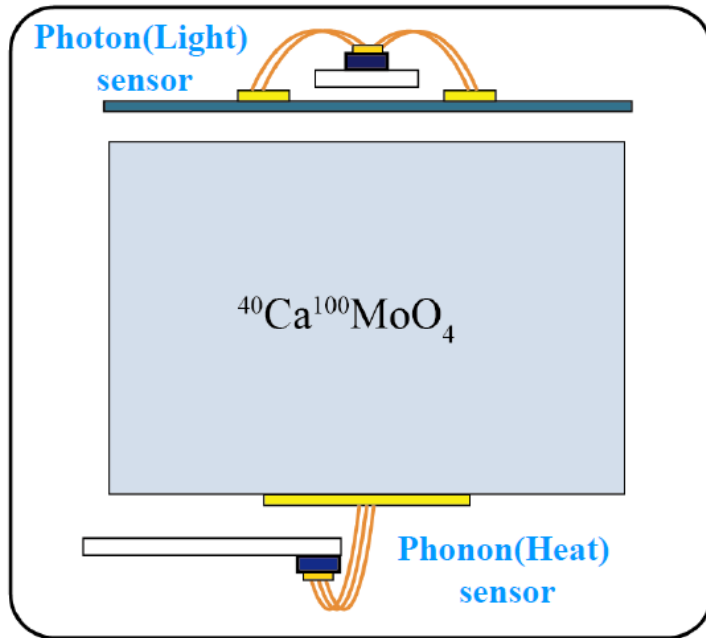
NTD-Ge baseline for photon and phonon channel

MMC R&D for photon channel

MMC for photon and phonon channel

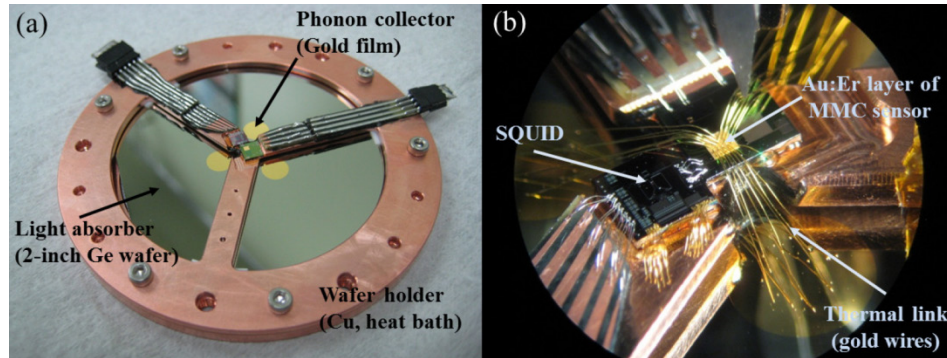
Approach used in AMoRe

Technical Design Report for the
AMoRE 0 $\nu\beta\beta$ Decay Search Experiment
[arXiv:1512.05957](https://arxiv.org/abs/1512.05957) [physics.ins-det]



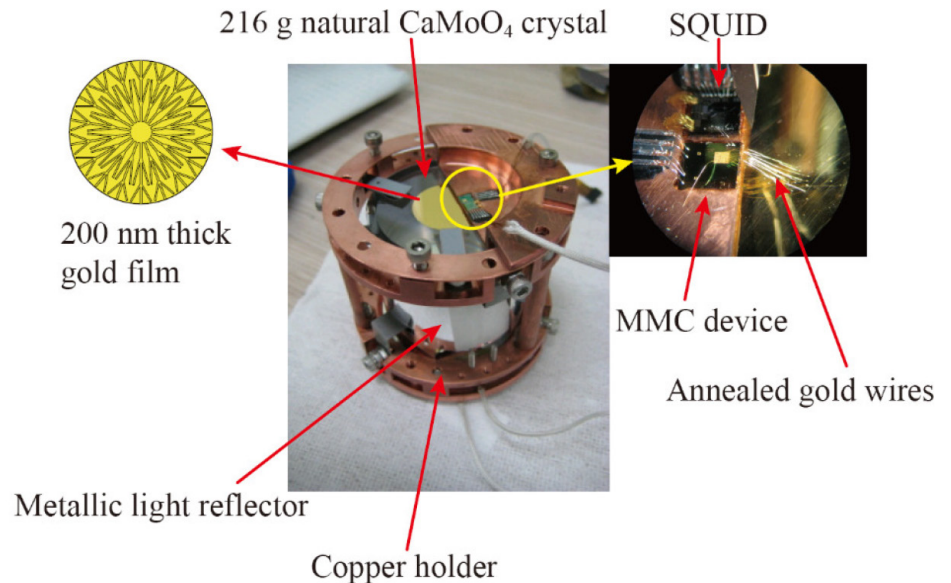
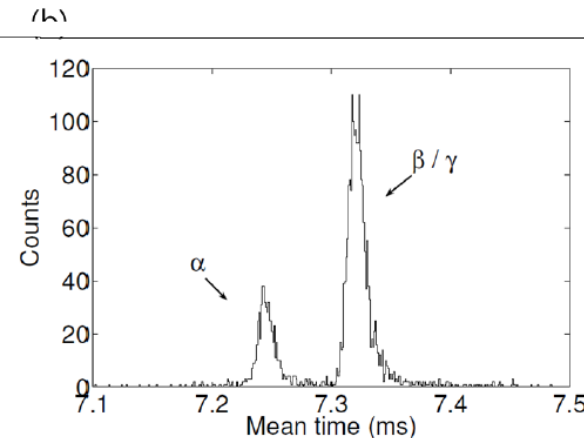
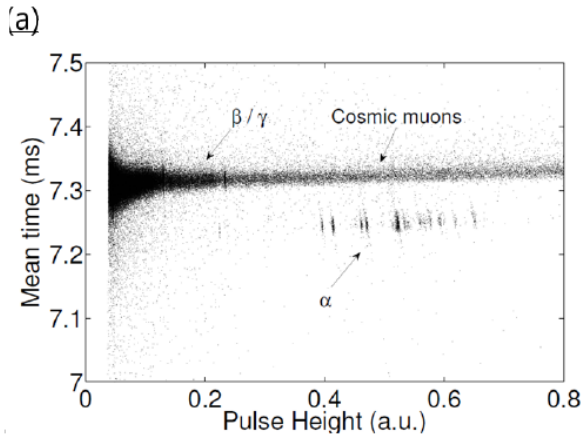
Approach used in AMoRe

Technical Design Report for the
 AMoRE $0\nu\beta\beta$ Decay Search Experiment
[arXiv:1512.05957](https://arxiv.org/abs/1512.05957) [physics.ins-det]



545 eV @ 6 keV

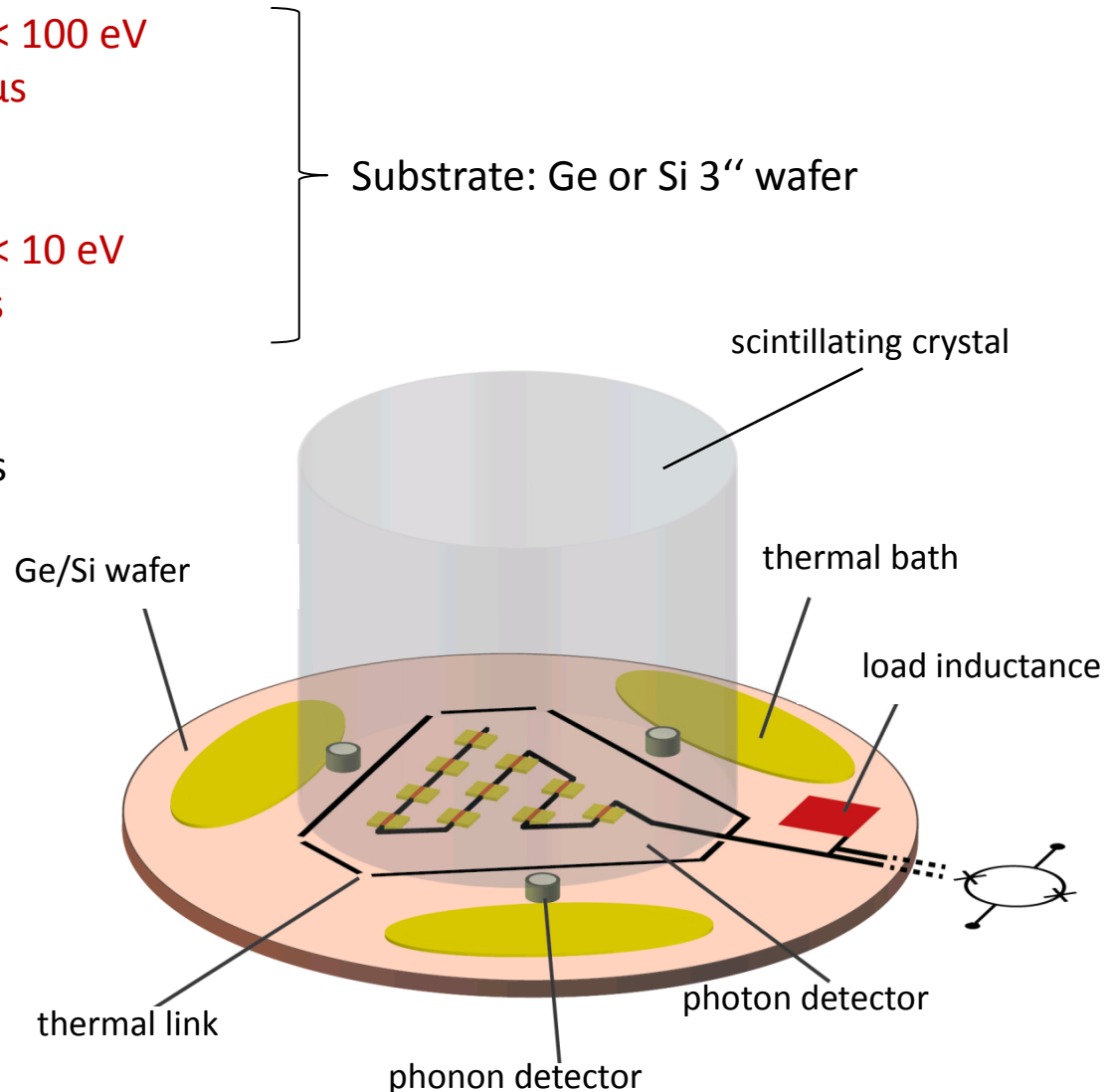
11 keV @ 2615 keV
 6.5 keV @ 583 keV



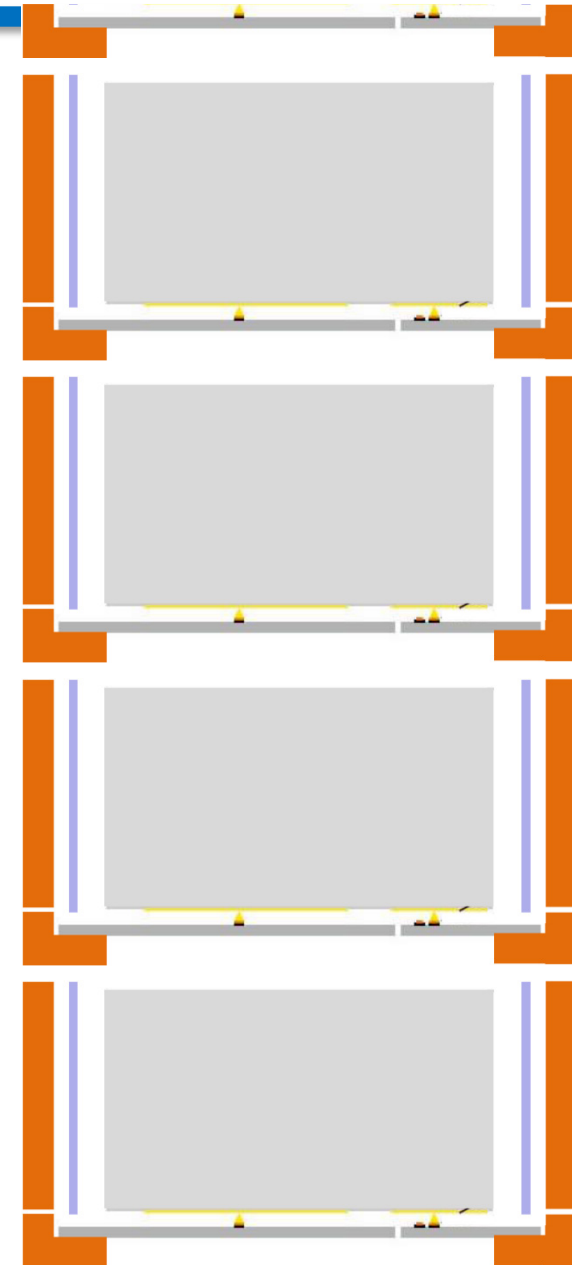
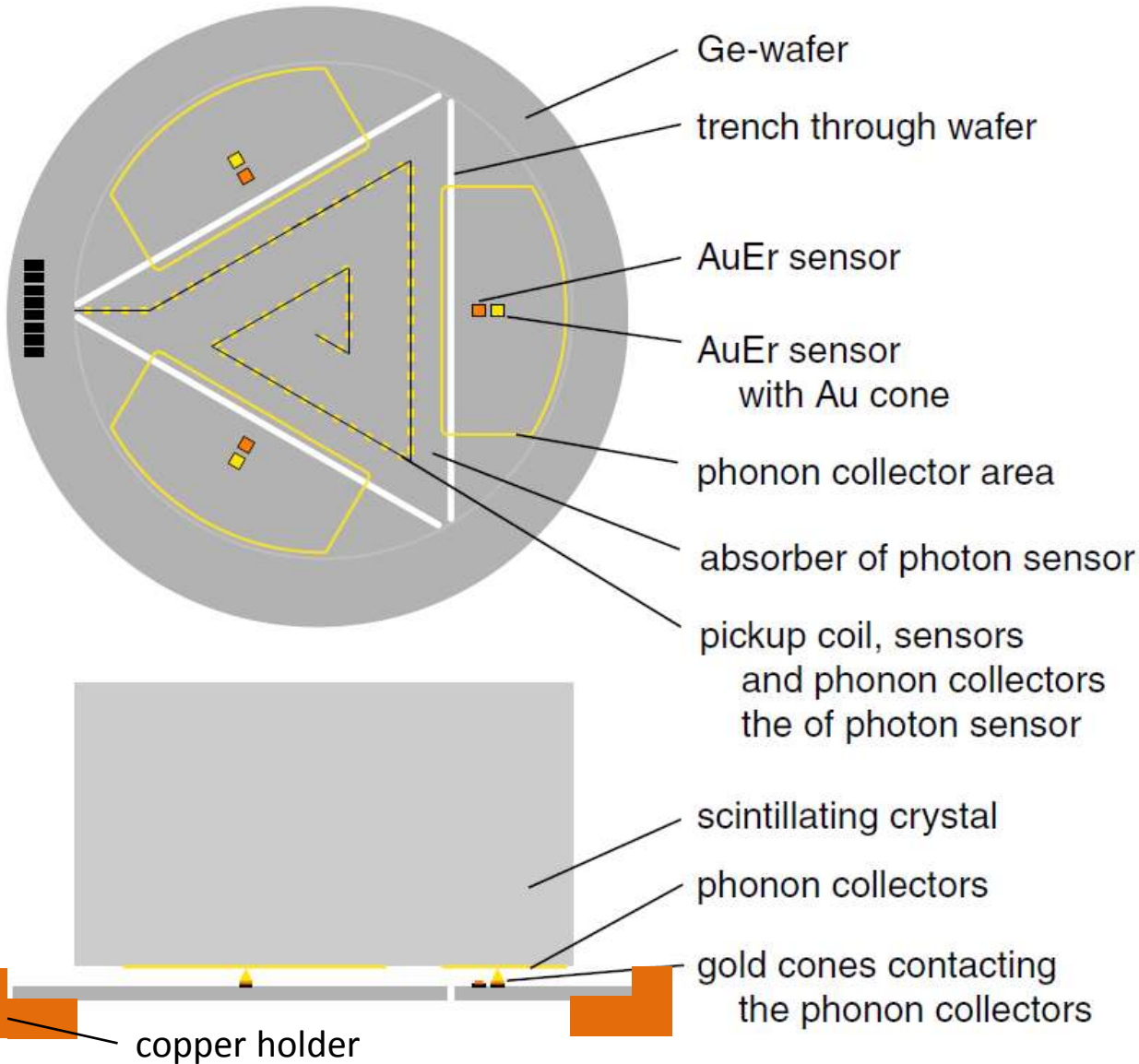
Combined Photon and Phonon Detector: P2

- Phonon detector:
 - energy resolution $\Delta E_{\text{FWHM}} < 100 \text{ eV}$
 - rise time $\tau < 200 \mu\text{s}$
- Photon detector:
 - energy resolution $\Delta E_{\text{FWHM}} < 10 \text{ eV}$
 - rise time $\tau < 50 \mu\text{s}$

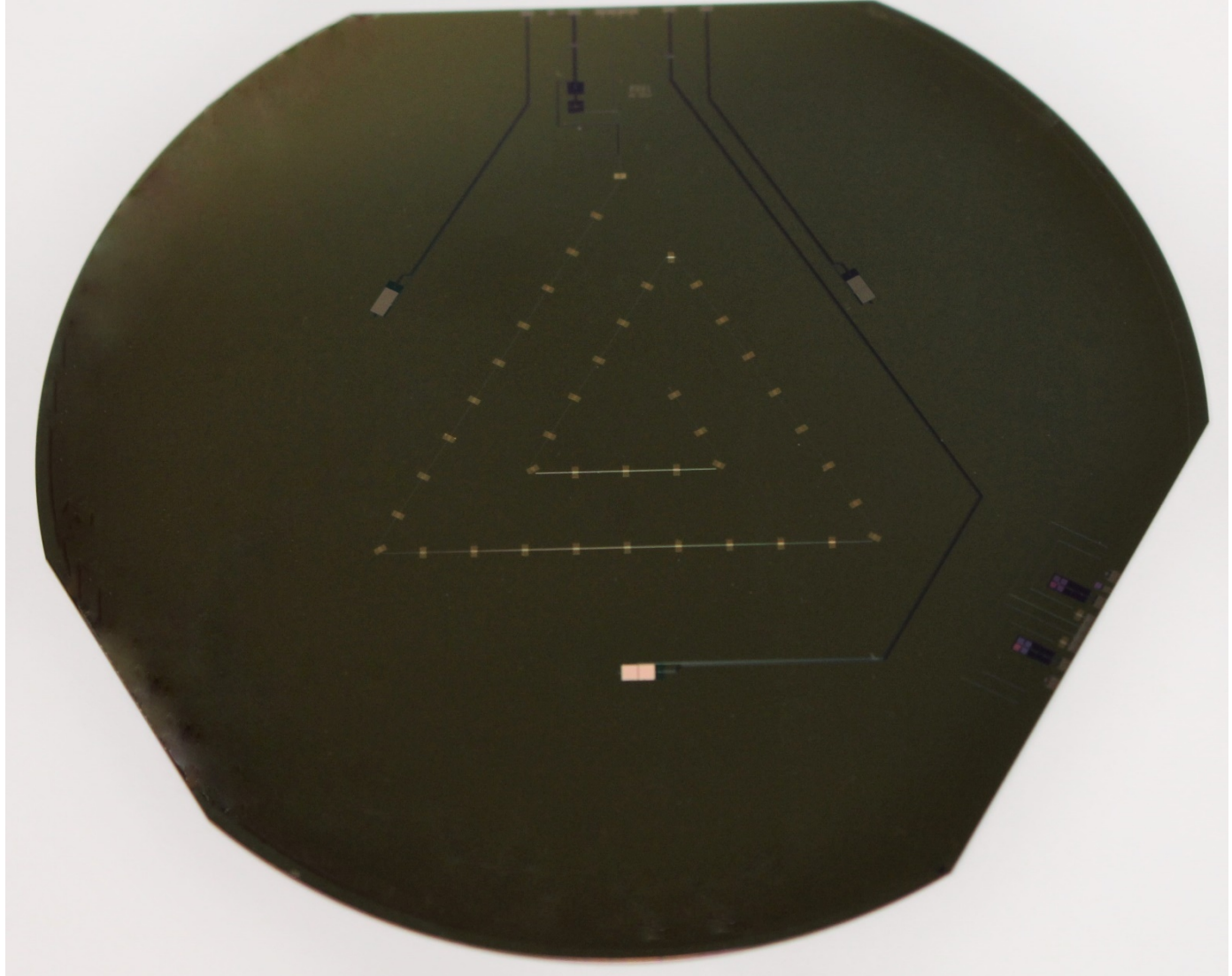
- A minimum of (contaminated?) parts
- Position sensitivity possible



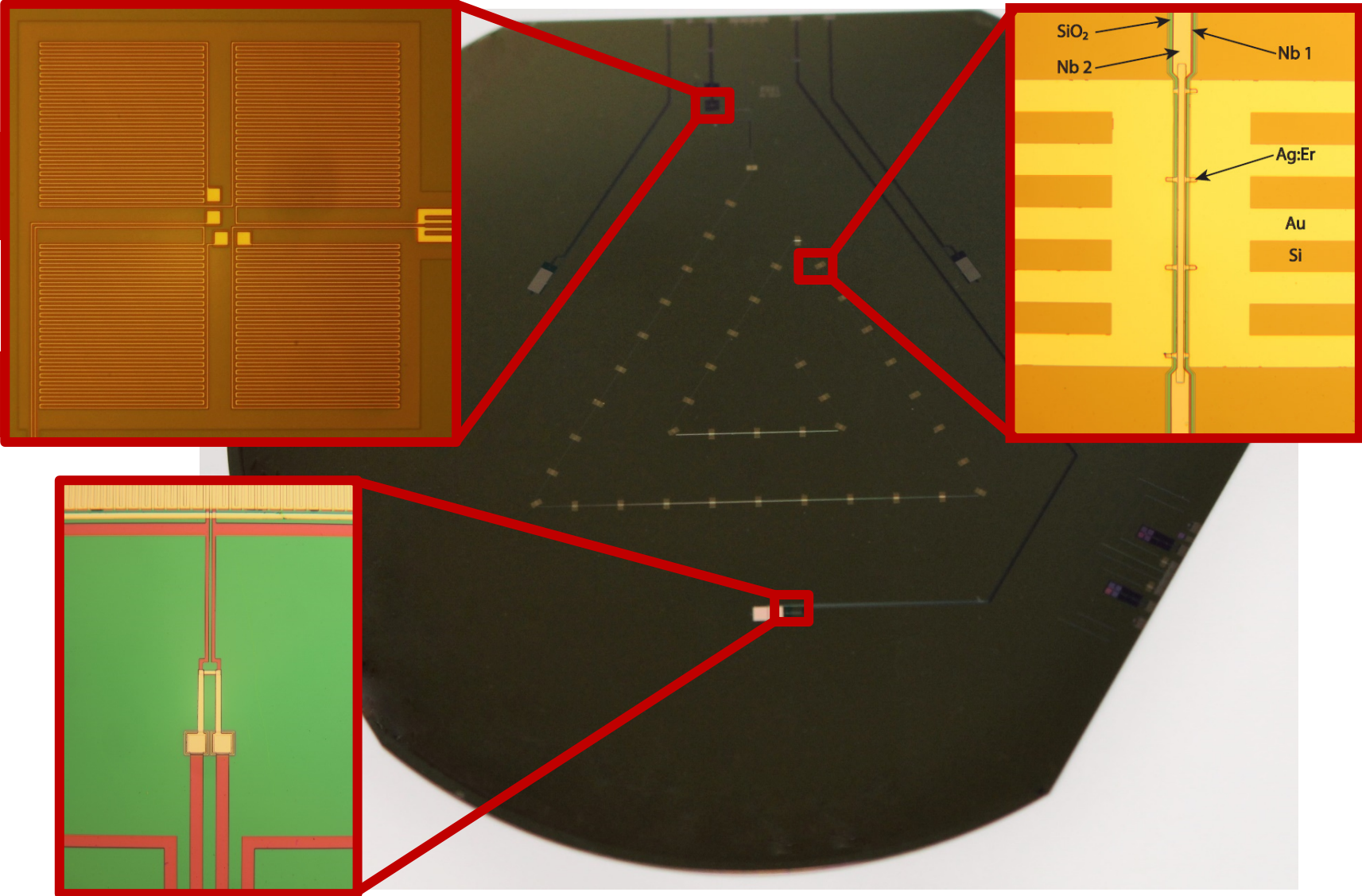
Combined photon and phonon detector: P2



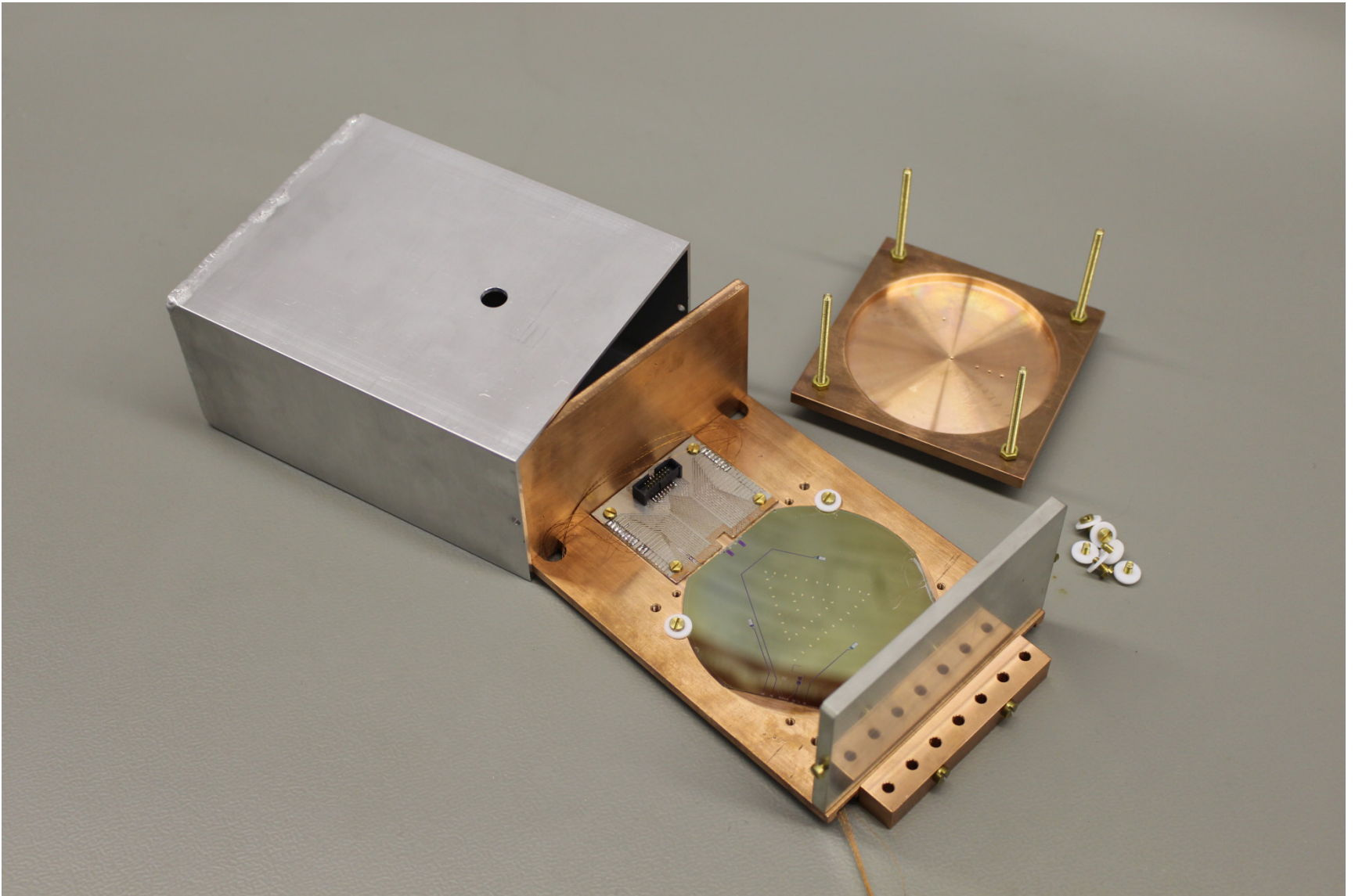
Integrated light and heat detectors P2



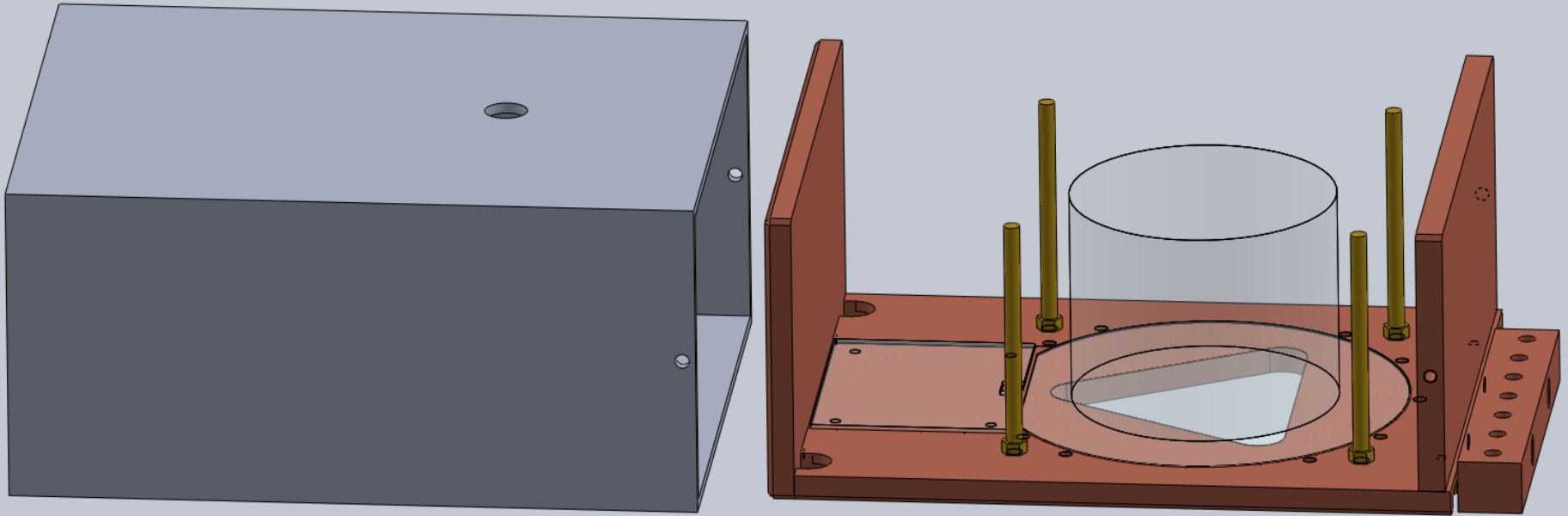
Integrated light and heat detectors P2



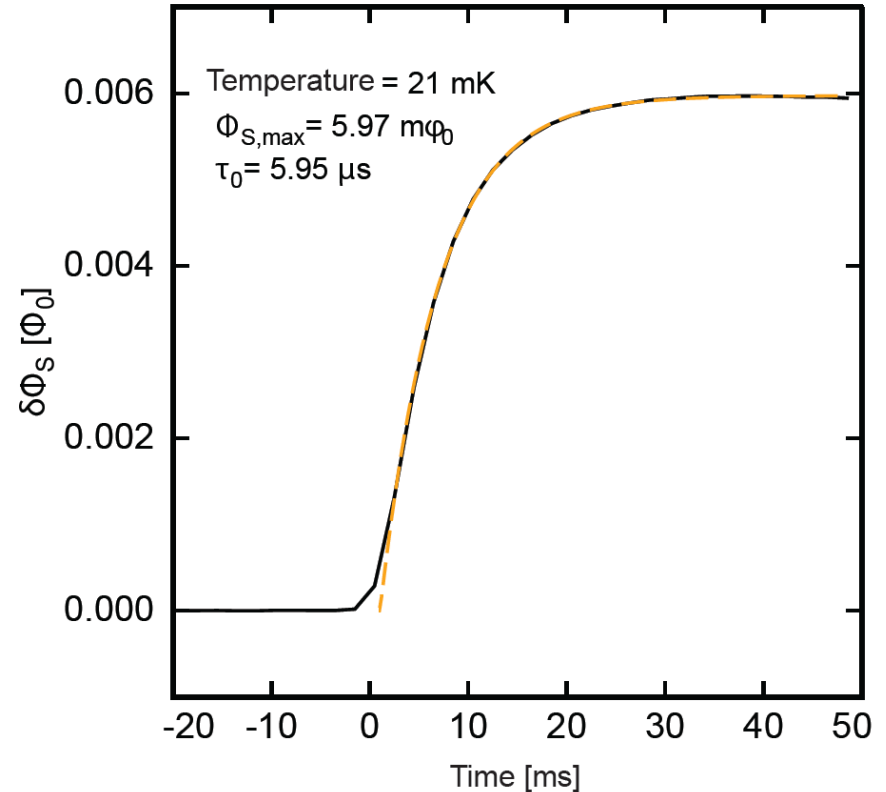
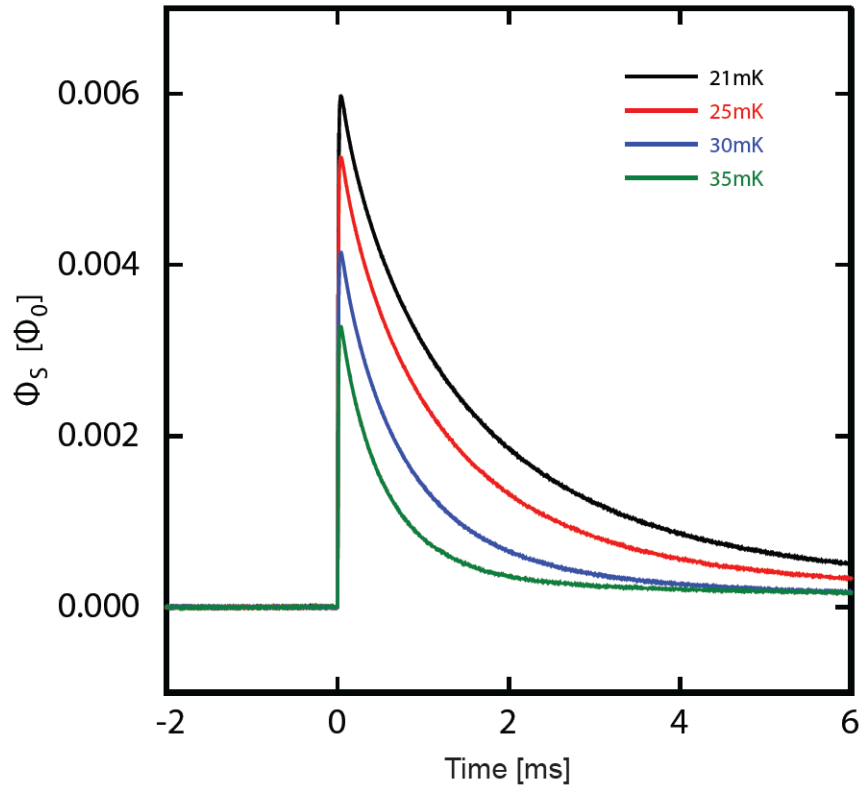
Experimental set-up for P2



Experimental set-up for P2



Photon detector: First tests with 6keV x-rays



Risetimes: **direct x-rays** $\sim 6 \text{ }\mu\text{s}$
for scintillation light $\sim 250\text{--}350 \text{ }\mu\text{s}$ (Saclay)

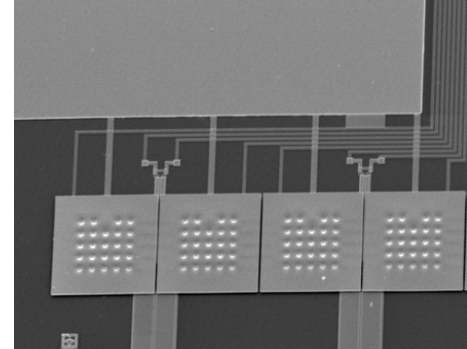
Conclusions and outlook

- Metallic magnetic calorimeters are **reliable** and **versatile** detectors
 - high resolution for all kinds of particles
 - wide range of energies
 - Fast signal rise time

- Direct determination of the **electron neutrino mass using ^{163}Ho** MMC have proved to **fulfil requirements**

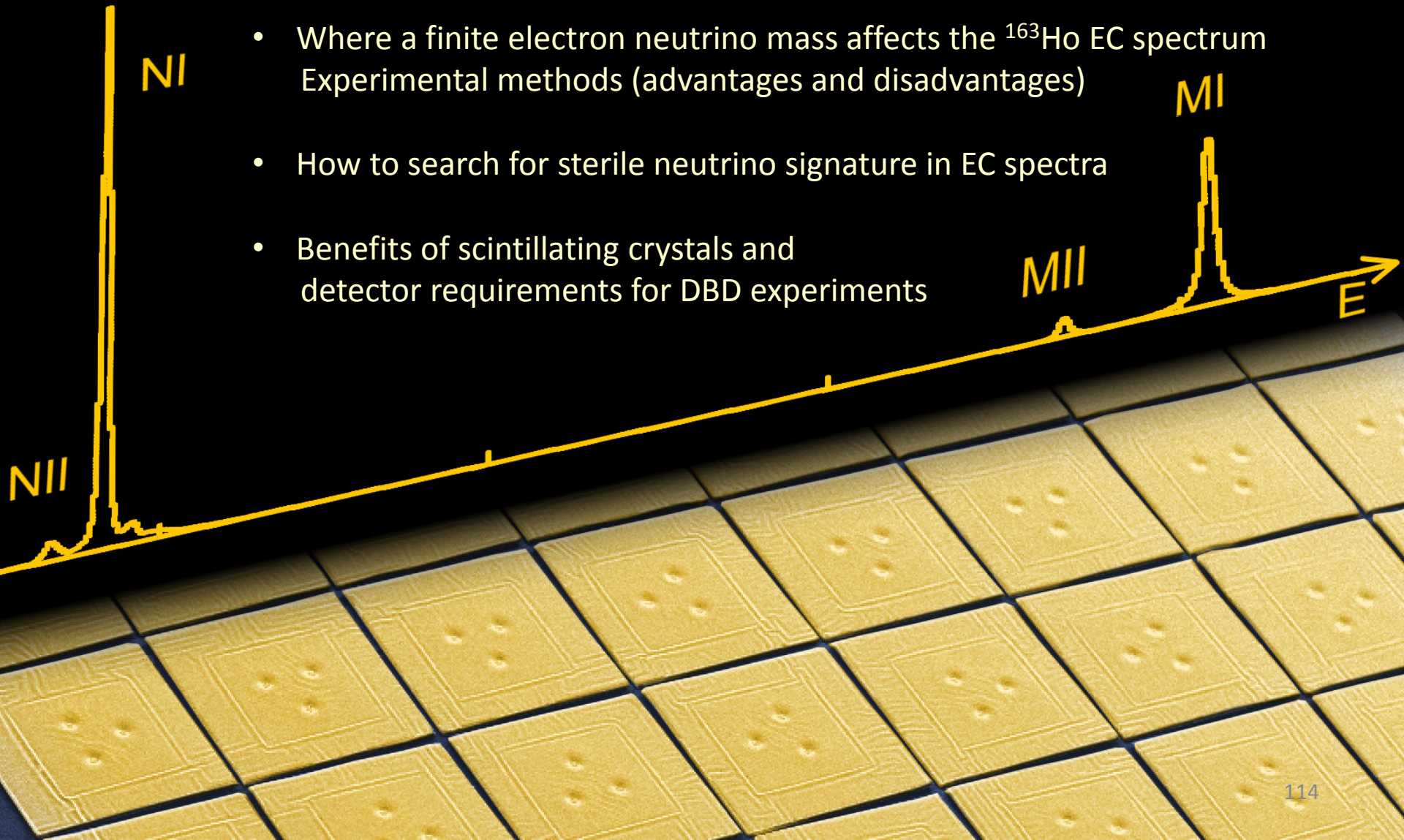
- eV-scale and keV-scale **sterile neutrinos** can be investigated through calorimetric measurement of **EC spectra**

- MMC-based **photon and phonon detectors** could bring significant benefit for large mass ^{100}Mo -based DBD experiments



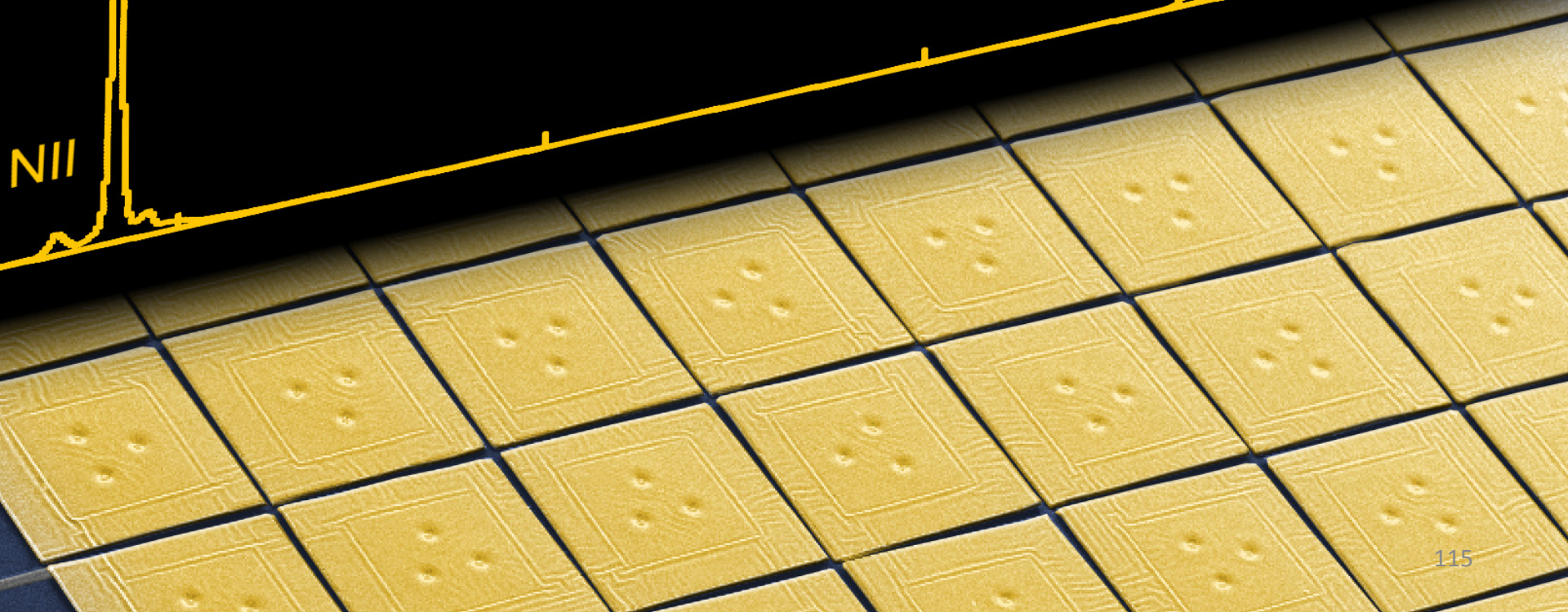
Take-home messages

- Working principle of MMCs and their performance
- Where a finite electron neutrino mass affects the ^{163}Ho EC spectrum
Experimental methods (advantages and disadvantages)
- How to search for sterile neutrino signature in EC spectra
- Benefits of scintillating crystals and detector requirements for DBD experiments



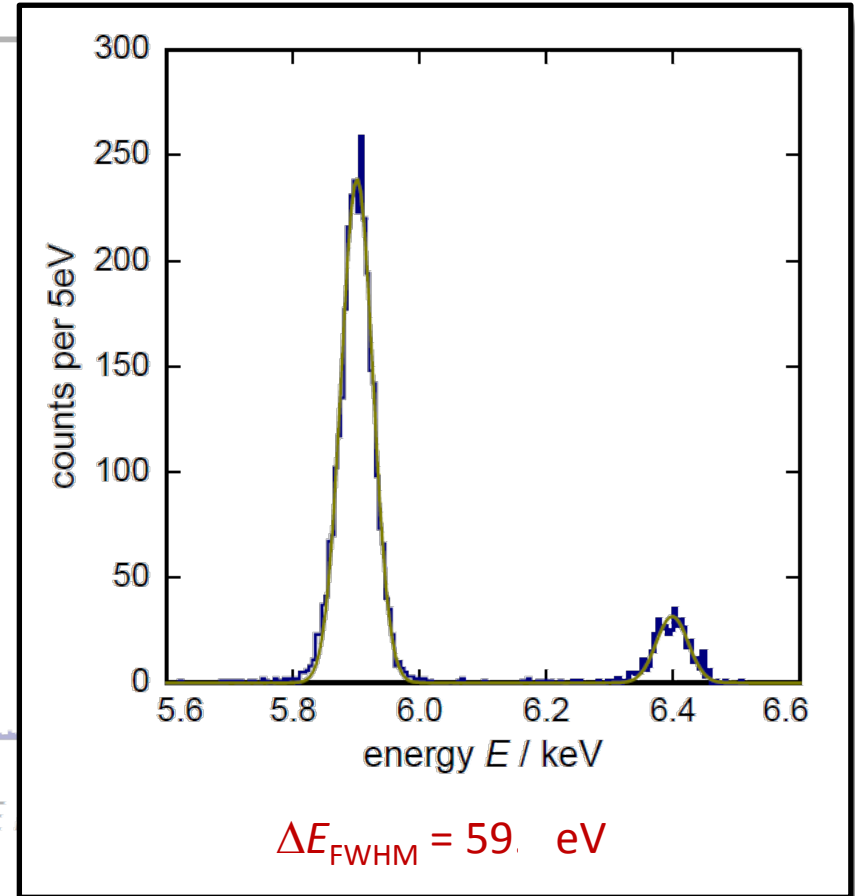
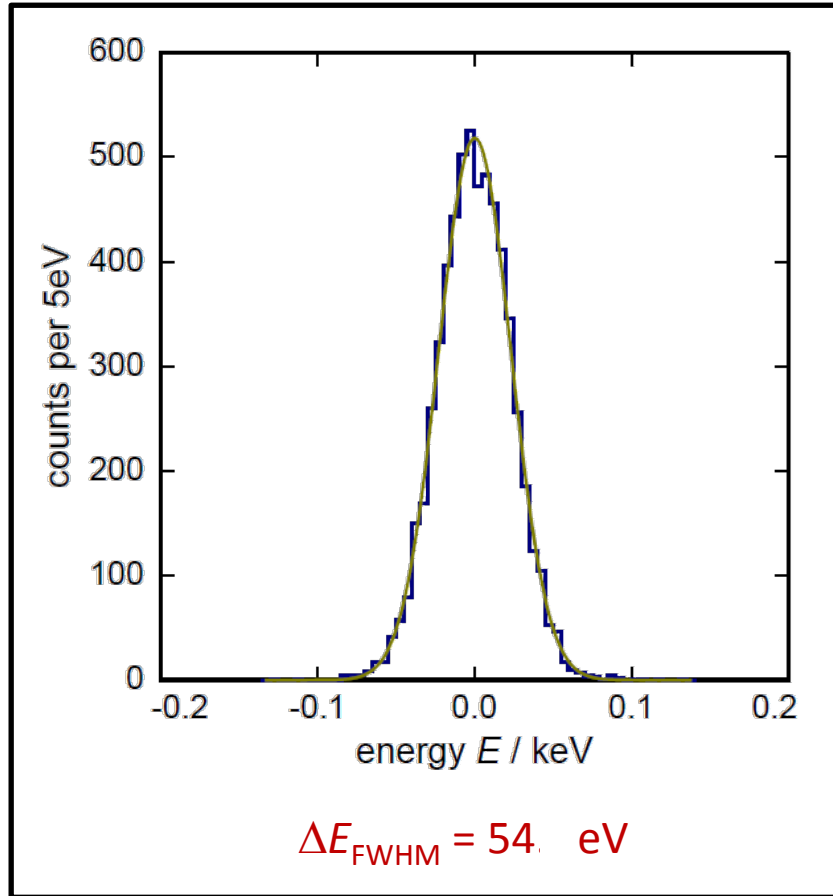


Thank you!



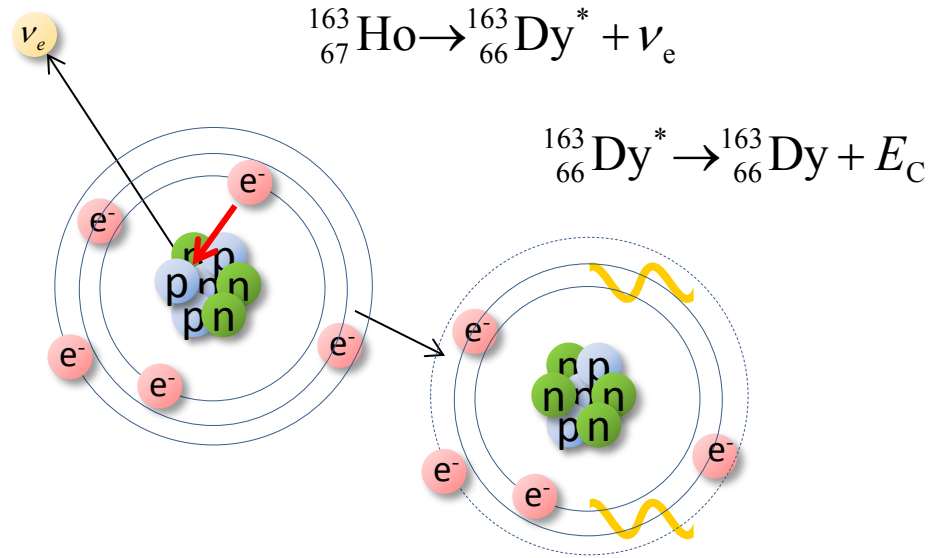
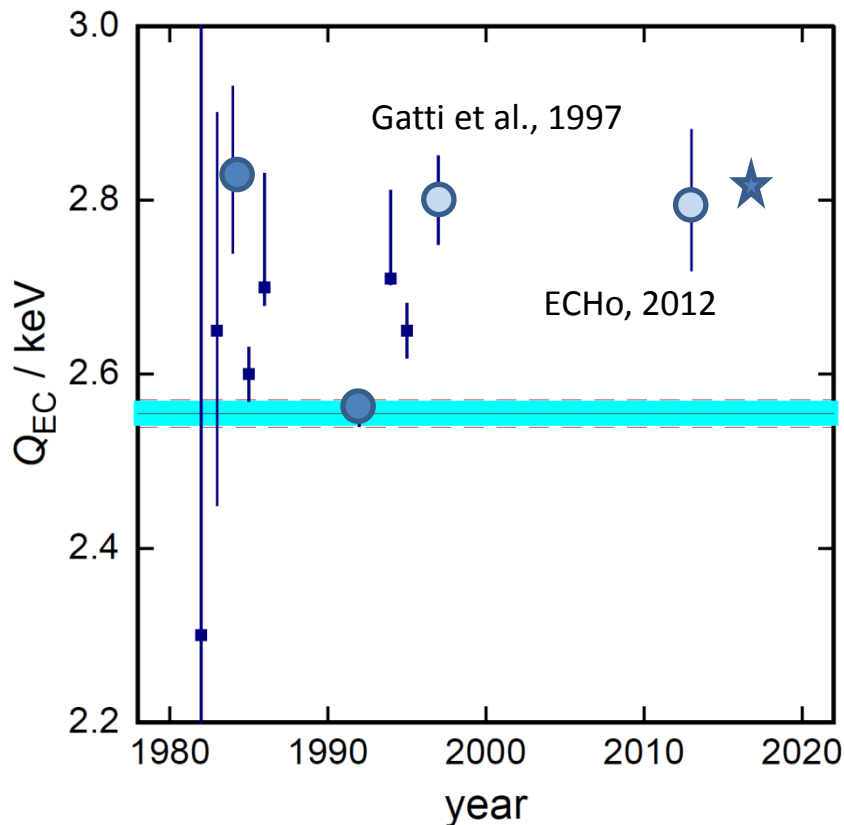
MMCs: Microwave SQUID multiplexing

measurement of the spectrum of ^{55}Fe to determine the energy resolution



Electron capture in ^{163}Ho : Q_{EC} determination

- Calorimetric measurements
- Measurements of x-rays
- ★ $Q_{\text{EC}} = m(^{163}\text{Ho}) - m(^{163}\text{Dy})$



• $\tau_{1/2} \cong 4570$ years ($2 \cdot 10^{11}$ atoms for 1 Bq)

• $Q_{\text{EC}} = (2.833 \pm 0.030^{\text{stat}} \pm 0.015^{\text{syst}})$ keV

S. Eliseev et al., *Phys. Rev. Lett.* **115** (2015) 062501

Electron capture in ^{163}Ho : Q_{EC} determination

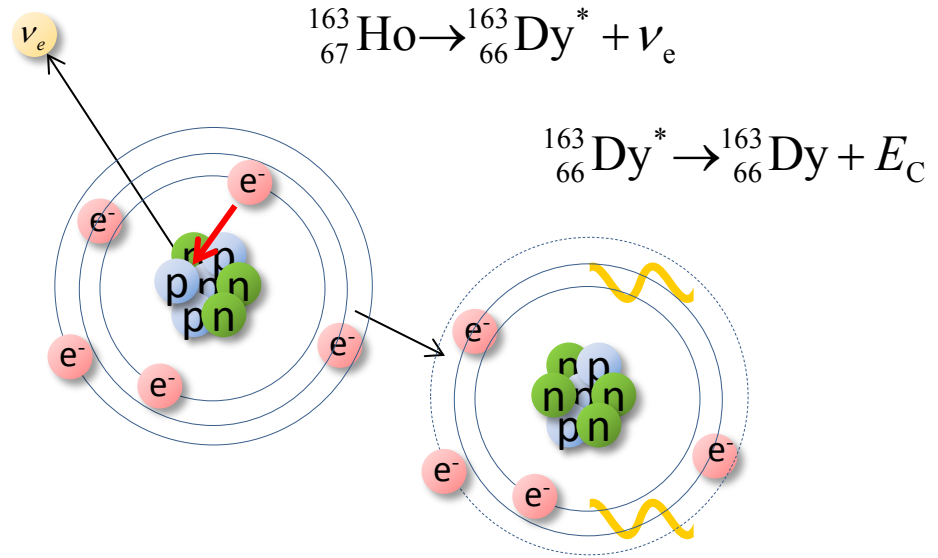
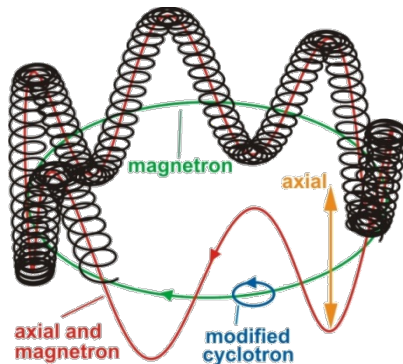
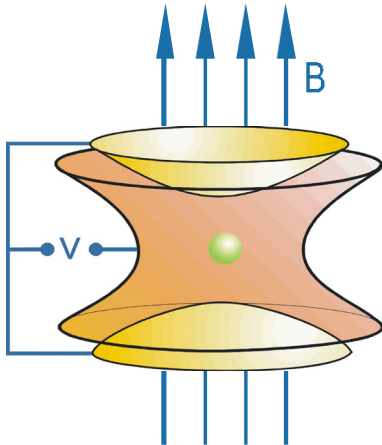
- Calorimetric measurements
- Measurements of x-rays
- ★ $Q_{\text{EC}} = m(^{163}\text{Ho}) - m(^{163}\text{Dy})$

Penning Trap Mass Spectroscopy

@TRIGA TRAP (Uni-Mainz) (*)

@SHIPTRAP (GSI – Darmstadt) (**)

$$v_c = \frac{qB}{m}$$



• $\tau_{1/2} \cong 4570$ years ($2 \cdot 10^{11}$ atoms for 1 Bq)

• $Q_{\text{EC}} = (2.833 \pm 0.030^{\text{stat}} \pm 0.015^{\text{syst}})$ keV

S. Eliseev et al., *Phys. Rev. Lett.* **115** (2015) 062501 (**)

F. Schneider et al., *Eur. Phys. J. A* **51** (2015) 89 (*)

High purity ^{163}Ho source in ECHO

Requirement : $>10^6 \text{ Bq} \rightarrow >10^{17}$ atoms

➤ (n, γ)-reaction on ^{162}Er

- High cross-section



- Radioactive contaminants



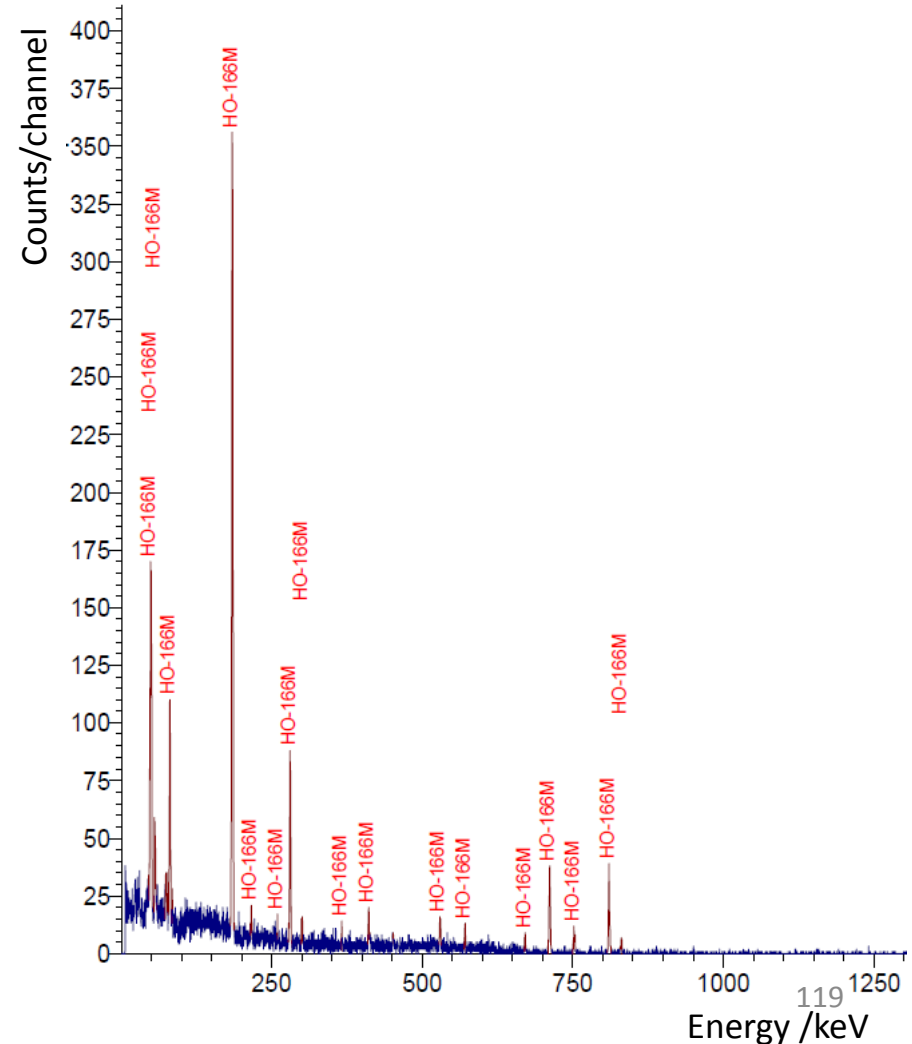
Er161 3.21 h 3/2-	Er162 0+	Er163 75.0 m 5/2-	Er164 0+	Er165 10.36 h 5/2-	Er166 0+
EC	0.14	EC	1.61	EC	33.6
Ho160 25.6 m 5+	Ho161 2.48 h 7/2-	Ho162 15.0 m 1+	Ho163 4570 y 7/2-	Ho164 29 m 1+	Ho165 7/2-
EC *	EC *	EC *	EC *	EC, β^- *	100

➤ Excellent chemical separation

Only $^{166\text{m}}\text{Ho}$

➤ Available ^{163}Ho source:

$\sim 10^{18}$ atoms



High purity ^{163}Ho source in ECHO

Requirement : $>10^6 \text{ Bq} \rightarrow >10^{17}$ atoms

➤ (n,γ)-reaction on ^{162}Er

- High cross-section



- Radioactive contaminants



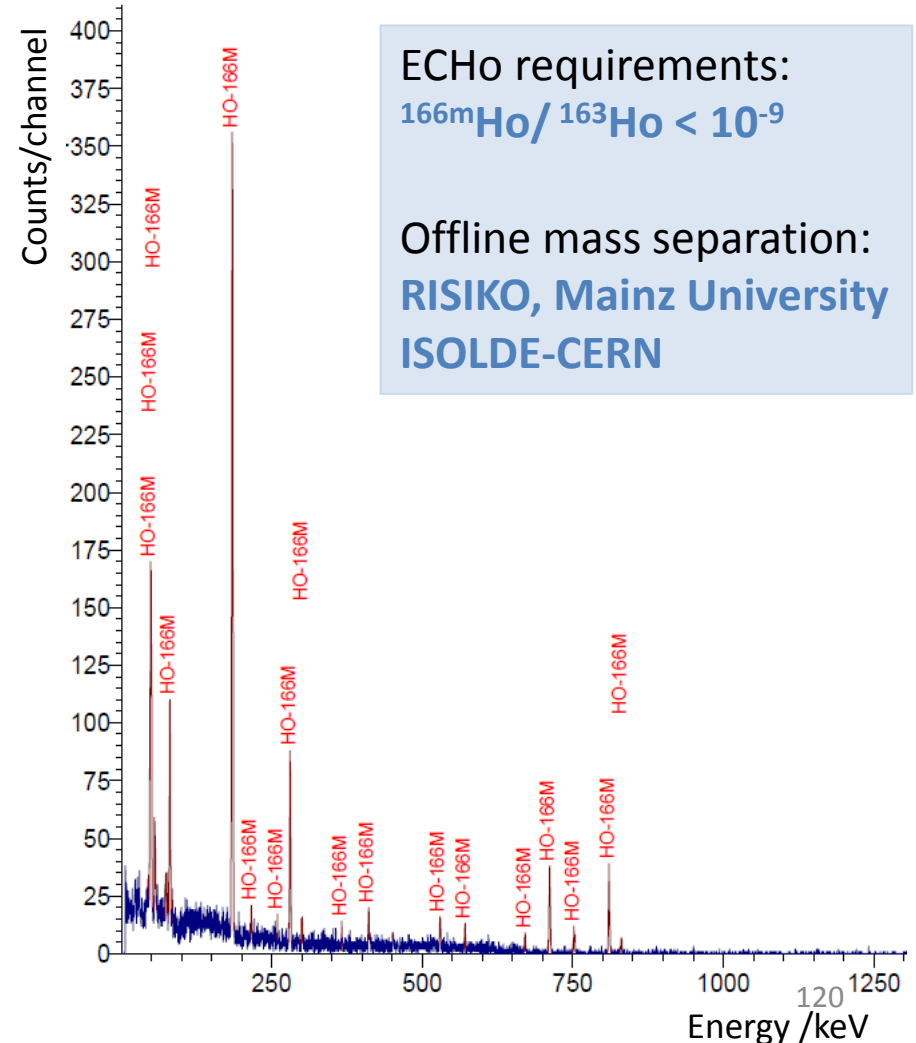
Er161 3.21 h 3/2-	Er162 0+	Er163 75.0 m 5/2-	Er164 0+	Er165 10.36 h 5/2-	Er166 0+
EC	0.14	EC	1.61	EC	33.6
Ho160 25.6 m 5+	Ho161 2.48 h 7/2-	Ho162 15.0 m 1+	Ho163 4570 y 7/2-	Ho164 29 m 1+	Ho165 7/2-
EC *	EC *	EC *	EC *	EC,β *	100

➤ Excellent chemical separation

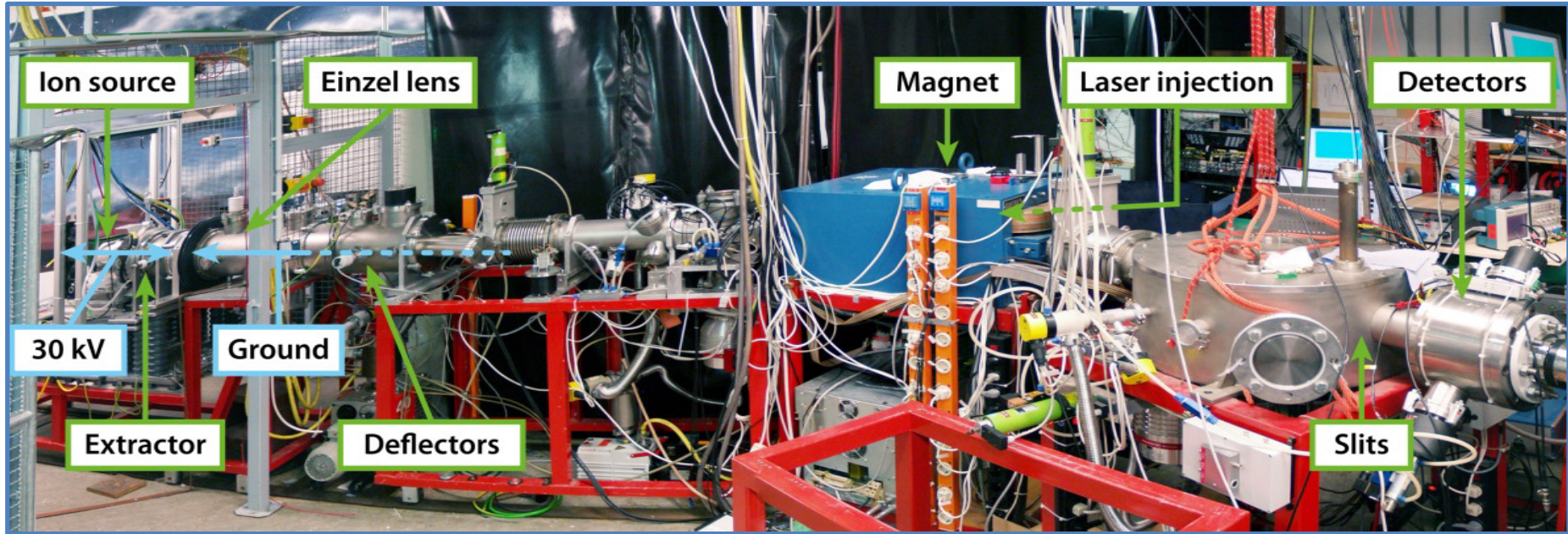
Only $^{166\text{m}}\text{Ho}$

➤ Available ^{163}Ho source:

$\sim 10^{18}$ atoms



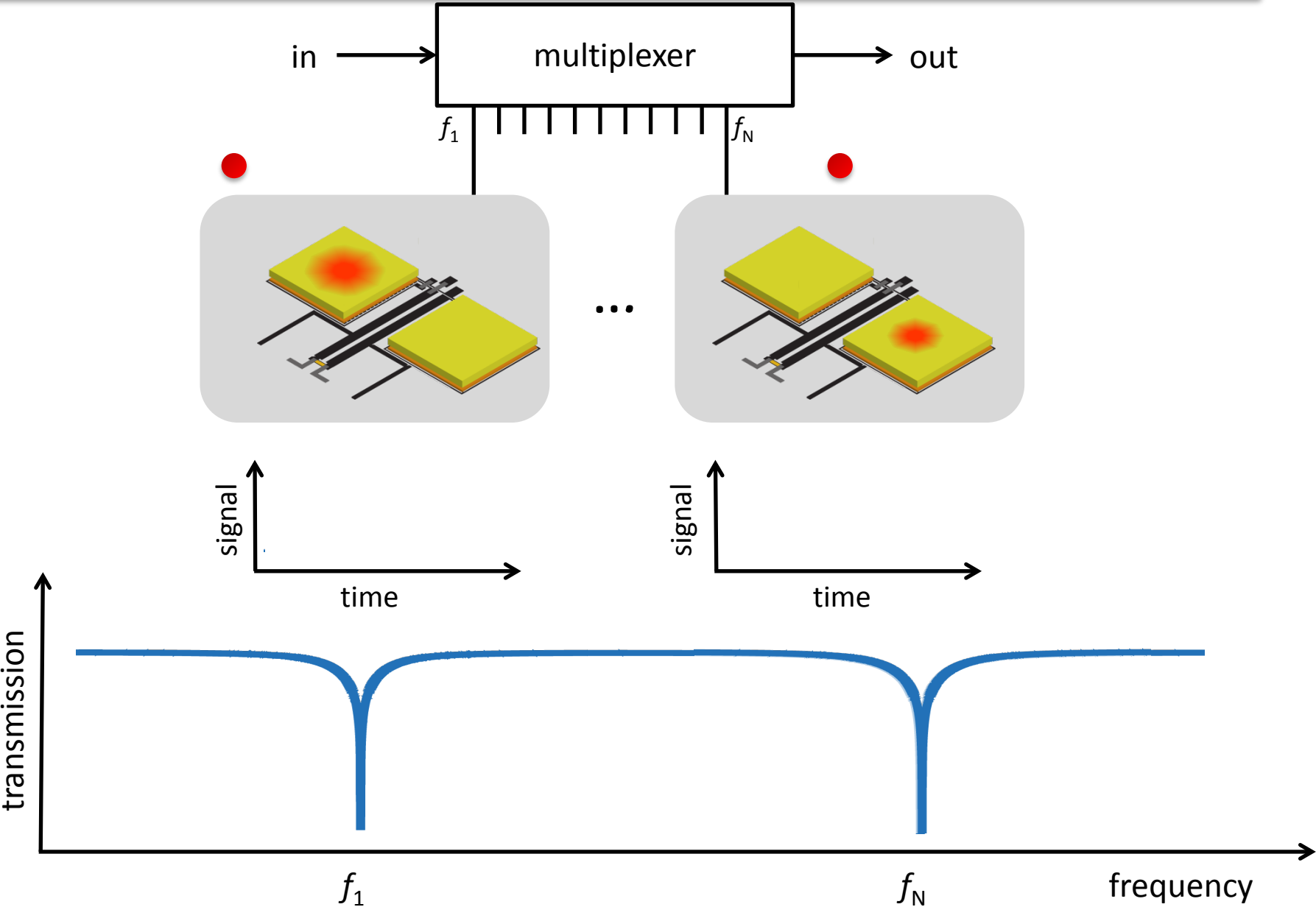
Mass separation and ^{163}Ho ion-implantation



RISIKO @ Physics Institute, Mainz University

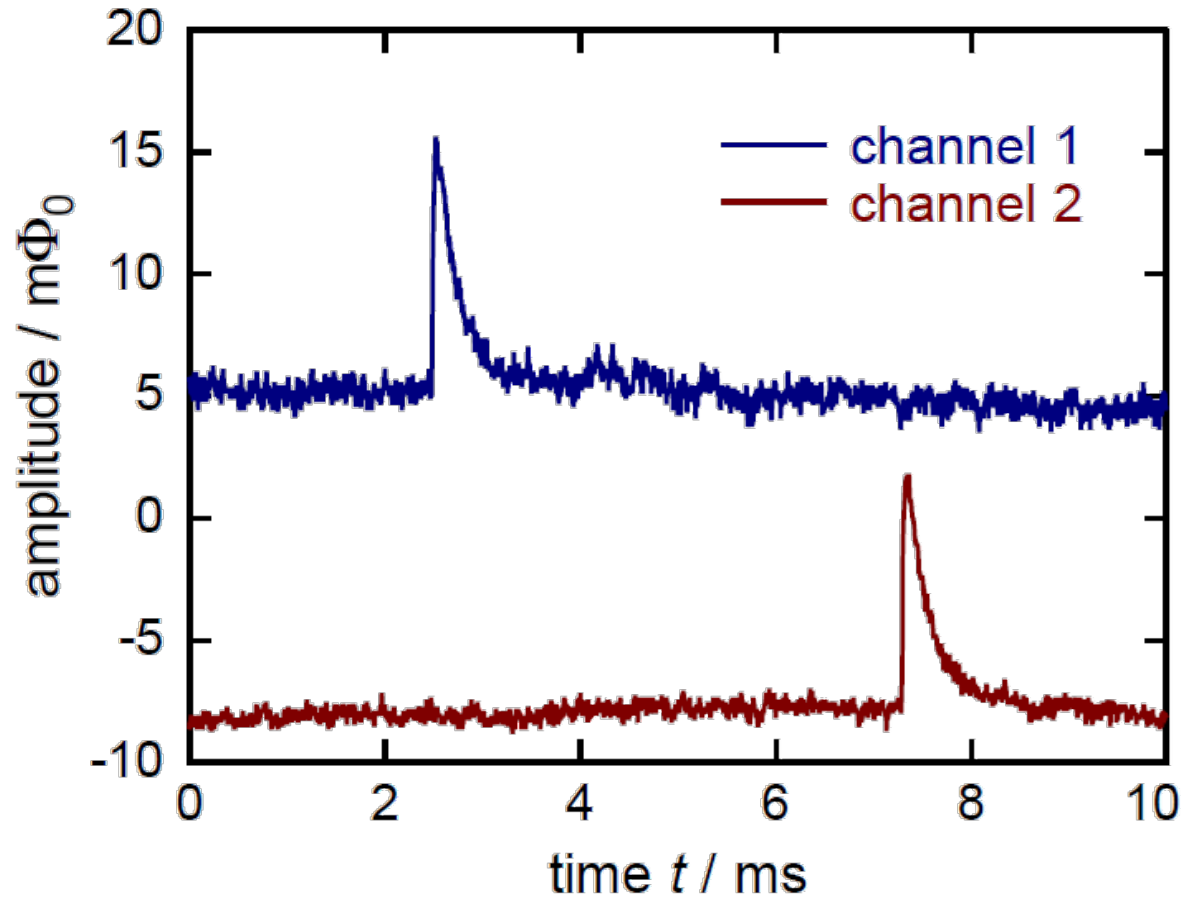
- Resonant laser ion source efficiency **42%**
- Suppression of neighboring masses **> 700**
→ $^{166\text{m}}\text{Ho}/^{163}\text{Ho} < 10^{-5}$
- Optimization of beam focalization

MMCs: Microwave SQUID multiplexing



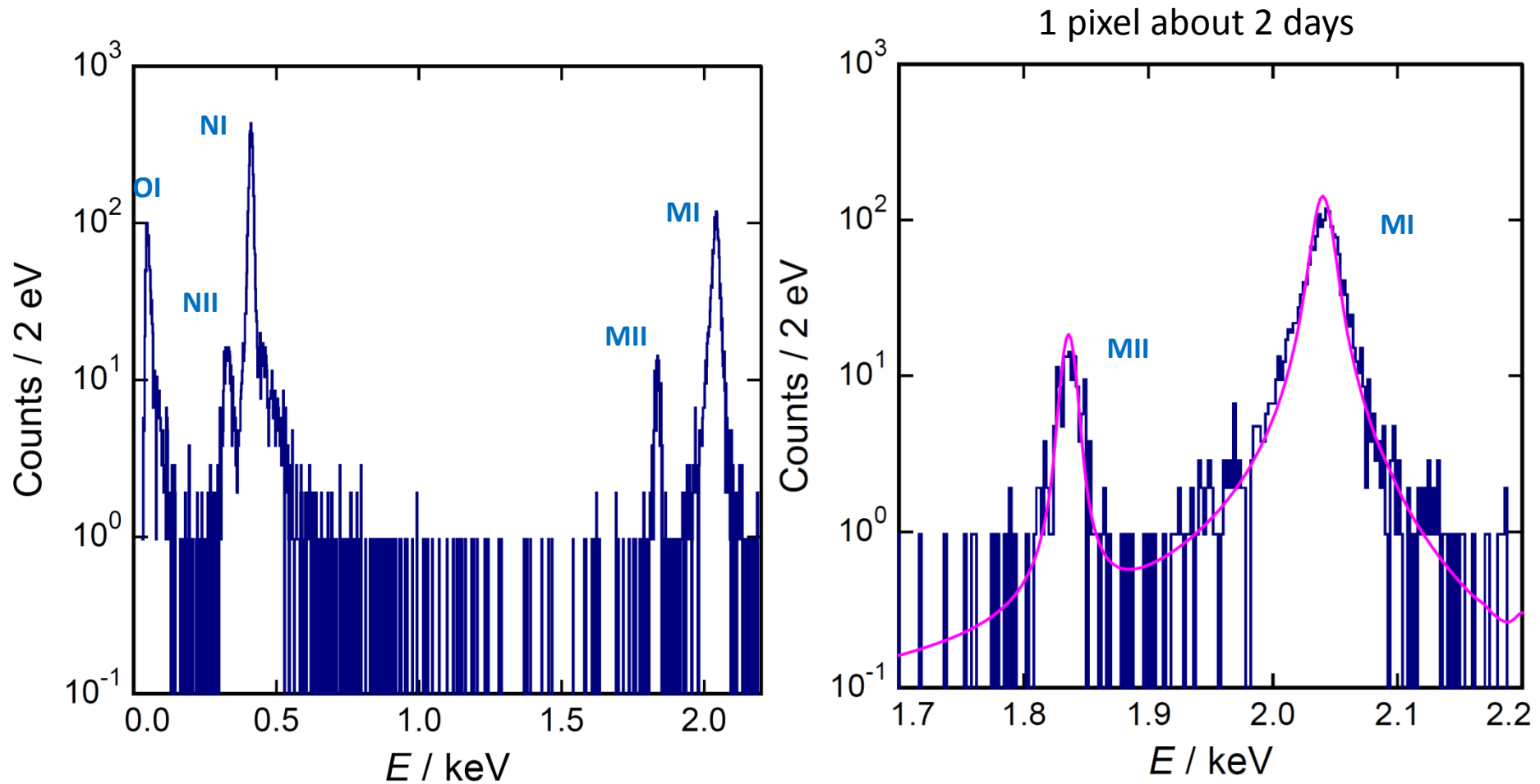
MMCs: Microwave SQUID multiplexing

simultaneous acquisition of signals from two independent detectors using a μ MUX



➡ very first demonstration of multiplexed MMC readout

163Ho off-line implantation: results



- Activity per pixel $A \sim 0.1 \text{ Bq}$
- Energy resolution $\Delta E_{\text{FWHM}} \sim 10 \text{ eV}$
- No strong evidence of radioactive contamination in the source
- Symmetric detector response

Where to improve

High purity ^{163}Ho source:

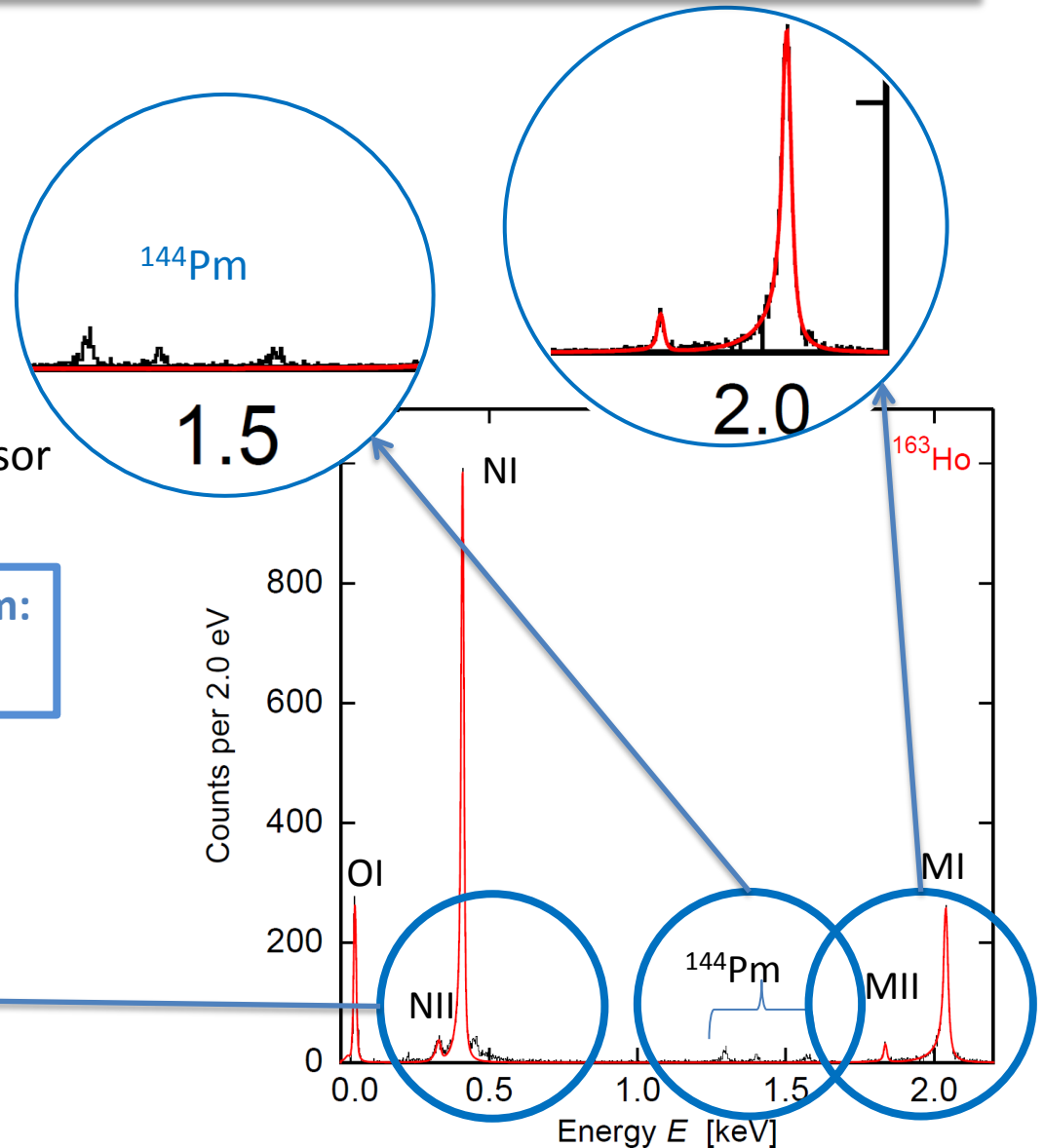
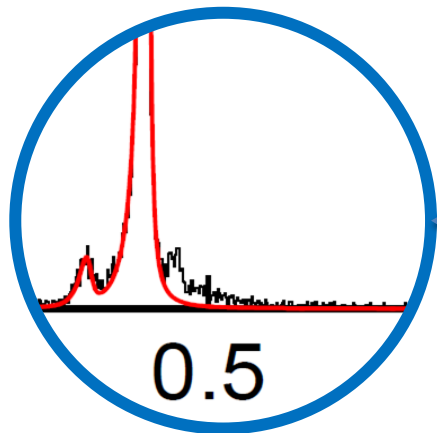
- Background reduction

Detector design and fabrication:

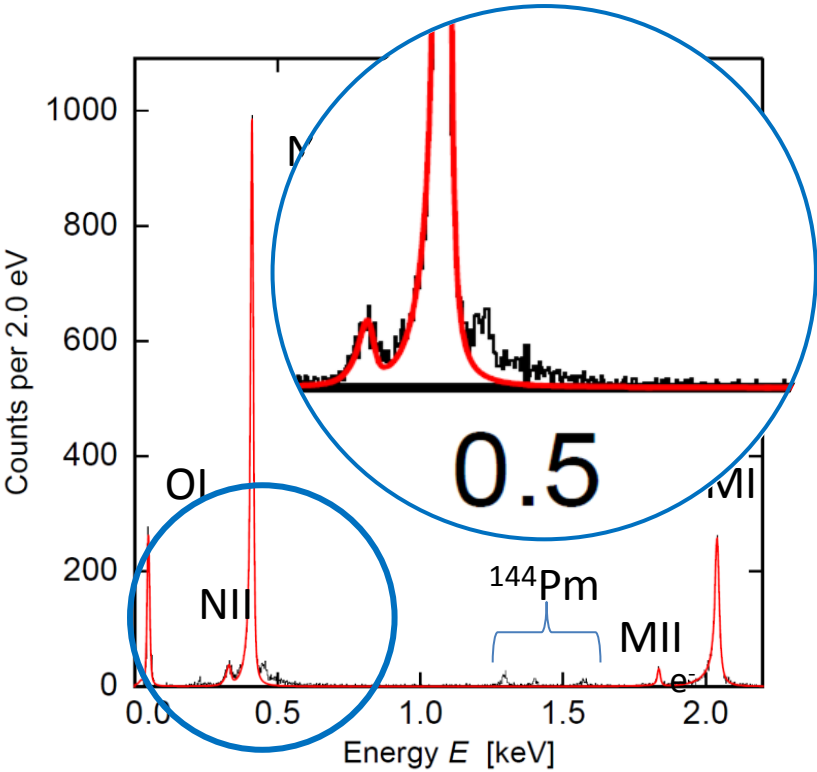
- Increase activity per pixel
- Stems between absorber and sensor

Understanding of the ^{163}Ho spectrum:

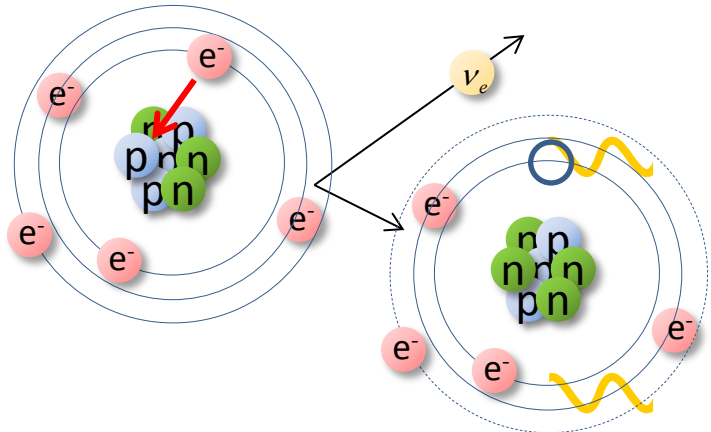
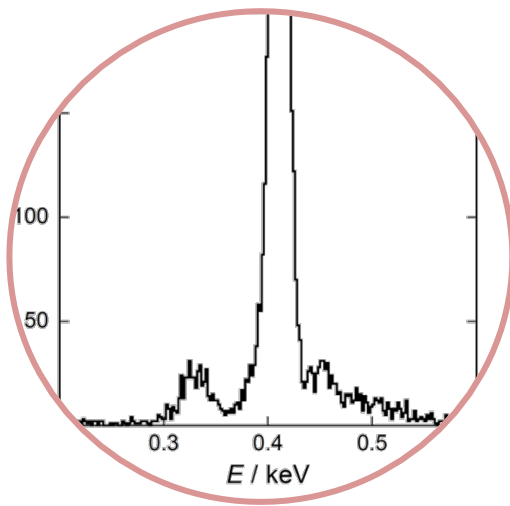
- Investigate undefined structures



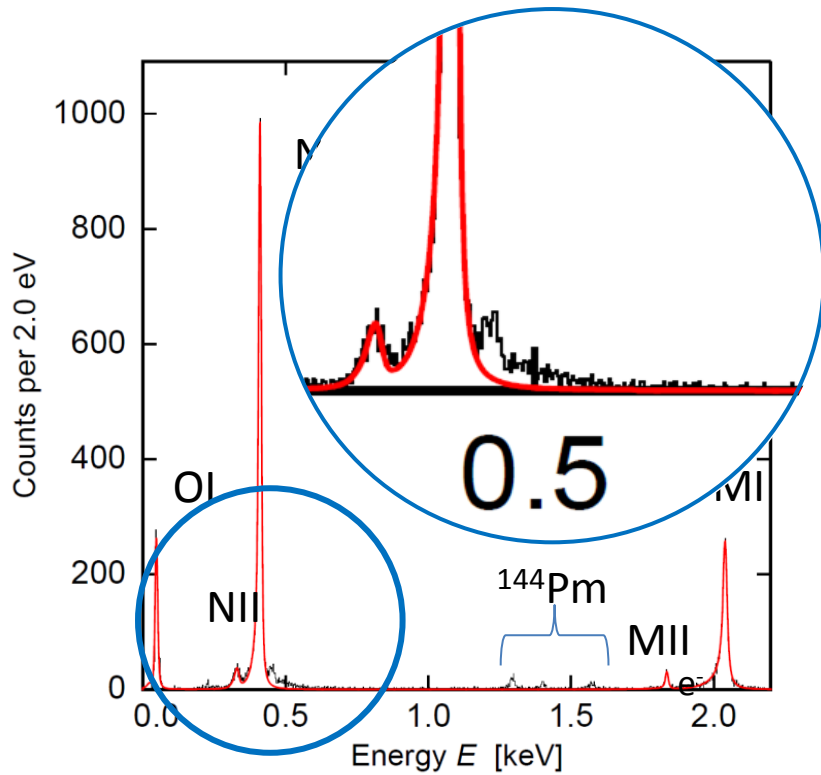
Characterisation of spectral shape



Structures present also in new data

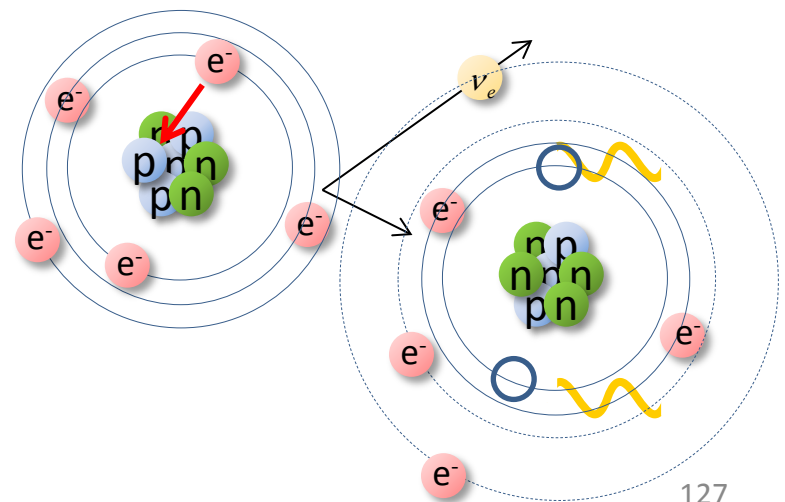


Characterisation of spectral shape

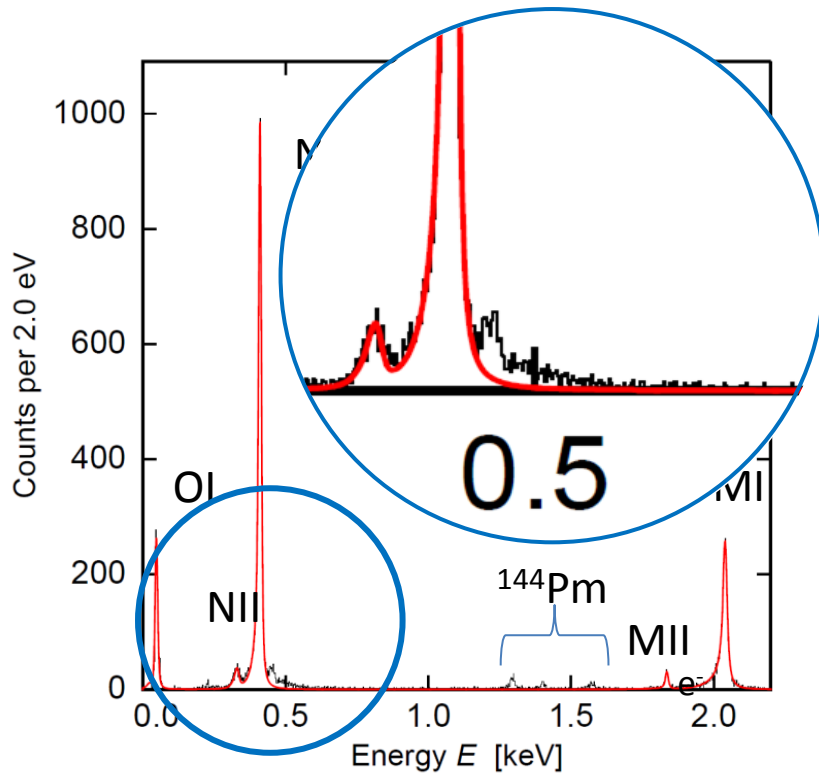


Two-holes excited states: shake-up

- A. Faessler et al.
J. Phys. G **42** (2015) 015108
- R. G. H. Robertson
Phys. Rev. C **91**, 035504 (2015)
- A. Faessler and F. Simkovic
Phys. Rev. C **91**, 045505 (2015)
- A. Faessler et al.
Phys. Rev. C **91**, 064302 (2015)
- A. De Rujula and M. Lusignoli
arXiv:1601.04990v1 [hep-ph] 19 Jan 2016
- A. Faessler et al.
Phys. Rev. C **95**, (2017) 045502

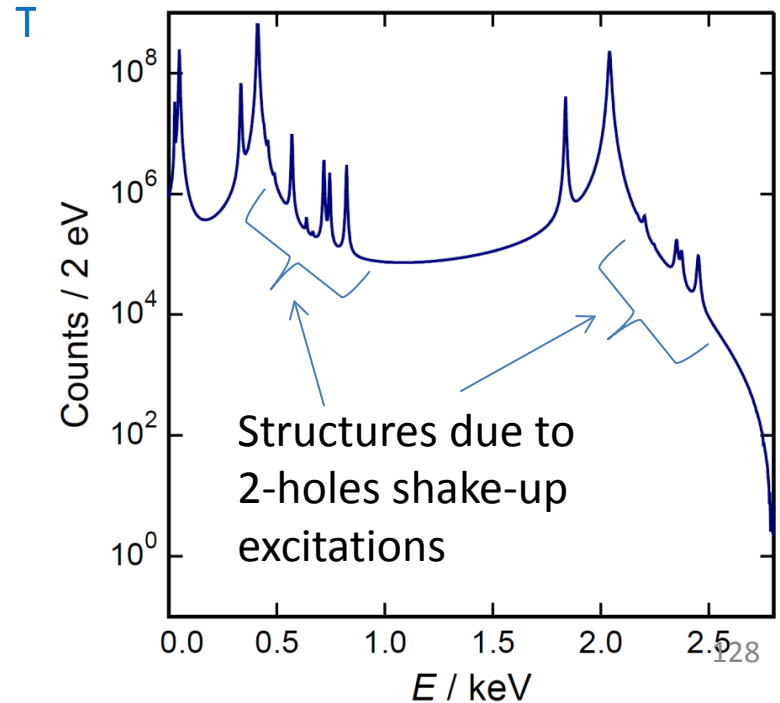


Characterisation of spectral shape

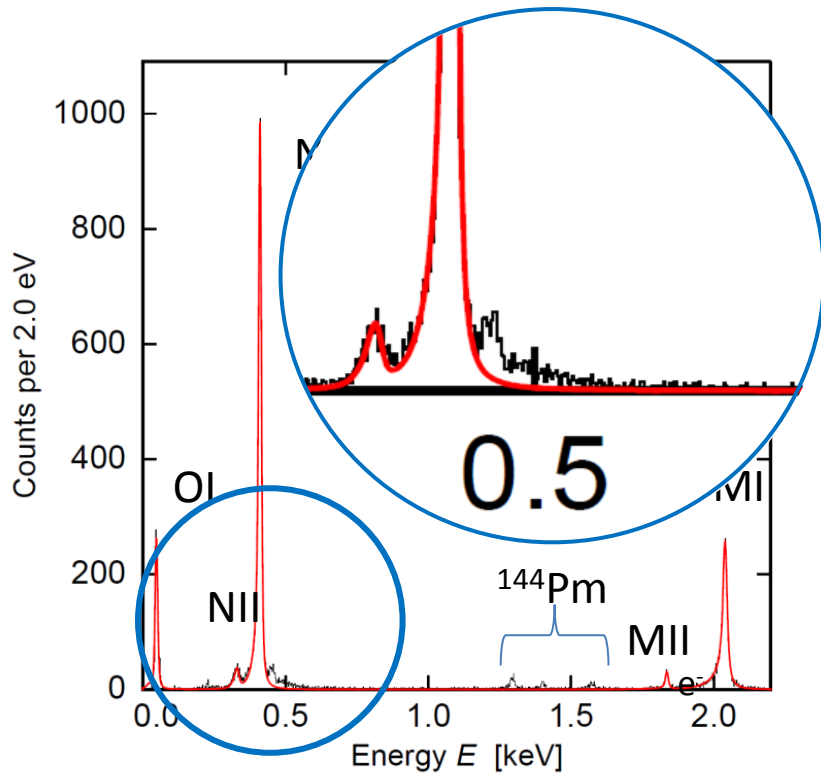


Two-holes excited states: shake-up

- A. Faessler et al.
J. Phys. G **42** (2015) 015108
- R. G. H. Robertson
Phys. Rev. C **91**, 035504 (2015)
- A. Faessler and F. Simkovic
Phys. Rev. C **91**, 045505 (2015)
- A. Faessler et al.
Phys. Rev. C **91**, 064302 (2015)
- A. De Rujula and M. Lusignoli
arXiv:1601.04990v1 [hep-ph] 19 Jan 2016
- A. Faessler et al.
Phys. Rev. C **95**, (2017) 045502

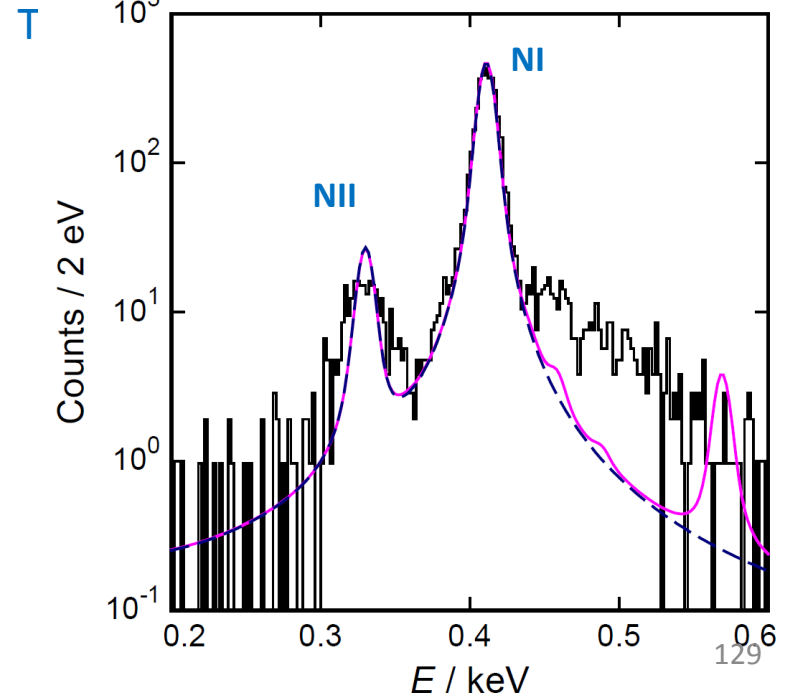


Characterisation of spectral shape

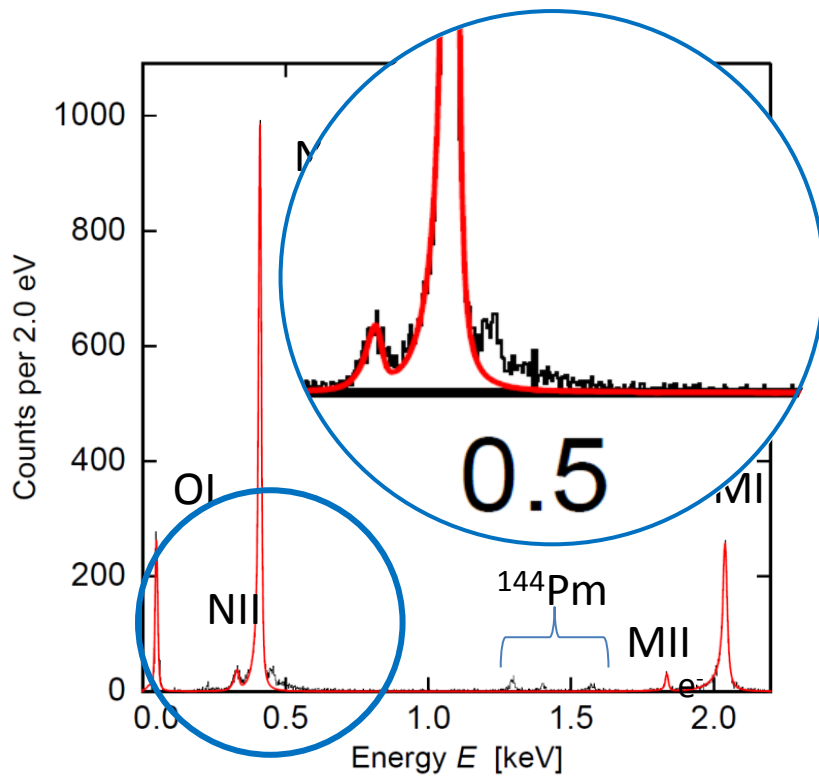


Two-holes excited states: shake-up

- A. Faessler et al.
J. Phys. G **42** (2015) 015108
- R. G. H. Robertson
Phys. Rev. C **91**, 035504 (2015)
- A. Faessler and F. Simkovic
Phys. Rev. C **91**, 045505 (2015)
- A. Faessler et al.
Phys. Rev. C **91**, 064302 (2015)
- A. De Rujula and M. Lusignoli
arXiv:1601.04990v1 [hep-ph] 19 Jan 2016
- A. Faessler et al.
Phys. Rev. C **95**, (2017) 045502



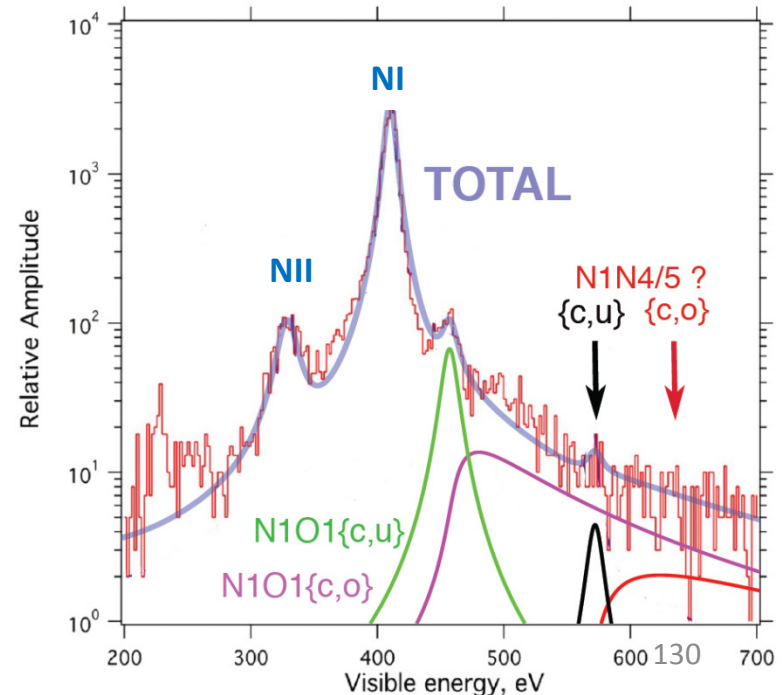
Characterisation of spectral shape



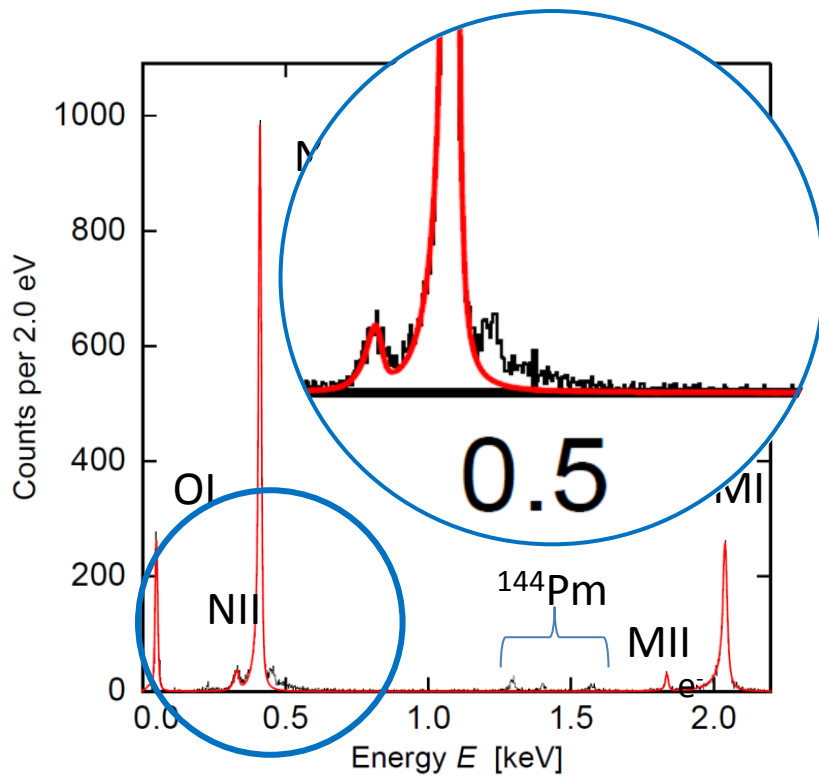
Two-holes excited states: shake-up
shake-off

High statistics and high energy resolution spectra will provide information on the spectral shape

- A. Faessler et al.
J. Phys. G **42** (2015) 015108
- R. G. H. Robertson
Phys. Rev. C **91**, 035504 (2015)
- A. Faessler and F. Simkovic
Phys. Rev. C **91**, 045505 (2015)
- A. Faessler et al.
Phys. Rev. C **91**, 064302 (2015)
- A. De Rujula and M. Lusignoli
[arXiv:1601.04990v1 \[hep-ph\]](https://arxiv.org/abs/1601.04990v1) 19 Jan 2016
- A. Faessler et al.
Phys. Rev. C **95**, (2017) 045502



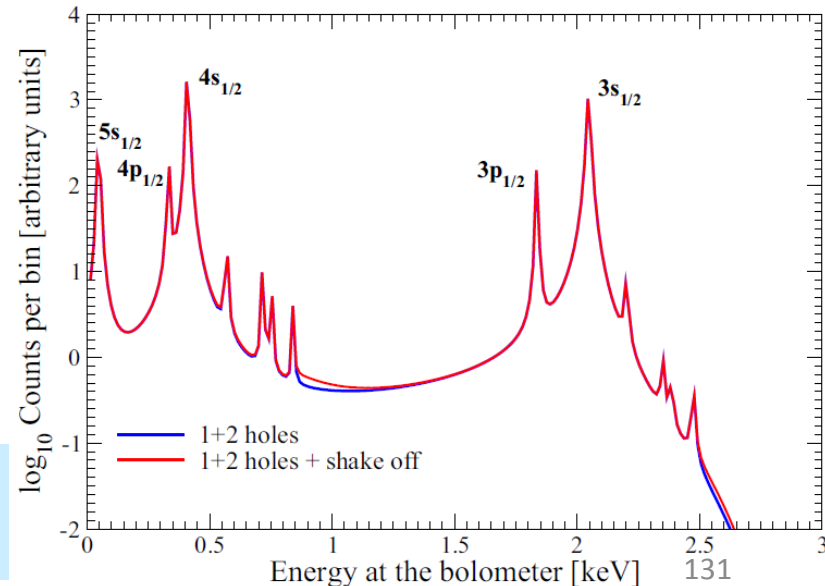
Characterisation of spectral shape



Two-holes excited states: shake-up
shake-off

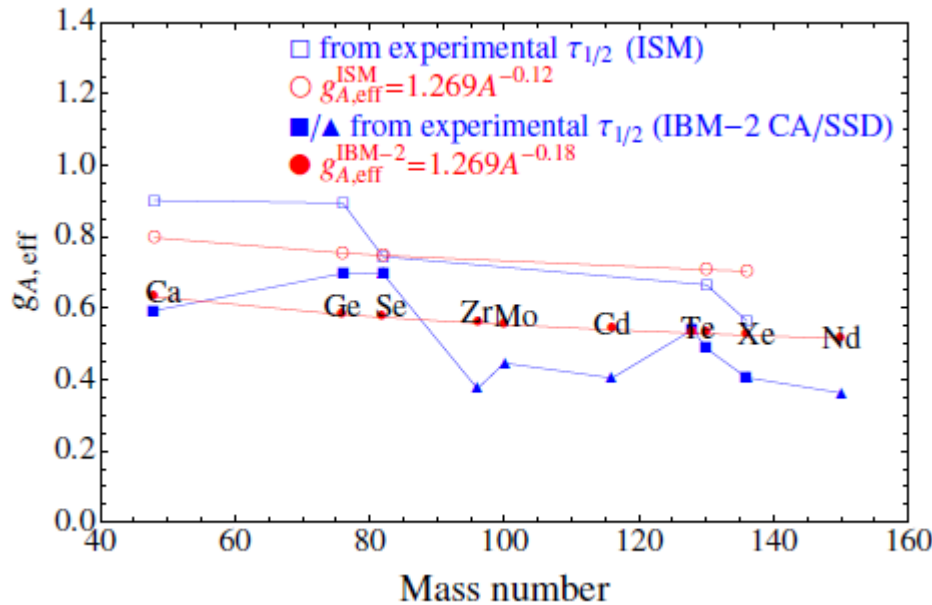
High statistics and high energy resolution spectra will provide information on the spectral shape

- A. Faessler et al.
J. Phys. G **42** (2015) 015108
- R. G. H. Robertson
Phys. Rev. C **91**, 035504 (2015)
- A. Faessler and F. Simkovic
Phys. Rev. C **91**, 045505 (2015)
- A. Faessler et al.
Phys. Rev. C **91**, 064302 (2015)
- A. De Rujula and M. Lusignoli
arXiv:1601.04990v1 [hep-ph] 19 Jan 2016
- A. Faessler et al.
Phys. Rev. C **95** (2017) 045502



Neutrinoless double beta decay - ν mass

The halflife for $0\nu 2\beta$ decay depends on the neutrino mass



Nuclear matrix element

Phase space term

$$\left(\tau_{1/2}^{0\nu}\right)^{-1} = \left|\frac{m_{\beta\beta}}{m_e}\right|^2 \left|M_{\nu}^{0\nu}\right|^2 G^{0\nu}$$

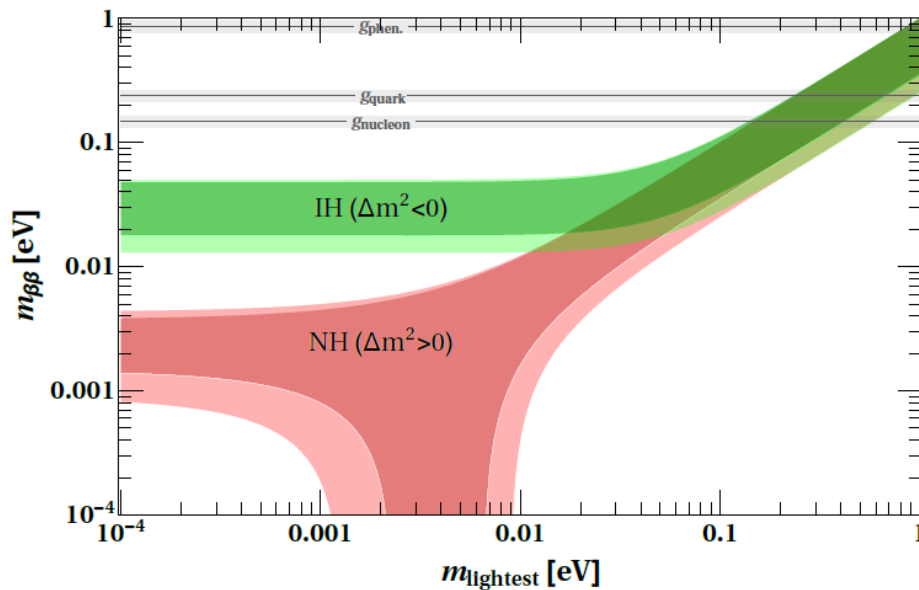
$$m_{\beta\beta}^2 = \left|\sum U_{ei}^2 m(\nu_i)\right|^2$$

Uncertainties to evaluate the effective Majorana mass due to:

- Nuclear matrix element
- Quenching of the axial vector coupling constant g_A

Neutrinoless double beta decay - ν mass

The halflife for $0\nu 2\beta$ decay depends on the neutrino mass



Nuclear matrix element

Phase space term

$$\left(\tau_{1/2}^{0\nu}\right)^{-1} = \left|\frac{m_{\beta\beta}}{m_e}\right|^2 \left|M_{\nu}^{0\nu}\right|^2 G^{0\nu}$$

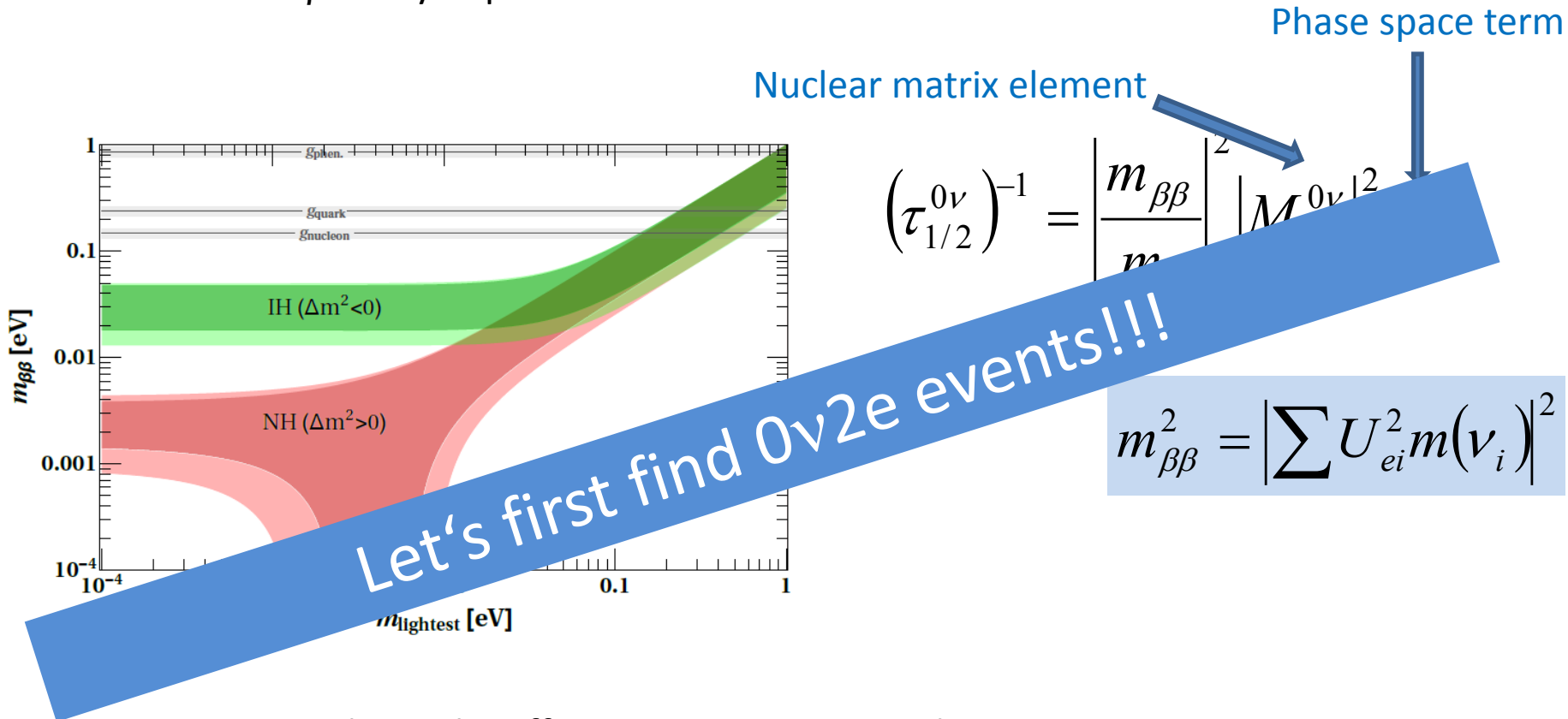
$$m_{\beta\beta}^2 = \left|\sum U_{ei}^2 m(\nu_i)\right|^2$$

Uncertainties to evaluate the effective Majorana mass due to:

- Nuclear matrix element
- Quenching of the axial vector coupling constant g_A

Neutrinoless double beta decay - ν mass

The halflife for $0\nu 2\beta$ decay depends on the neutrino mass



Uncertainties to evaluate the effective Majorana mass due to:

- Nuclear matrix element
- Quenching of the axial vector coupling constant g_A

Fight against background

Direct reduction of background activity

- Select and use ultra-pure materials
- Minimize all passive (non “source”) materials
- Avoid material re-contamination (machining, manipulation, storage)
- Fabricate ultra-clean materials (underground fab if needed)
- underground labs — reduced muon flux & related induced activations

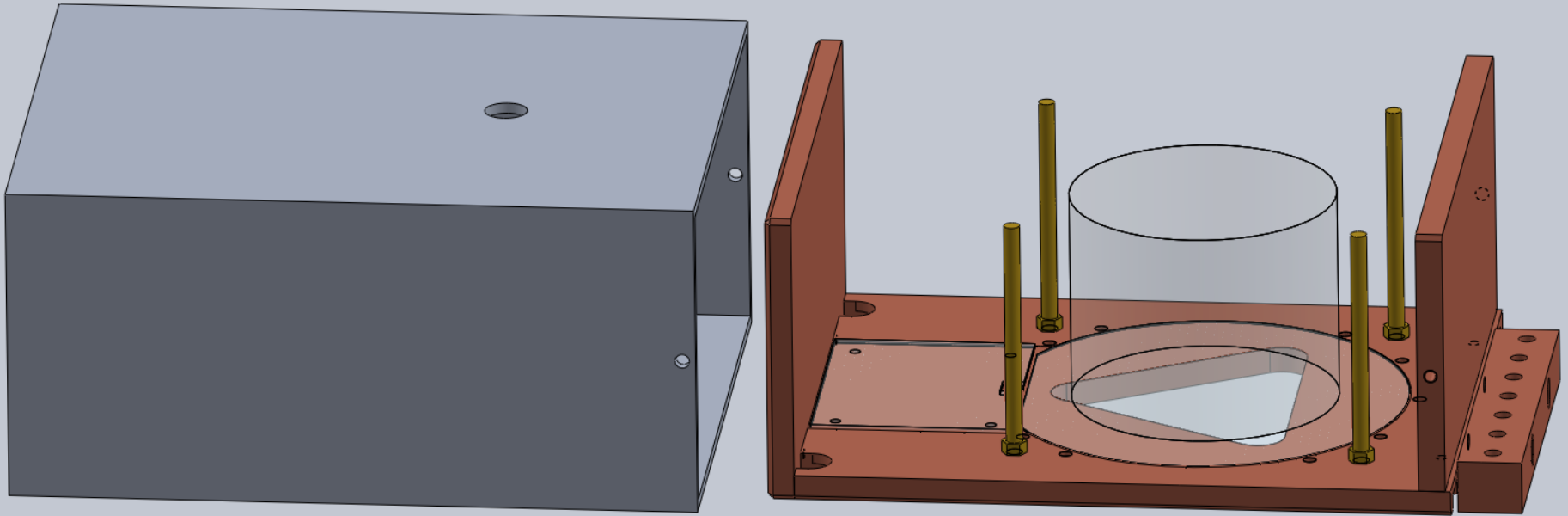
Discrimination techniques

- Energy resolution
- Active veto detector
- Tracking (topology)
- Particle ID, angular, spatial, time correlations
- Fiducial Fits
- Granularity (arrays)
- Pulse shape discrimination (PSD)
- Ion Identification

Methods	
TPCs (liquid, gas)	^{136}Xe
Doped Liquid Scintillators	$^{136}\text{Xe}, ^{130}\text{Te}$
Solid state detectors	$^{76}\text{Ge}, ^{116}\text{Cd}$
Bolometers (+ enhancements)	$^{130}\text{Te}, ^{82}\text{Se}, ^{100}\text{Mo}, ^{116}\text{Cd}$
Foils with tracking chambers	$^{82}\text{Se}, ^{150}\text{Nd}, ^{100}\text{Mo}$

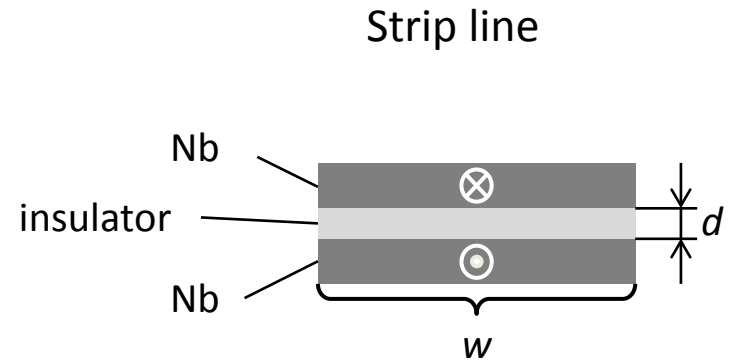
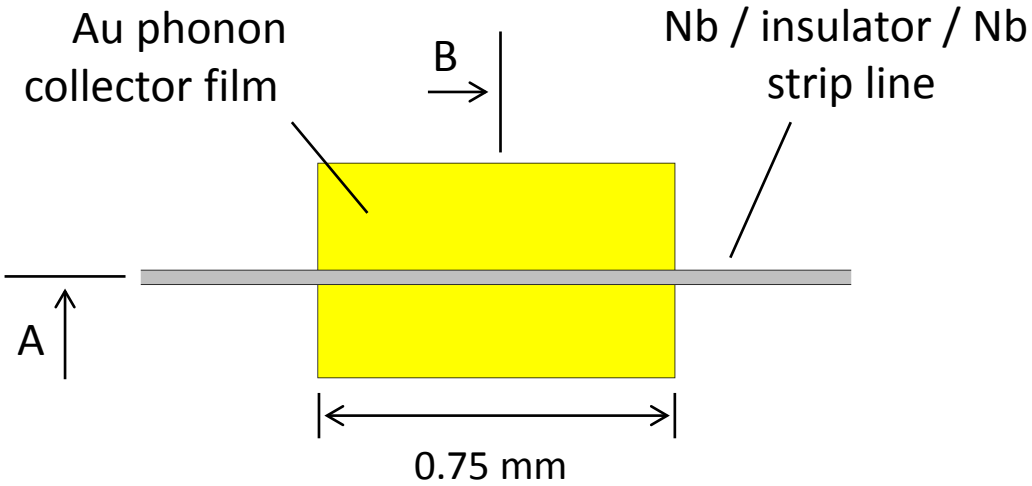
Both approaches are needed

Experimental set-up for P2

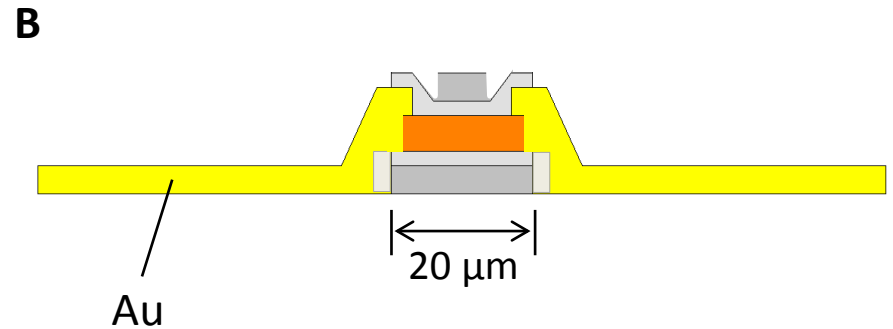
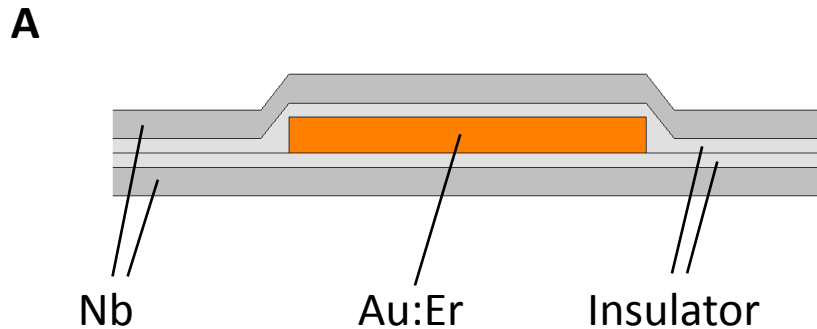


First prototype of photon detector

Top view:



Cross section:



Wärme- & Lichtdetektor

Temperatur = -273°C

